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1 Water in type I chondrules of Paris CM chondrite

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10

11 12 Abstract

13
14 Hydrogen isotopic ratio and water concentration have been measured with the NanoSIMS in olivine,
15 pyroxene and mesostasis phases in individual chondrules from the carbonaceous chondrites Paris (CM2),
16 Renazzo (CR2) and ordinary chondrite Bishunpur (LL3). On average, chondrule pyroxenes in Renazzo,
17 Bishunpur and Paris contain 893 ± 637 ppm (1SD), 879 ± 536 ppm and 791 ± 227 ppm H₂O. Chondrule
18 olivines from Renazzo and Bishunpur vary from 156 ± 44 ppm to 222 ± 123 ppm. Olivines in the Paris
19 chondrules have high water concentration (603 ± 145 to 1051 ± 253 ppm H₂O) with a minimum mean value
20 of 645 ± 99 ppm. δD ranges from $-212\pm 125\text{‰}$ to $15\pm 156\text{‰}$ and from $-166\pm 133\text{‰}$ to $137\pm 176\text{‰}$ in Renazzo
21 and Bishunpur chondrule olivines, pyroxenes and mesostases, respectively. In Paris chondrules, δD ranges
22 from $-398\pm 23\text{‰}$ to $366\pm 35\text{‰}$; this represents an extreme variation over 764‰ . Paris olivines and
23 pyroxenes are either enriched or depleted in deuterium relative to the mesostasis and no systematic
24 isotopic pattern is observed. Simple model of chondrules hydration during parent body hydrothermal
25 alteration is difficult to reconcile with such isotopic heterogeneity. It is proposed that a hydrous
26 component, having a δD of c.a. -400‰ , in the chondrule precursors, has been outgassed at 800-900°C in
27 the gas phase. Nevertheless, a residual water fraction remains trapped in Paris chondrules. Quantitative
28 modeling supports this scenario.

29 30 31 1. Introduction

32
33 Some carbonaceous chondrites (CC) hold hydrated minerals and contain up to 10 wt.% H₂O (Kerridge,
34 1985). They are the most plausible source of water delivery to the Earth (Marty, 2012; Walsh et al., 2011)
35 as they display, on average, a similar D/H ratio as the Earth oceans, i.e. 155.76×10^{-6} (Robert, 2006). Thus,
36 the carbonaceous chondrites are the key samples to constrain the origin of water on Earth. CM
37 carbonaceous chondrites constitute an interesting class for studying aqueous alteration. Indeed, the CM
38 chondrite group embraces type 1 (i.e. heavily aqueously altered) and type 2 (i.e. less aqueously altered)
39 chondrites. Manganese-chromium dating of carbonates in CM chondrites gave a reliable aqueous
40 alteration age of $4563.4\pm 0.4/-0.5$ million years (Fujiya et al., 2012); younger age than previous
41 estimations (de Leuw et al., 2009). This CM carbonate age suggests the late onset of aqueous activities in
42 the Solar System and allows Fujiya et al. (2012) to estimate that CM parent body accreted roughly 3.5
43 Myr after the birth of the Solar System.

44 Hydrated minerals and organic matter carry hydrogen in aqueously altered chondrites. Aqueous alteration
45 of silicate minerals in these chondrites is evidenced by the occurrence of hydrous phyllosilicates, as
46 serpentines, saponites, and smectites, which are the predominant phase in matrices of CCs (Garenne et al.,
47 2014; Beck et al., 2010), often associated with carbonates and oxides (Howard et al., 2011) The petrologic
48 chondrite classification reflects varying degrees of aqueous alteration in different groups of chondritic
49 meteorites (Van Schmus and Wood, 1967). Classification of CMs is established based on a sequence of
50 aqueous alteration (Alexander et al., 2013; Browning et al., 1996; Rubin et al., 2007).

51 Whether alteration on CM chondrite took place in the protosolar nebula or on the parent body is still a
52 matter of debate: Ciesla et al. (2003) have argued for the hydration of silicate dust in the icy regions of the
53 nebula during the passage of shock waves and Bishoff (1998) favors pre-accretionary alteration in the
54 nebula, based on the relationship between fine-grained rims and their surrounding components. On the
55 other hand, several observations support the hypothesis that the aqueous alteration occurred after the
56 accretion onto the parent body (Brearley, 2003). For example, the existence of alteration zones, as Fe-rich
57 aureoles around metal grains (Hanowski and Brearley, 2001), fluid inclusions (Fieni et al., 1978; Zolensky
58 et al., 1999), supports this hypothesis. The oxygen isotopic data also indicate a parent body alteration:
59 Young et al. (1999) explained the oxygen isotopic diversity of CC by simple down-temperature fluid flow
60 within the parent body, based on the systematic relationships between the bulk oxygen isotopic
61 composition of CC and the degree of aqueous alteration.

62 Chondrites show a systematic enrichment in deuterium relative to the molecular hydrogen of the
63 protosolar nebula (Boato, 1954). The D/H ratio in matrix clays of chondrites has been assumed to reflect
64 the D/H ratio of water that has circulated during the aqueous alteration of the parent body. Based on
65 petrographic and mineralogical analyses in the heavily aqueously-altered Alan Hills 81002 CM2
66 chondrite, Hanowski and Brearley (2001) argued for the alteration of chondrules in a parent body
67 environment. Nevertheless, hydrogen isotopic measurements performed on the Bishunpur and Semarkona
68 LL3 ordinary chondrites and Renazzo CR2 carbonaceous chondrite showed large D/H heterogeneities
69 among matrix clays and also between silicate minerals in chondrules (Deloule and Robert, 1995; Deloule
70 et al., 1998). An isotopic reequilibration during hydrothermal alteration on the parent body by a
71 homogenous aqueous reservoir cannot explain alone the heterogeneities of D/H ratio in these chondrites.
72 D-enrichment of chondrite water could result from 2 scenarios (Bonafant et al., 2013): (i) enrichment is an
73 heritage of water formed in the outer solar system or presolar molecular clouds (Deloule and Robert,
74 1995; Deloule et al., 1998), variations observed in chondrites reflect different water reservoirs having
75 different D/H ratios ; (ii) light water was formed in the warmer inner solar system and got enriched by
76 secondary processes during aqueous alteration on the parent body including exchange with D-rich
77 organics, oxidation of metal or other processes able to fractionate the initial D/H of water (Eiler and
78 Kitchen, 2004; Alexander et al., 2012)

