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## Line mixing calculation in the  $v_6 Q$ -branches of  $N_2$ -broadened CH<sub>3</sub>Br **at low temperatures**

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#### **Abstract:**

In an early study [Tran H, Jacquemart D, Mandin JY, Lacome N, JQSRT 2008;109:119-31], line mixing effects of the  $v_6$  band of methyl bromide were observed and modeled at room temperature. In the present work, line mixing effects have been considered at low temperatures using state-to-state collisional rates which were modeled by a fitting law based on the energy gap and a few fitting parameters. To validate the model, several spectra of methyl bromide perturbed by nitrogen have been recorded at various temperatures (205- 299K) and pressures (230-825 hPa). Comparisons between measured spectra and calculations using both direct calculation from relaxation operator and Rosenkranz profile have been performed showing improvement compared to the usual Lorentz profile. Note that the temperature dependence of the spectroscopic parameters has been taken into account using results of previous studies.

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#### **1. Introduction**

Methyl bromide is the major contributor to bromine in the stratosphere and the main organobromide in the lower atmosphere. Atmospheric CH3Br main anthropogenic sources are fumigation processes used in agriculture and automobiles using leaded petrol [1], and some natural ones are the oceans [2, 3] and the biomass burning [4, 5]. This molecule significantly contributes to ozone depletion since  $CH_3Br$  is dissociated by UV radiation, producing Br radicals that catalyze the destruction of ozone [6]. These bromine atoms are 50-60 times more destructive of ozone than the chlorine atoms coming from the chlorofluorocarbons compounds (CFCs) [7]. For this reason, since  $2005$  the use of CH<sub>3</sub>Br has been phased out under the Montreal protocol.

As the CH<sub>3</sub>Br  $v_6$  band is located in an atmospheric window (band center at 954.86 cm<sup>-1</sup>), it is the best candidate for possible atmospheric remote sensing. For this reason, this band is the subject of numerous laboratory spectroscopic studies in order to determine spectroscopic constants [8-12], as well as different line parameters (positions, intensities, broadenings) [13-17]. In particular, a complete set of measured line parameters (positions, intensities, room temperature air- and self-broadenings) for the  $CH_3Br$   $v_6$  band is given in Ref. [15], and the temperature dependence of the  $N_{2}$ - and self-broadening coefficients can be found in Ref. [17]. However, it is well known that *Q*-branch absorption is significantly affected by line mixing effects when the pressure increases. In a previous study performed at room temperature in the  $v_6$  region of CH<sub>3</sub>Br [16], significant deviations were observed when comparing the sum of individual line contributions with Lorentz (or Voigt) profiles and measured spectra even at moderate pressures. A theoretical approach was thus proposed to model line mixing effects within the  $v_6 Q$ -branches of N<sub>2</sub>- and self- perturbed CH<sub>3</sub>Br at room temperature. It was validated by comparing calculations with measured spectra at different conditions of pressure and at room temperature. The results showed great improvement with respect to Lorentz lineshapes.

In order to complete these studies, this work is dedicated to the calculation at low temperatures of line mixing effects in the  $v_6$  Q-branches of CH<sub>3</sub>Br. The theoretical model described in Ref. [16] has been extended for different temperatures. Comparisons between the experiments and the calculation**s** have been performed. Measured spectra and experimental considerations are presented in Section 2. The theoretical model is described in Section 3. Data used and comparisons between calculations and measured spectra are presented in Section 4.

#### **2. Experimental measurements**

The rapid scan Bruker IFS 120 HR interferometer of the Laboratoire de Dynamique, Interactions et Réactivité (LADIR) was used to record N<sub>2</sub>-broadened spectra of CH<sub>3</sub>Br at low temperatures. The unapodized spectral resolution (FWHM) used is about  $5.5 \times 10^{-3}$  cm<sup>-1</sup>, corresponding to a maximum optical path difference of 90 cm. The interferometer was equipped with a Ge/KBr beamsplitter, a MCT photovoltaic detector, a Globar source, and an optical filter covering the 800 - 1100  $cm^{-1}$  spectral region. The experimental conditions of the recorded spectra are summarized in Table 1. Under these conditions, the rotational structure of CH3Br transitions is not completely resolved, as it can be observed on Fig. 1 which presents the whole studied spectral domain. For all spectra, the whole optical path, outside the cell, was under vacuum. A White-type cell of 1 m base length with a total absorption path of 415 cm was used with KCl windows. The commercial gas sample, purchased from Fluka, with a stated purity of 99.50% in natural abundances, was used without further purification. In order to obtain low temperatures, a continuous circulation through the secondary envelop of the cell of methylcyclohexane cooled by liquid nitrogen was used. The pressure of the gas was measured with a capacitance MKS Baratron manometer with an accuracy estimated to be  $\pm 1\%$ . Averaging the 100 scans, the signal to noise ratio is nearly equal to 100.