79 The CM2 Paris is a fresh sample that suffered only limited parent-body transformations, as revealed by
80 preservation of amorphous silicates in the matrix. It is considered as the least altered CM chondrite
81 currently known, classified as a CM2.7 (Hewins et al., 2014; Marrocchi et al., 2014). This sample offers a
82 unique opportunity to study the initial stages of aqueous alteration on the CM parent body by investigating
83 well preserved type I chondrules, which are not prone to aqueous alteration and are the least altered
84 component in CM chondrites. Moreover, Paris exhibits heterogeneous alteration areas (Hewins et al.,
85 2014), which bring us a large set of chondrules ranging from low to high alteration rate. Here we

86 investigated the water content and the D/H ratio in different phases (olivine, pyroxene and mesostasis) of
87 12 type I chondrules and the surrounding matrix from Paris. In comparison, 4 chondrules from Bishunpur
88 (LL3) and 3 chondrules from Renazzo (CR2) were also analysed. Our aim was to determine if chondrules
89 and matrix underwent the same aqueous alteration, by using the D/H ratio, or if they were subjected to
90 different hydration episodes. The high lateral resolution of NanoSIMS allows measuring water content
91 and hydrogen isotopic ratio in contiguous olivine, pyroxene and mesostasis of chondrules.

92 2. Samples

93 2.1. Objects of study

94 The Paris chondrite is a breccia with varying degrees of alteration and thus it contains both well preserved
95 and altered chondrules. Three sections were analyzed in three sessions: one mounted in epoxy (i.e. Session
96 I- chondrules 4, 6, 9, 19, 24 and 27) and two epoxy-free (i.e. Session II: chondrules 1 and 7; Session III:
97 chondrules A, B, D and E). The latest were chosen to assess vacuum issues on the hydrogen
98 contamination in addition to the effect of epoxy -expected to impregnate the sample at a fine scale- that
99 could bias our measurements (Aubaud et al., 2007; Hauri et al., 2006; Stephant et al., 2014). The three
100 sections of Paris were used for D/H analyses but only the one of session III were used for water content
101 measurements. Renazzo and Bishunpur, mounted in epoxy, were also analysed as they exhibit
102 heterogeneities of the D/H ratios in their chondrules, documented by previous studies (Deloule and
103 Robert, 1995; Deloule et al., 1998) (i.e. Session IV – chondrules X, Y, Z and I, J, K, L). As such, they are
104 used as references in this study. Renazzo is a CR2 chondrite which contains abundant type I chondrules
105 (Weisberg et al., 1993) with intermediate level of alteration, evidenced by the replacement of chondrule
106 mesostases by phyllosilicates. Bishunpur is a LL3.1 ordinary chondrite whose aqueous alteration was
107 limited to the production of smectite in the matrix (Alexander et al., 1989). Type I chondrules were
108 imaged using a JEOL JSM 840-A SEM installed at MNHN. Chemical characterization of olivine and
109 pyroxene from 6 Paris chondrules (in the epoxy mount) were carried out using a CAMECA SX-100
110 EMPA at the Camparis facility, in Paris, France.

111 Paris contains markedly more type I than type II chondrules and their size ranges from ~100 to ~500 μm .
112 In addition, type I chondrules in Paris have various petrologic characteristics. We selected 12 chondrules
113 of different textural types and various degrees of aqueous alteration in order to survey the range of
114 chondrules present in Paris. Backscattered electron images of these 12 chondrules are presented in Figure
115 1 to illustrate this diversity: barred olivines (chondrules 19, 24 and 1), radial pyroxene (chondrule 4), and
116 porphyritic chondrules. Degree of chondrule aqueous alteration is established on the amount of metal
117 globules, glassy stage of mesostasis and presence of alteration aureole. Moreover, various degrees of
118 aqueous alteration are observed in chondrule mesostasis. For instance, chondrule 19 is a well-preserved
119 object with glassy mesostasis and metal globules whereas chondrule 27 exhibits an altered mesostasis in
120 its center and lacks of Fe-rich globules (Fig.1). X-ray elemental maps obtained by EDX allow us to select
121 the target areas for NanoSIMS analysis; we aimed at areas with contiguous occurrences of olivine,
122 pyroxene and mesostasis. The EMPA measurements (Table 1) indicate that all olivines type I chondrules
123 from Paris are magnesium-rich ($Fa < 2$). Low-Ca pyroxenes show limited compositional range (Fs ranges
124 from 0.7 to 17.5; Wo ranges from 0.4 to 5.9 with an exception at 38.3). We also chose 4 representative
125 type I chondrules in Bishunpur and 3 in Renazzo.

126 2.2. Standards

127 Silicate standards were used for calibration of water contents and for instrumental mass fractionation on
128 D/H ratio. These standards match different phases analyzed in chondrules: pyroxene KBH1, olivine KLV-
129 23, glass DR15-1-5, glass DR20-1-1, amphibole Mont Emma and amphibole Kipawa (Clog, 2010;
130 Deloule, 1991; Koga et al., 2003). Their chemical compositions are reported in Table 2. Samples were
131 mounted in such a manner to reduce a contribution of epoxy to the OH signal while having a nice polished
132 surface: holes were drilled in a 10 mm aluminium disk with a 2 mm diameter drill bit. In each hole,
133 samples were mounted individually with epoxy and then polished down to 0.25 μm diamond paste.
134 Mounts were carbon-coated, thickness 20 nm. Chondrites and standards were outgassed for one week in
135 the NanoSIMS airlock before each analytical session.

136 3. Methods

137 3.1 Quantification of water content

138 Elemental measurements were performed with the Cameca NanoSIMS 50 at MNHN, Paris. $^{16}\text{OH}^-$, $^{28}\text{Si}^-$,
139 $^{24}\text{Mg}^{16}\text{O}^-$ and $^{27}\text{Al}^{16}\text{O}^-$ secondary ions were imaged and quantified using a Cs^+ primary beam. Special care
140 was dedicated to correct the contribution of water contamination: we analyzed consecutively the same
141 areas with two different primary beam intensities, 1.2pA and then 15pA, in order to estimate the part of
142 OH contamination (Stephant et al., 2014). Indeed, using a correlation between the primary beam intensity
143 and the OH/Si $^-$ ratio, we calculated the OH/Si $^-$ at the “infinite” primary beam intensity, supposed free of
144 any OH contamination (see supplementary information for raw data). The primary beam was rastered over
145 a $20\times 20\mu\text{m}^2$ surface area divided in 256×256 pixels with a $200\mu\text{m}$ aperture diaphragm for 60 cycles. The
146 counting time was set to 1ms/px. Prior to the analysis, a $25\times 25\mu\text{m}^2$ surface area was presputtered for 20
147 minutes using a high primary beam current (i.e. 280pA and 230pA for the session III and IV, respectively)
148 and a $750\mu\text{m}$ aperture diaphragm. The mass resolution power was ~ 6100 for $^{16}\text{OH}^-$ and ~ 7500 for $^{28}\text{Si}^-$
149 using the entrance slit ES3 and the aperture slit AS2. 39 areas were imaged in type I chondrules from Paris
150 during session III, 14 from Bishunpur and 11 from Renazzo during session IV. Vacuum in the analytical
151 chamber was 9.3×10^{-10} Torr for session III and 1.1×10^{-9} Torr for session IV.