As a result, 9 high-resolution spectra of  $CH_3Br/N_2$  mixtures were recorded at different total pressures (~ from 230 to 825 hPa) and temperatures (from 205 to 299 K) (see Table 1).

#### **3. Theoretical model**

To model line mixing (LM) effects, a similar theory as what was done in Ref. [16] is used. Impact approximation and binary collisions are assumed. The Doppler width is neglected compared with the collisional broadening in the studied pressure range. Within this frame and accounting for line mixing (LM) effects, the absorption coefficient,  $\alpha^{LM}$ , in cm<sup>-1</sup>(CGS system), for CH<sub>3</sub>Br diluted in N<sub>2</sub> at wavenumber  $\sigma$  and temperature T is given by:

$$
\alpha^{LM}(\sigma, P_{CH_3Br}, P_{N_2}, T) = \frac{8\pi^2 \sigma}{3hc} P_{CH_3Br}^{iso} [1 - \exp(-hc\sigma / k_B T)]
$$
  
\n
$$
\times \sum_{k} \sum_{\ell} \rho_k(T) d_k d_\ell \times \text{Im} \left\{ \langle \ell | [\Sigma - L_0 - i(P_{CH_3Br} W_{CH_3Br/CH_3Br}(T) + P_{N_2} W_{CH_3Br/N_2}(T))]^{-1} | k \rangle \right\}
$$
\n(1)

where  $k_B$  is the Boltzman constant, and the summation is over all  $k$  and  $\ell$  absorption lines.  $P_{CH3Br}$  and  $P_{N2}$  are respectively CH<sub>3</sub>Br and N<sub>2</sub> partial pressures. Note that,  $P_{CH3Br}$  is the total pressure of two isotopologues,  $CH_3^{79}Br$  and  $CH_3^{81}Br$ .  $P_{CH_3Br}^{iso}$  corresponds to the partial pressure of the considered absorbing isotopologue (i.e iso = 79 or 81, respectively 50.54 or 49.46 % of the total pressure of  $CH_3Br$ ). As the two isotopologues have similar abundances and absorb in the same spectral region, they have to be calculated simultaneously, so that the whole absorption coefficient is actually the sum of the absorption coefficients due to each isotopologue.  $\Sigma$ ,  $L_0$  and W are operators in the Liouville (line) space and  $\langle \langle \ell | ... | k \rangle$ designate a matrix element.  $\rho_{\ell}$  and  $d_{\ell}$  are the relative population of the initial level and the dipole matrix element of line  $\ell$ . The latter is related to the integrated intensity  $S_{\ell}$  by the following expression:

$$
S_{\ell}(T) = \frac{8\pi^3}{3hc} \rho_{\ell}(T) \times \sigma_{\ell} \times [1 - \exp(-hc\sigma_{\ell}/k_{B}T)] \times d_{\ell}^{2}.
$$
 (2)

 $\Sigma$  and  $L_0$  are diagonal and respectively associated with the current wavenumber and the position  $\sigma_{\ell}$  of individual transitions, i.e.:

$$
\langle \langle \ell | \mathbf{\Sigma} | k \rangle \rangle = \sigma \times \delta_{k,\ell} , \qquad \langle \langle \ell | \mathbf{L}_0 | k \rangle \rangle = \sigma_k \times \delta_{k,\ell} . \tag{3}
$$

where  $\delta_{k,l}$  is the Kronecker symbol and W is the relaxation operator which contains information on the effects of collisions on the spectral line shape, it depends on the vibrational band, on the temperature, and on the collision partner. Its off-diagonal elements account for the interferences between absorption lines (line mixing) whereas the real and imaginary parts of the diagonal ones are respectively the collisional broadening and shifting coefficients,  $\gamma$ and  $\delta_{\ell}$ 

$$
\langle \langle \ell | \mathbf{W}(T) | \ell \rangle \rangle = \gamma_{\ell}(T) - i \delta_{\ell}(T). \tag{4}
$$