152 Water concentrations in the different phases of chondrules were obtained using a calibration of OH/Si $^-$
153 versus water concentration; water concentration is reported as $[\text{H}_2\text{O}]$ and represents H_2O ppm / SiO_2 wt. %.
154 The calibration lines were forced to the origin: using our method to estimate the OH/Si $^-$ ratio at infinite
155 primary beam intensity, the surface contamination contribution on the OH/Si $^-$ value is withdrawn (see the
156 discussion of this calibration in Stephant et al., 2014). Different calibrations were used depending on the
157 silicate mineralogy (Tenner et al., 2009) (Fig. 2): matrix effects are so strong that ferromagnesian minerals
158 and amphiboles show similar OH/Si $^-$ ratios for very distinct water concentrations (less than 200 ppm up
159 for the former to compare with 1 wt. % for the latter). Thus it is crucial to precisely identify each analyzed
160 mineral to accurately quantify its water content. To this end, $^{24}\text{Mg}^{16}\text{O}^-/^{28}\text{Si}^-$ and $^{27}\text{Al}^{16}\text{O}^-/^{28}\text{Si}^-$ are used to
161 decipher between olivine, pyroxene and mesostasis (see supplementary information). Standard
162 measurements for each session of analysis (i.e. III and IV) are reported in supplementary information.
163 Uncertainties for each calibration line were evaluated using a code written on the R program (Thomen et
164 al., 2014). The reported errors on meteorite measurements are dominated by uncertainties on the
165 calibration slopes. The counting statistics on the OH/Si $^-$ ratio are negligible compared to the calibration
166 slope uncertainties.

167 3.2 Hydrogen isotopic ratios

168 Hydrogen isotopic measurements were performed by imaging H^+ and D^+ over the same areas studied for
169 elemental analyses to provide both isotopic and elemental data at exactly the same location. Thereby, D/H
170 images could be superimposed on elemental images. The primary beam was rastered over a $20 \times 20 \mu m^2$
171 surface area divided in 256×256 pixels with a $300 \mu m$ aperture diaphragm for 60 cycles with a counting
172 time set to 1ms/px. The primary beam intensities was ~ 200 - $240 pA$ for the three sessions on Paris
173 chondrite and $29 pA$ for the session IV (i.e. Bishunpur and Renazzo). Prior to each analysis, a $25 \times 25 \mu m^2$
174 surface area was presputtered for 5 minutes using a high primary beam current (i.e. 230 - $280 pA$) with the
175 $750 \mu m$ aperture diaphragm. The entrance and the aperture slits were both set at the position 1. The
176 instrumental isotopic fractionation factor for hydrogen was determined with the set of silicate standards
177 previously described at each session (supplementary information). During sessions I, II and III, no matrix
178 effect was noticed for the D/H ratio using NanoSIMS at the precision of our measurements (Hu et al.,
179 2015). However, in the session IV, a significant difference in the instrumental isotopic fractionation was
180 observed for the amphibole and the glass standards. That can be due to the lower primary beam intensity
181 during this session compared to the three prior sessions. The instrumental isotopic fractionation factor was
182 also calculated as the average of these standards. The uncertainty on the instrumental isotopic
183 fractionation corresponds to the standard deviation of terrestrial standards (see supplementary
184 information). Errors at the 2σ level (quadratic sum) combine the uncertainty on the instrumental mass
185 fractionation factor and the counting statistic error on each measurement (dominated by the total number
186 of counts detected for D^+). The D/H ratios are expressed in δ units (where $\delta D = [(D/H_{\text{sample}})/(D/H_{\text{SMOW}}) - 1]$
187 $\times 1000$).

188 4. Results

189 4.1 Water concentration in nominally anhydrous minerals (NAMs)

190 Distribution of water concentrations in olivines and pyroxenes of type I chondrules (Table 3) are presented
191 in Figure 3 for each chondrite. Mean water concentrations in pyroxenes are 893 ± 637 ppm (1SD, $n=4$),
192 879 ± 536 ppm (1SD, $n=11$) and 791 ± 227 ppm (1SD, $n=8$) for Renazzo, Bishunpur and Paris, respectively.
193 In olivine, water concentration varies markedly between Renazzo, Bishunpur and Paris: 156 ± 44 ppm
194 (1SD, $n=7$), 222 ± 123 ppm (1SD, $n=7$) and 860 ± 132 ppm (1SD, $n=14$), respectively. Relative uncertainties
195 for these measurements are $\pm 25\%$ for Paris NAMs and $\pm 50\%$ for Bishunpur and Renazzo NAMs
196 (supplementary information). It must be noted that significant uncertainties in the determinations of the
197 water content in NAMs are due to the fact that our standards – being terrestrial nominally anhydrous
198 minerals - have much lower water concentrations than those measured in Paris chondrules. This implies a
199 large extrapolation towards large values hence larger uncertainties. Using the uncertainties on the NAMs
200 water calibration line, the lowest estimations for pyroxene mean water concentrations become 447 ± 318
201 ppm, 439 ± 268 ppm and 557 ± 147 ppm for Renazzo, Bishunpur and Paris, respectively. Regarding the
202 olivines, the lowest estimations for mean water concentrations are 78 ± 22 ppm, 111 ± 61 ppm and 645 ± 99
203 ppm, for Renazzo, Bishunpur and Paris, respectively. Therefore, Paris olivines appear to be more hydrated
204 than those in Bishunpur and Renazzo chondrules and do contain a significant amount of water for a
205 nominally anhydrous mineral. For pyroxenes, Bishunpur and Renazzo exhibit a more scattered
206 distribution than Paris. The small numbers of chondrule analyzed for Bishunpur (i.e. 4) and Renazzo (i.e.
207 3) compared to Paris doesn't allow us to make a robust statement on the water content of these two
208 meteorites. Moreover, water concentrations in Paris chondrules were determined in epoxy-free sections

209 (session III), allowing us to rule out a significant contribution from the epoxy in the OH signal and a
210 subsequent bias in the determination of the water content (Aubaud et al., 2007).