Since the pressure of CH<sub>3</sub>Br is very small compared to the  $N_2$  pressure (0.24-0.4% of the total pressure), the term taking into account the line mixing due to the self-perturbation of CH<sub>3</sub>Br in Eq. (1) is neglected. Therefore, the W<sub>CH3Br/CH3Br</sub> matrix becomes diagonal and only line mixing effects due to collisions of  $CH_3Br$  with  $N_2$  are considered. The real off-diagonal elements of  $\mathbf{W}_{CH_3Br/N_2}$  (the imaginary part is neglected), are modeled using the state-to-state inelastic collisional rates of the lower states as

$$
\langle \langle \ell | \mathbf{W}_{CH_3Br/N_2}(T) | k \rangle \rangle = -A_{\ell,k} \times K(i_{\ell} \leftarrow i_k, T) \text{ , with } (k \neq \ell \text{ ),}
$$
 (5)

where  $K(i_k \leftarrow i_k, T)$  are the collisional transfer rates from the initial level,  $i_k$ , of the *k* line to the initial level,  $i_{\ell}$ , of the  $\ell$  line (*k* and  $\ell$  belonging to the same isotopologue). The empirical  $A_{\ell,k}$ parameters which relate state-space to line-space depend in general on the considered  $k$  and  $\ell$ lines. In order to simplify the problem,  $A_{\ell,k}$  is considered to depend only on the type of the considered band and on the buffer gas. Furthermore, we assumed that *A* is independent of the temperature, so that the value of the preceding work at room temperature [16] has been retained, i.e.  $A = 0.7$  for all *Q*-branches of the  $v_6$  band.

In order to model the state-to-state collisional rates  $K(i_k \leftarrow i_k, T)$ , a Modified Exponential Gap (MEG) fitting law [18-20] has been used

$$
K(i_{\ell} \leftarrow i_{k}, T) = a_{1} \left( \frac{1 + a_{4} (E_{k} / a_{2} k_{B} T)}{1 + a_{4} (E_{k} / k_{B} T)} \right)^{2} \exp \left( - \frac{a_{3} |E_{k} - E_{\ell}|}{k_{B} T} \right) \left( \frac{T}{T_{0}} \right)^{n} .
$$
 (6)

where  $a_1, a_2, a_3, a_4$  and *n* are fitting parameters and  $T_0$  is a reference temperature. Equation (6) is applied for the downward rates, i.e.,  $E_{\ell} \langle E_{k}$ , where *E* is the rotational energy of the initial level of line  $(\ell \text{ or } k)$ . The upward rates are calculated by using the detailed balance

$$
K(i_k \leftarrow i_\ell, T) = \frac{\rho_k(T)}{\rho_\ell(T)} K(i_\ell \leftarrow i_k, T), \tag{7}
$$

Since the value of  $T_0$  is 296 K,  $n$  was ignored in the preceding work [16] performed at room temperature. The values of  $a_1$ ,  $a_2$ ,  $a_3$  and  $a_4$  parameters are kept unchanged from those obtained at room temperature [16] by least-squares fitting the sum rule to the  $\gamma_k$  values for the  ${}^RQ_0$  branch. Note that the values of *a*<sub>1</sub>, *a*<sub>2</sub>, *a*<sub>3</sub> and *a*<sub>4</sub> are the same for all  $v_6 Q$ -branches and for the two isotopologues. The value of  $n$  is then obtained by manually fitting the sum rule using the N<sub>2</sub>-broadenings coefficients for the  ${}^RQ_0$  branch obtained at low temperatures [17], i.e.

$$
\gamma_k^{CH_3Br/N_2}(T) = \frac{1}{2} \left( \sum_{\ell \neq k} K(i_\ell \leftarrow i_k, T) + \sum_{\ell \neq k} K(f_\ell \leftarrow f_k, T) \right), \tag{8}
$$

with  $K(i_{\ell} \leftarrow i_{k}, T)$  given by Eqs. (6,7).