211 4.2 Hydrogen isotopic ratio in chondrules

212 4.2.1 Comparison between Paris, Renazzo and Bishunpur

213 Distribution of the hydrogen isotopic composition of minerals in chondrules in Paris, Renazzo and
214 Bishunpur (Table 3) are shown in Figure 4. The mean δD value for these three meteorites fall within the
215 terrestrial domain (Lécuyer et al., 1998). Paris exhibits a wide spread of δD from $-398 \pm 23\text{‰}$ to $366 \pm 35\text{‰}$
216 (i.e. covering a range of 764‰) with a mean value of $-88 \pm 141\text{‰}$ (2SD, $n=70$). In Paris, the δD of the
217 surrounding matrix is $-69 \pm 32\text{‰}$. The possible contamination of the Paris samples by terrestrial moisture
218 can be ruled out since chondrules show both enrichment and depletion of deuterium relative to the
219 terrestrial value. The Paris mounts without epoxy (i.e. session II and III; chondrules A, B, D, E, 1 and 7)
220 shows the largest range of δD values. In Renazzo, the δD in olivines, pyroxenes and mesostases range
221 from -212 ± 125 to $15 \pm 156 \text{‰}$ with a mean value of $-131 \pm 115\text{‰}$ (2SD, $n=21$). An image recorded in the
222 matrix gives a value of $316 \pm 202\text{‰}$. These determinations are consistent with previous studies in Renazzo
223 (Deloule and Robert, 1995) showing that the matrix is enriched in deuterium (-37 to 733‰) relative to
224 chondrules (-294 to 284‰). Bishunpur minerals range from $-166 \pm 133\text{‰}$ to $137 \pm 176\text{‰}$ with a mean value
225 of $-66 \pm 148\text{‰}$ (2SD, $n=16$). This range is significantly smaller compared to previous studies on Bishunpur
226 chondrules: $-371\text{‰} < \delta D < 1934\text{‰}$ (Deloule and Robert, 1995; Deloule et al., 1998). We will focus our
227 attention in the upcoming discussion on the emblematic case of Paris for which the δD variations reach
228 764‰ .

229 4.2.2 Inter-chondrule variations in Paris

230 A broad inter-chondrule isotopic distribution is observed in Paris: in the mesostasis $-292 \pm 32\text{‰}$
231 $< \delta D < 366 \pm 35\text{‰}$ with a mean value $\delta D = -44 \pm 346\text{‰}$ (2SD, $n=26$), in olivines $-398 \pm 23\text{‰} < \delta D < 76 \pm 36\text{‰}$
232 with a mean value of $-136 \pm 258\text{‰}$ (2SD, $n=24$) and in pyroxenes $-342 \pm 23\text{‰} < \delta D < 155 \pm 53\text{‰}$ with a mean
233 value of $-84 \pm 250\text{‰}$ (2SD, $n=20$). The δD in Paris chondrules spans over a range of 764‰ ; this
234 corresponds to a large isotopic fractionation that is never observed in terrestrial samples. Pyroxenes and
235 olivines exhibit the same range. Mesostasis seems to be slightly enriched in deuterium compared to
236 adjacent olivine and pyroxene. However, individual chondrules exhibit more complex variations as
237 discussed hereafter.

238 4.2.3 Intra-chondrule variations in Paris

239 In addition to the broad range of δD observed among chondrules, intra-chondrule heterogeneities are also
240 pronounced in Paris. Indeed, an isotopic variation of 658‰ is observed between the different phases of the
241 chondrule E and up to 664‰ for the chondrule A. More surprisingly, isotopic variation among olivines in
242 chondrule A reaches 454‰ in less than $\sim 200 \mu\text{m}$. In Figure 5, NanoSIMS images of AlO^+/Si^- and D/H of a
243 $20 \times 20 \mu\text{m}^2$ area in the chondrule A are shown. Olivine and mesostasis that are contiguous exhibit a
244 difference in δD of 317‰ with no visible gradient (the NanoSIMS resolution does not exceed $\sim 500\text{nm}$). In
245 most cases, the mesostasis is enriched in deuterium compared to olivine and pyroxene (chondrules A, B,
246 D, E, A, 7, 9) with the exception of chondrules 24, 6 and 27. As such, no systematic between mesostases
247 δD and silicates δD is observed, suggesting there is no equilibrium between mesostasis and silicates

248 minerals. Furthermore, no correlation is observed between the degree of aqueous alteration experienced by
249 the chondrule, as revealed by its petrology, and the δD distribution of the chondrule. In the less altered
250 chondrule (19), there is no variation, within errors, of the δD value between the mesostasis and olivine and
251 pyroxene (i.e. $-34 \pm 79\%$ on average, 2SD). However, chondrules A, E and 7 exhibit a low degree of
252 aqueous alteration illustrated by abundant metal grains and they show a large isotopic heterogeneity
253 between their minerals. In addition, chondrule 27 is heavily altered (its mesostasis is altered into
254 phyllosilicates) while δD variations do not exceed 232% between mesostasis and minerals. An important
255 point to notice is that chondrules A, E and 1, showing the largest range of δD values, were all measured in
256 epoxy-free mounts. As a conclusion of this isotopic survey, δD variations are intrinsic to each chondrule
257 but not related to their degree of aqueous alteration.

258 5. Discussion

259 5.1 Implication of the δD distribution for chondrule formation models

260

261 5.1.1 δD fractionation among coexisting minerals

262 The isotopic fractionation coefficient α between two phases corresponds to the ratio of their D/H values.
263 Figure 6 reports the isotopic fractionation coefficients α between olivine and mesostasis versus those
264 between pyroxene and mesostasis, reported as $10^3 \times \ln \alpha$ in per mil. Fractionation coefficients are calculated
265 for adjacent phases (see supplementary information). Therefore, each point corresponds to a NanoSIMS
266 image containing the 3 phases in a $20 \times 20 \mu\text{m}^2$ surface area. A trend is observed with both positive and
267 negative values. The slope of the correlation is not far from unity, indicating that pyroxene and olivine
268 have similar δD . The important point to notice is the large range of fractionation coefficient values (i.e.
269 $10^3 \times \ln \alpha$ ranging for -431% to $+283\%$). This correlation requires a process that would generate a large
270 fractionation factors between minerals during chondrule history. Furthermore, this correlation has no link
271 to the degree of aqueous alteration and is not simply related to hydrothermal alteration, which may imply
272 that chondrules followed different fractionation history.

273 5.1.2 Low-temperature and hydrothermal alteration on the parent body under isotopic equilibrium 274 conditions

275
276 The observed fractionation factors are too large to account for an isotopic fractionation at equilibrium
277 during H isotopic exchange between H-bearing minerals and water. Significant intra-chondrules
278 heterogeneities among the same phases are up to 764% . In terrestrial environments, such variations are
279 never observed among alteration products and fractionation factors between minerals and water are
280 restricted to 60% (Saccocia et al., 2009). This should be regarded as clear evidence against a simple
281 equilibrium isotope fractionation between phases of the same chondrule.

282 5.1.3 Modelling the chondrule alteration by an isotopically homogeneous fluid.

283
284 NAMs (i.e. olivine and pyroxenes) are either enriched or depleted in D relative to the mesostasis (Fig. 6).
285 This suggests there is no equilibrium between mesostasis and these minerals. This results in a crossover
286 rarely observed in natural environment, which should be regarded as a dependable constraint for
287 chondrules formation models. We first discuss common geochemical situations that cannot yield such a
288 crossover.

289 The progressive isotopic re-equilibration of the minerals with the mesostasis by diffusion of an external
 290 fluid to the chondrule cannot account for the isotopic crossover observed in the Figure 6 since isotopic
 291 fractionation will be unidirectional (Kyser and O'Neil, 1984). In other terms, the expected correlation for
 292 diffusion in the Figure 6 is either enriched or depleted relative to 0 but never on both sides.

293 However, we considered here diffusion by a reservoir of water with homogeneous isotope composition.
 294 Bonal et al. (2013) have shown that the δD of the matrix water is highly heterogeneous at the micrometer
 295 scale in CR chondrite. Piani et al. (2015) have shown that water-ice precursors with a large range of δD
 296 compositions could have accreted into the parent body and been responsible for matrix δD heterogeneities
 297 at the micron scale. Secondary processes on the parent body have also been suggested to explain this
 298 spatial variability (Alexander et al., 2010; Alexander et al., 2012). This δD spatial heterogeneity at a small
 299 scale could potentially have induced the heterogeneities observed among single chondrules. However, this
 300 model holds three issues. (1) The fractionation coefficients between ferromagnesians (i.e. olivine and
 301 pyroxene) and mesostasis show significant variations (cf. Fig. 6). This means that the δD of the mesostasis
 302 can be distinct from that in the pyroxene or olivine next to this mesostasis. Even if the aqueous alteration
 303 occurs at the micrometer scale, phases having common boundaries should record a similar δD value; plus
 304 or less the isotopic fractionation at equilibrium between phases (Saccocia et al., 2009). (2) The δD
 305 measured in the matrix covers a more restricted range (i.e. $-66 \pm 70\%$; 2SD) compared to the broad
 306 distribution of δD in chondrule phases; (3) This model does not account for the high water content in
 307 olivines, as it seems quite unrealistic to incorporate a minimum of 645 ppm H_2O inside an olivine at
 308 chondritic parent body conditions of pressure and temperature. This problem will be discussed in the next
 309 section.

310 Hence, hydration of the chondrule by parent body water does not fit the requirements to explain the
 311 isotopic crossover observed in the Figure 6. However, as shown in the next section, it may be explained by
 312 the evaporation of H before the formation of the parent body, provided that this H was present in
 313 chondrules at the moment of their formation.

314 5.2.3. Loss of H from chondrules during their cooling in the gas phase.

315 Here we propose that distillation of H took place during the melting and the cooling of the chondrule in
 316 the gas phase. This outgassing model has two implications: (1) the initial [H] concentration and (2) the
 317 degree of distillation caused by H_2 escape can vary from chondrule to chondrule. As the analyzed
 318 chondrules show various textural and petrological features, there is no reason that they share common
 319 initial water content and that they underwent the same degree of distillation.

320
 321 Rayleigh distillation induced by the loss of H_2 from melted chondrules can be written as:
 322

$$D/H(t)_{phase} = D/H(0)_{phase} \times f_{phase}^{(\alpha_{H_2-H_2O})^{-1}}$$

323
 324 With $D/H(0)$ the initial hydrogen isotope ratio of the mineral; $D/H(t)$ the hydrogen isotopic ratio at instant
 325 t and f_{phase} is the remaining fraction of H in the mineral (i.e. from 1 to 0); $\alpha_{H_2-H_2O}$ is the isotopic
 326 fractionation between H_2 and H_2O in phases. The lowest measured value on chondrule phases i.e. for $\delta D =$
 327 -398% is used to set $D/H(0)_{phase}$. We assume that each phase follows its own f Rayleigh coefficient i.e.

328 that NAMs and mesostasis evolve independently since they do not share the same level of advancement of
329 the diffusion for H species (Ingrin and Blanchard, 2006) and initial water concentration. Indeed,
330 mesostasis can be either depleted or enriched relative to olivine and pyroxene, function of their relative
331 degree of degassing. Calculations are reported in Figure 6 for f_{phase} values randomly generated for
332 mesostasis and for NAMs. These calculations reproduce the wide range of isotopic fractionation variations
333 with $\alpha_{H_2-H_2O}$ value of ≈ 0.9 for olivine and pyroxene and ≈ 0.85 for mesostasis. These α values are
334 markedly different from unity and therefore consistent with the large H₂-H₂O isotopic fractionation factor
335 (Simon et al., 2011). From these α values, we calculate a reducing temperature of the system during the
336 isotope fractionation around 900-1100 °C (Simon et al., 2011).

337
338 An alternative scenario could involve the reduction of water during parent body processes (Alexander et
339 al., 2010). Lécluse and Robert (1994) showed that the isotope exchange between H₂O and H₂ is still fast
340 down to 200-300°C, i.e. at hydrothermalism temperatures. If we consider that a distillation took place
341 around 150°C in the parent body (Clayton and Mayeda, 1999; Guo and Eiler, 2007), the fractionation
342 factor between H₂ and H₂O should be far ($\alpha_{H_2-H_2O} \approx 0.3$, by extrapolating from Simon et al., 2011) from
343 our calculated value ($\alpha_{H_2-H_2O} \approx 0.9$). In other terms, it seems unrealistic to argue that water reduction took
344 place on the parent body while the hydrogen isotopic system registered temperatures around 900-1100 °C.
345 We thus propose that chondrules have degassed their juvenile water prior to their incorporation in their
346 parent body, presumably in the nebula, following the reduction of their water at temperature around 900-
347 1100°C.