In order to simplify computations for possible atmospheric applications, the absorption coefficient was also calculated using the Rosenkranz profile [21]. Expanding Eq. (1) to first order, we obtain:

$$
\alpha^{LM}(\sigma, P_{\text{CH}_3\text{Br}}, P_{\text{N}_2}, T) = \frac{1}{\pi} P_{\text{CH}_3\text{Br}} \sum_{\ell} S_{\ell}(T) \frac{\sigma}{\sigma_{\ell}} \frac{\gamma_{\ell} + (\sigma - \sigma_{\ell})Y_{\ell}}{(\sigma - \sigma_{\ell})^2 + {\gamma_{\ell}}^2},
$$
(9)

where  $Y_{\ell}$  is a line mixing parameter representing the coupling between line  $\ell$  and all other lines of the same branch. The  $Y_{\ell}$ 's are given by:

$$
Y_{\ell} = 2P_{N_2} \sum_{k \neq \ell} \frac{d_k}{d_\ell} \frac{\langle \langle k | \mathbf{W}_{\mathrm{CH}_3 \mathrm{Br}/\mathrm{N}_2}(T) | \ell \rangle \rangle}{\sigma_{\ell} - \sigma_k} \ . \tag{10}
$$

Calculations using the Lorentz lineshapes, the direct calculation (Eq. (1)) and the Rosenkranz profile (Eqs. (9-10)) have been performed. Comparisons between these calculations and the experimental spectra are presented in the next section.

#### **4. Data used, results and discussion**

In this work, the low temperature line mixing effects in the  $Q$ -branches of  $v_6$  band have been investigated using the line parameters obtained in Ref. [15] at 296 K together with the average temperature dependence of the self- and  $N_2$ -broadening coefficients of Ref. [17]. Note that this set of parameters corresponds to the actual data of the GEISA and HITRAN 2008 editions. For the temperature dependence of line intensities, the following expression has been used:

$$
S_{\ell}(T) = S_{\ell}(T_0) \exp\left[-\frac{hcE_{\ell}}{k_B} \left(\frac{1}{T} - \frac{1}{T_0}\right)\right] \frac{\left[1 - \exp\left(\frac{hcE_{\ell}}{T}\right)\right]}{1 - \exp\left(\frac{hcE_{\ell}}{T_0}\right)} \frac{Q_{tot}(T_0)}{Q_{tot}(T)},\tag{11}
$$

where  $Q_{\text{tot}}(T)$  and  $Q_{\text{tot}}(T_0)$  are the total partition functions of CH<sub>3</sub>Br at temperatures *T* and  $T_0$ respectively. Partition functions for a given temperature *T* have been obtained as the product of the vibrational  $(Q_{vib})$  and rotational  $(Q_{rot})$  partition functions, and using spectroscopic constants of Ref. [11].

For the broadening parameters at temperature *T*, the following equations have been used:

$$
\gamma_{\ell}^{CH_3Br/CH_3Br}(T) = \left(\frac{T_0}{T}\right)^{n_{\ell}^{CH_3Br/CH_3Br}} \gamma_{\ell}^{CH_3Br/CH_3Br}(T_0)
$$
\n
$$
\gamma_{\ell}^{CH_3Br/N_2}(T) = \left(\frac{T_0}{T}\right)^{n_{\ell}^{CH_3Br/N_2}} \gamma_{\ell}^{CH_3Br/N_2}(T_0)
$$
\n
$$
\gamma_{\ell}(T) = P_{CH_3Br}\gamma_{\ell}^{CH_3Br/CH_3Br}(T) + P_{N_2}\gamma_{\ell}^{CH_3Br/N_2}(T)
$$
\n(12)

where  $\gamma_{\ell}^{CH_3Br/CH_3Br}(T_0)$  and  $\gamma_{\ell}^{CH_3Br/N_2}(T_0)$  are the self- and N<sub>2</sub>-broadening coefficients for CH<sub>3</sub>Br at 296K (taken from Ref. [15]), and  $n_e^{CH_3Br/CH_3Br}$  $n_e^{CH_3Br/CH_3Br}$  and  $n_e^{CH_3Br/N_2}$  their temperature dependence modelled by a *J*-dependent polynomial function (Eqs. (4,5) of Ref. [17]).