348
349 The lack of correlation between water content and the δD demonstrates that each chondrules contained
350 heterogeneous initial water concentrations or δD or both. As analyzed chondrules are of different textural
351 types, we can assume that each chondrule gets its own initial water content.

352 353 5.2 Incorporation of high water contents in NAMs

354 5.2.1 Data on Earth

355 It is now established that nominally anhydrous minerals are a major host for water in terrestrial mantle
356 (Bell and Rossman, 1992; Bolfan-Casanova, 2005; Hirschmann et al., 2005; Peslier, 2010). Martin and
357 Donnay (1972) were the first to propose a water incorporation model based on OH ions replacing O in the
358 structure of pyroxenes. Water storage has been extensively studied in NAMs. Natural pyroxene water
359 concentrations range from few 10s up to ~ 1100 ppm of H₂O, depending on the geological occurrence
360 (Rauch and Keppler, 2002; Skogby, 2006; Sundvall and Stalder, 2011). Natural olivine contains
361 significantly less water than pyroxene with contents ranging roughly between few and 400 ppm of H₂O
362 (Bell and Rossman, 2003; Beran and Libowitzky, 2006). Measured water concentrations in NAMs of Paris
363 chondrules (mostly olivines) are markedly higher than those in terrestrial ones. Thus water contents
364 measured in olivine seem too high to come only from water in olivine.

365 However, several experiments show that olivine can incorporate significant amount of water at high
366 pressure, high temperature (Kohlstedt et al., 1996; Mosenfelder et al., 2011; Ohtani et al., 2001). Férot and
367 Bolfan-Casanova (2012) measured olivine water contents of 4690 ± 1655 ppm for experiments performed
368 at 9 GPa and 1175 °C. Smyth et al. (2006) found water solubility in olivines up to 8900 ppm at 12 GPa
369 and 1250°C.

370 OH-rich inclusions in olivines and pyroxenes of magmatic origin could account for the large content we
371 determine. For instance, Khisina et al. (2001) have shown that hydrogen incorporation in olivine exists in
372 two modes: intrinsic mode into olivine structure (Wright, 2006) and extrinsic incorporation as nanometer-
373 sized inclusions of OH bearing silicates. In pyroxene, hydroxyl can occur in the form of narrow lamellae
374 of amphibole (Skogby et al., 1990; Veblen and Buseck, 1981).

375 5.2.2 Implications for chondrule olivines of Paris

376 From the experimental data on Earth's olivine, the striking water contents measured in Paris olivines
377 could suggest that olivines were formed at high-pressure, high-temperature conditions in order to
378 incorporate such amount of water; assuming that this water is under the form of hydroxylated groups in
379 the structure of olivines. The model proposed by Libourel and Krot (2007) suggested that olivines are
380 inherited from mantle planetesimals fragmented by impacts. In this scenario, mantle olivine can
381 incorporate large amounts of water under conditions of high pressure and high temperature. As suggested
382 by Libourel and Krot, pyroxenes could result from the destabilization of olivine caused by the
383 incorporation of SiO diffusing from the gas phase. Qualitatively, this model accounts for the similar δD
384 values in the pyroxene and in its parent olivine. Nevertheless, there are restrictions to this hypothesis.
385 First, this scenario by Libourel and Krot is based on chondrules having a granoblastic texture that is not
386 found in all our studied chondrules. Secondly, the kind of high pressure needed to incorporate a minimum
387 of 645 ppm of water in olivine seems difficult to reconcile with a small parent body.

388 The speciation of hydrogen is not known through NanoSIMS analyses and it is limited by its spatial
389 resolution (100nm). Nanometer-sized inclusions cannot be identified by NanoSIMS imaging. We assumed
390 that the hydrogen occurs under the form of hydroxylated groups in the chondrule olivines. However, glass
391 inclusions in olivines can be found in any type of chondrules (Varela et al., 2002; Florentin et al., 2016).
392 These inclusions occur at the micrometer scale and their water contents should be more similar to the
393 mesostasis one, in the range of few wt.% H₂O. Indeed, such inclusions should appear on our NanoSIMS
394 images of AlO⁻/Si⁻, MgO⁻/Si⁻ and OH⁻/Si⁻, as we have a resolution down to 100 nm at 1.2 pA. We do not
395 observe any variations in these three ratios during the duration of our analyses, meaning that we do not
396 cross any glass inclusions of this type.

397 We propose that type I chondrules precursors could be composed of glassy grains, clays minerals or
398 altered ferromagnesians, i.e. hydrous matrix-like material. These precursors would have been
399 incompletely degassed after chondrule cooling and before incorporation into the parent body. Water
400 should have been incorporated into chondrules before their accretion into the parent body, as it seems
401 unrealistic to add a large amount of water in chondrule olivines at parent body conditions.

402 5.3. Open questions

403 The Rayleigh distillation model seems to indicate that chondrules would have degassed their primordial
404 hydrogen, distinct from the alteration water, resulting from reduction of water into H₂ at $\approx 900^{\circ}\text{C}$. Still,
405 there are some interesting unsolved issues:

- 406 (i) The primordial source of hydrogen in chondrules is unknown. The hydrogen isotopic ratio
407 of this source could be as low as -398‰, i.e. the lower δD value measured in chondrule
408 phases. This value is markedly distinct from both the nebula H₂ and the mean chondrite

hydrogen isotopic signature. There are few components with such low δD among solar system objects. The unique organic matter contained in the Abee meteorite in one of them (Remusat et al. 2012; Yang and Epstein 1983). This material signs the occurrence, in the early solar system, of a H-reservoir that was at isotopic equilibrium with molecular H_2 in the neutral environment (Lécluse and Robert, 1994). This source of H could also be responsible of the end-member δD value advocated to interpret the linear variations observed at the scale of the bulk CCs in a D/H vs. C/H diagram: at C/H=0, $\delta D = -444 \pm 23\%$ (Alexander et al., 2012). According to the present study, in bulk CCs, organic free minerals (chondrules) could be the carriers of a H-reservoir with a low δD .

- (ii) This model implies that NAMs had a higher water concentration before the distillation since H_2 was lost from chondrule. Therefore chondrule precursors were hydrated before their melting/crystallization into the form of chondrules and were incompletely outgassed during the rapid cooling of the chondrules. Incorporation of high water content could have been achieved because chondrule precursors included hydroxylated minerals.

It should be emphasized that the presently proposed distillation model does not preclude a late alteration (hydration) of chondrules by circulating parent body water causing a limited homogenization of the initial chondrule δD values. Indeed, - if the fluid was isotopically homogenous - the isotope fractionation between minerals and water is limited to 50 ‰ and thus should not have significantly altered the observed chondrule δD distribution that spans over 764‰.

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604

605

606 **Figure captions**

607 Figure 1: SEM images of 12 type I chondrules in Paris from the section with epoxy - Chondrules 1, 19 and
 608 27: barred olivine. Chondrule 19 exhibits a well-preserved glassy mesostasis. Chondrule 4: radial
 609 pyroxene. Others chondrules are porphyritic. They presented various stages of aqueous alteration. Some
 610 of them like chondrule 6 or D exhibit well-preserved Fe grains, sign of low aqueous alteration. In the other
 611 hand, chondrule 27 has its mesostasis transformed into phyllosilicates and an aqueous alteration rim grew
 612 around him.

613 Figure 2: OH^-/Si^- vs $\text{H}_2\text{O}/\text{SiO}_2$ (ppm/wt.%) calibration for the quantification of water contents. Dashed
 614 lines are regression lines calculated with the R program.

615 Figure 3: Water content distribution in olivine and pyroxene chondrule minerals of Renazzo, Bishunpur
 616 and Paris chondrites. Y axis represents the cumulate number of analyses.

617 Figure 4: Distribution of δD values in the olivines, pyroxenes and mesostases of Renazzo, Bishunpur and
 618 Paris chondrules.

619 Figure 5: NanoSIMS images of an area ($20 \times 20 \mu\text{m}^2$) of Paris chondrule A. (a) $^{27}\text{Al}^{16}\text{O}^-$ image permits to
 620 distinguish mesostasis (Mes) and olivine (Ol). (b) D^-/H^- ratio image. We defined ROIs for each phase
 621 (blue lines) that fit in both images. The ROI is chosen smaller than the phases due to the different image
 622 resolution and the possible shifting occurring between the elemental images and the isotopic image.

623 Figure 6 : Fractionation coefficients between olivine and mesostasis versus pyroxene and mesostasis
 624 reported in per mil. Black square symbols represent our analyses and gray square symbols stand for the
 625 distillation numerical simulation explained in section 4.2.3. Arrows indicate direction of fractionation for
 626 hydrous alteration of the Paris chondrules, which is unidirectional. The direction will depend on the
 627 alteration parameters.

628

629 **Table captions**

630 Table 1 – EMPA analyses of olivines and pyroxenes from 6 Paris chondrules.

631 Table 2 – Chemical compositions of standards: (a) from Koga et al. (2003), (b) from Clog (2010), (c)
 632 from Deloule et al. (1991).

633 Table 3 – NanoSIMS data of water concentrations and hydrogen isotopic ratios (expressed in δD) of
 634 mesostases, olivines and pyroxenes from 3 Renazzo chondrules (session IV), 4 Bishunpur chondrules
 635 (session IV) and 12 Paris chondrules (session III). For sessions II and III, the water concentrations are not

636 reported in the table because the data were obtained at low primary beam intensity and background
637 correction could not be done.

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(wt.%)	MgO	Al ₂ O ₃	SiO ₂	K ₂ O	CaO	TiO ₂	Cr ₂ O ₃	MnO	FeO	NiO	Total	Fa	Fs	Wo
Chondrule 27														
Olivine	54	0	41.4	0	0.2	0.02	0.44	0.15	1.11	0.04	97.6	1.2		
Olivine	54.2	0	40.8	0	0.22	0.02	0.55	0.2	1.01	0.06	97.2	1.1		
Pyroxene	31.3	1.2	45.3	0.04	0.78	0.22	1.8	0.17	11.87	1.32	100.9		17.5	0.4
Pyroxene	34.2	0.8	51.9	0	0.8	0.27	0.61	0.2	2.53	0.08	91.9		3.9	1.6
Chondrule 19														
Olivine	54.3	0.2	42.3	0	0.29	0.03	0.13	0.08	0.8	0.02	98.3	0.8		
Pyroxene	37.5	1.2	57	0	0.47	0.19	0.4	0.08	0.97	0.09	97.9		1.4	0.9
Pyroxene	47.9	0.5	46.8	0.01	0.24	0.1	0.37	0.14	1.56	0.03	97.9		1.8	0.4
Chondrule 4														
Pyroxene	36.3	0.6	55.7	0.01	0.62	0.08	0.85	0.26	1.77	0	96.4		2.6	1.2
Pyroxene	33.8	0.7	53.4	0.02	0.71	0.11	1.59	0.27	5.33	0.32	97.3		8.0	1.4
Chondrule 24														
Olivine	54.6	0.2	41.1	0.01	0.58	0.1	0.34	0.09	0.87	0	98	0.9		
Olivine	54.6	0.2	40.8	0	0.5	0.07	0.24	0	0.89	0	97.4	0.9		
Pyroxene	35.9	2.2	56.5	0.02	1.78	0.3	0.55	0	0.44	0	97.8		0.7	3.4
Pyroxene	34.1	2.9	56.6	0	2.99	0.3	0.56	0.07	0.63	0.02	98.2		1.0	5.9
Pyroxene	21.7	2	52.7	0	18.86	1.01	0.53	0.14	0.36	0.06	97.5		0.6	38.3
Chondrule 6														
Olivine	52.1	0.1	40.6	0	0.47	0.05	0.43	0.17	1.38	0.05	95.8	1.5		
Olivine	54.6	0	41.2	0.02	0.19	0	0.53	0.13	1.46	0.12	98.3	1.5	3.6	1.0
Pyroxene	37.6	1.3	56.8	0.02	0.54	0.27	0.4	0	0.88	0.05	98.1		1.3	1.0
Pyroxene	38.2	1.3	53.6	0	0.55	0.2	0.5	0.15	2.6	0.17	97.5			
Chondrule 9														
Olivine	54.4	0.1	40.9	0	0.31	0.11	0.46	0.14	1.2	0.03	97.8	1.2		
Olivine	54.5	0.1	41.2	0	0.31	0.01	0.48	0.12	1.61	0.07	98.4	1.6		
Pyroxene	32.5	2.1	50.7	0.03	0.97	0.25	0.81	0.25	6.84	0.04	95.3		10.4	1.9
Pyroxene	35.1	2	54.1	0.01	0.99	0.14	0.71	0.11	3.2	0.01	96.6		4.8	1.9
Pyroxene	30	0.5	47.7	0.01	1.2	0.13	1.61	0.37	4.13	0.04	85.9		7.0	2.6