Using Eqs. [6-8] and the values of  $\gamma_{\ell}^{CH_3Br/N_2}(T)$ , the parameter *n* (see Eq. (6)) has been found equal to 1.43 $\pm$ 0.05. As  $a_1$ ,  $a_2$ ,  $a_3$  and  $a_4$  parameters, *n* is also assumed to be independent on the considered branch. The calculated line broadening coefficients obtained in this way are presented in Fig. 2 for the  ${}^RQ_0$  and  ${}^PQ_1$  branches at different temperatures. As it can be observed, a good agreement is found, validating the fitted parameters of the MEG model. However, this agreement is not perfect, so that the renormalization procedure [22,23] has been performed.

For the calculation of the spectra, the apparatus function has been neglected because of the high pressures used in this work. Both isotopologues have been calculated using line mixing for transitions of the strong *Q*-branches (*ΔK*=0, *K*=0-6) and Lorentz profiles for the other transitions.

Comparisons between the measurements and calculations with and without the linemixing effects are shown on Figs. 1, 3 and 4. A detailed view of the  ${}^RQ_0$  branch in various conditions of pressures and temperatures is presented in Fig. 3, whereas the whole spectrum is shown on Fig. 1. As it can be observed in Fig. 3, Lorentzian lineshapes strongly underestimate absorption in the low *J* region of the *Q*-branch whereas the direct calculation (Eq. (1)) and the Rosenkranz profile (Eq. (9)) lead to better agreements. Similar results are obtained for all other *Q*-branches, as can be observed in Fig. 1 for the whole set of *Q*branches, or in Fig. 4 where detailed view of  ${}^RQ_2$ ,  ${}^RQ_0$  and  ${}^PQ_2$  is shown. Note that under the pressure range considered in this work, the use of the Rosenkranz profile leads to almost same results as those obtained by the direct calculation.

#### **5. Conclusion**

Several spectra of  $CH_3Br$  perturbed by  $N_2$  have been recorded at different temperatures (from 205 to 299 K) and pressures (from 230 to 825 hPa) using a Bruker 120 FTIR spectrometer. For the calculation of the spectra, the relaxation matrix has been built including temperature dependences. Both direct calculation and Rosenkranz profile have been used to model the experimental spectra. In both cases, the residuals are considerably reduced compared to the ones obtained by using Lorentz lineshapes. It is then clear that line mixing must be accounted for to correctly model the  $v_6 Q$ -branches of CH<sub>3</sub>Br.

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#### FIGURE CAPTIONS

**Figure 1**. Overview of the  $v_6$  band for the experimental spectrum of CH<sub>3</sub>Br/N<sub>2</sub> at  $T = 234$  K  $(P_{\text{CH3Br}} = 1.52 \text{ hPa}$  and  $P_{\text{tot}} = 476.1 \text{ hPa}$ . The residuals (obs-calc) have been plotted in the lower panels using Lorentz lineshapes (a), direct calculation of the relaxation operator (b), and Rosenkranz profile (c).

**Figure 2.** Comparison between N<sub>2</sub>-broadening coefficients of CH<sub>3</sub>Br calculated by the MEG model (open circles) at different temperatures and those obtained using Refs. [15,17] (solid squares).

**Figure 3**. Comparison between measured transmissions  $\bullet$  of CH<sub>3</sub>Br diluted in N<sub>2</sub> in the  $v_6$  ${}^RQ_0$  region and calculations without (---) (Lorentz lineshapes) and with (--) line-mixing effects using direct calculation (Eq. (1)). The corresponding residuals are plotted in the middle panels. For comparison, residuals with calculations using Lorentz  $(-)$  and Rosenkranz  $(-)$ profiles (Eqs. (9,10)) are shown in the lower panels.

**Figure 4**. Comparisons for different branches between the calculations without (**---**) (Lorentz lineshapes) and with  $(-)$  taking into account line mixing effect (direct calculation). The experimental spectrum ( $\blacksquare$ ) corresponds to *T*=265 K, *P*<sub>CH3Br</sub>=1.74 hPa, and *P*<sub>tot</sub>=545.50 hPa.

**Figure 5**. Measured transmissions ( $\bullet$ ) of CH<sub>3</sub>Br diluted in N<sub>2</sub> in the  $v_6{}^RQ_0$  region at different temperatures and pressures and calculations with  $(-)$  and without  $(-)$  line-mixing effects using the Rosenkrantz approximation. In lower panels are represented the corresponding (obscalc).

### TABLE CAPTIONS

**Table 1.** Experimental conditions and characteristics of the recorded spectra



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