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Table 1

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	H ₂ O (ppm)	δD (‰)	SiO ₂ (wt.%)	Al ₂ O ₃ (wt.%)	FeO (wt.%)	MnO (wt.%)	MgO (wt.%)	CaO (wt.%)	Na ₂ O (wt.%)	K ₂ O (wt.%)	Cr ₂ O ₃ (wt.%)	TiO ₂ (wt.%)	P ₂ O ₅ (wt.%)	Cl (wt.%)	Total (%)
Olivine ^(a)															
KLV-23	120	-	40.32	0.05	12.27	-	47.47	0.08	0.03	-	-	0.03	-	-	100.26
Pyroxene ^(a)															
KBH-1	186	-	54.68	4.73	5.88	-	32.92	0.86	0.12	-	-	0.11	-	-	99.32
Glasses ^(b)															
DR15-2-5	2581	-74	50.81	14.78	11.15	0.20	6.95	10.37	3.10	0.46	-	1.82	0.21	0.02	99.85
DR20-1-1	3988	-61	50.20	15.13	10.95	0.20	6.89	10.82	3.07	0.28	-	1.92	0.28	0.03	99.74
Amphiboles ^(c)															
Mt Emma	14300	-50	40.87	13.96	9.84	0.08	14.32	10.79	2.56	1.55	0.09	2.89	-	0.04	98.51
Kipawa	14500	-88	45.05	8.32	12.77	0.24	14.01	10.90	2.48	1.05	0.05	0.80	-	0.59	98.49

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643 Table 2

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Chondre	Phase	δD (‰)	H ₂ O (ppm)
Renazzo			
Session IV			
Chondrule X	mesostasis	-73 ± 146	-
	mesostasis	-81 ± 150	2284 ± 275
	olivine	-139 ± 132	115 ± 58
	mesostasis	15 ± 156	-
	pyroxene	-116 ± 137	366 ± 183
Chondrule Y	olivine	-188 ± 129	182 ± 91
	olivine	-115 ± 150	196 ± 98
	mesostasis	-212 ± 125	-
	mesostasis	-211 ± 121	32373 ± 3899
	pyroxene	-109 ± 142	1481 ± 740
	pyroxene	-120 ± 140	1407 ± 703
Chondrule Z	olivine	-192 ± 126	158 ± 79
	pyroxene	-161 ± 132	320 ± 160
	olivine	-120 ± 128	133 ± 67
	olivine	-125 ± 137	215 ± 108
	olivine	-148 ± 132	94 ± 47
	surrounding matrix	316 ± 202	-
Bishunpur			
Session IV			
Chondrule I	mesostasis	-166 ± 133	4702 ± 566
	olivine	-30 ± 151	121 ± 61
	olivine	-164 ± 137	159 ± 79
	mesostasis	-53 ± 146	4725 ± 569
	olivine	-149 ± 135	399 ± 199
	olivine	0 ± 156	159 ± 79
	pyroxene	52 ± 163	659 ± 330
	pyroxene	137 ± 176	1244 ± 622
	surrounding matrix	-113 ± 136	-
Chondrule J	pyroxene	-60 ± 145	1328 ± 664
	pyroxene	-118 ± 138	1796 ± 898
	pyroxene	-149 ± 131	1395 ± 698
	pyroxene	-81 ± 143	674 ± 337
	olivine	-115 ± 136	76 ± 38
	pyroxene	-95 ± 152	649 ± 324
	surrounding matrix	-119 ± 136	-
Chondrule K	olivine	-36 ± 148	-

	pyroxene	-18 ± 153	-
	mesostasis	-46 ± 148	-
Chondrule L	mesostasis	-122 ± 138	5641 ± 679
	pyroxene	-44 ± 149	1090 ± 545
	mesostasis	-49 ± 148	15107 ± 1819
	pyroxene	20 ± 159	567 ± 284
	olivine		316 ± 158
	olivine		326 ± 163
	pyroxene		0
	pyroxene		266 ± 131

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Chondre	Phase	δD (‰)	H ₂ O (ppm)
Session I			
Chondrule 4	mesostasis	-129 ± 31	-
	pyroxene	-121 ± 61	-
	mesostasis	-90 ± 31	-
	pyroxene	-86 ± 31	-
	mesostasis	-104 ± 32	-
	pyroxene	-89 ± 35	-
	surrounding matrix	-94 ± 100	-
Chondrule24	mesostasis	-99 ± 36	-
	pyroxene	-80 ± 43	-
	olivine	19 ± 54	-
	mesostasis	-252 ± 38	-
	olivine	-125 ± 50	-
	pyroxene	-12 ± 44	-
Chondrule 6	mesostasis	-97 ± 30	-
	olivine	-130 ± 35	-
	olivine	18 ± 45	-
	pyroxene	-34 ± 35	-
	surrounding matrix	-27 ± 39	-
Chondrule19	mesostasis	-23 ± 45	-
	olivine	-21 ± 38	-
	olivine	0 ± 46	-
	pyroxene	-91 ± 37	-
Chondrule27	mesostasis	-182 ± 28	-
	olivine	-24 ± 43	-

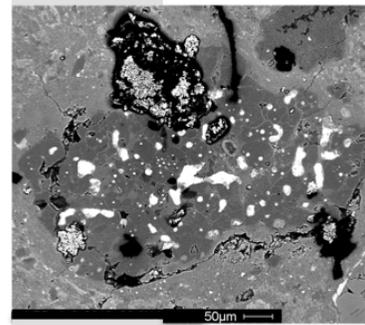
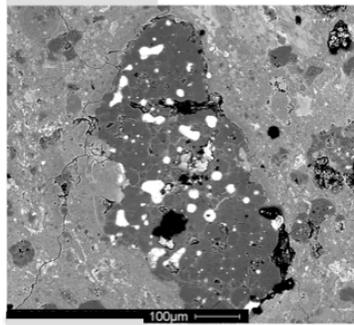
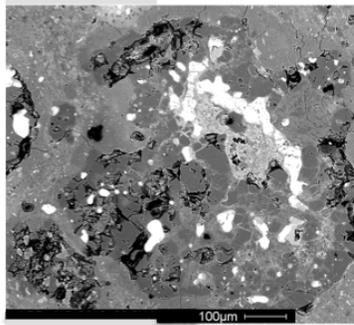
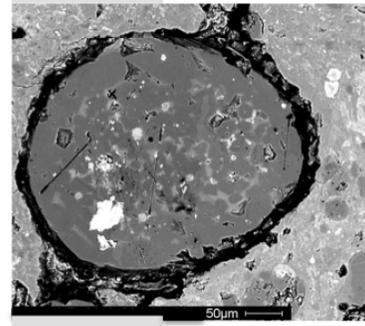
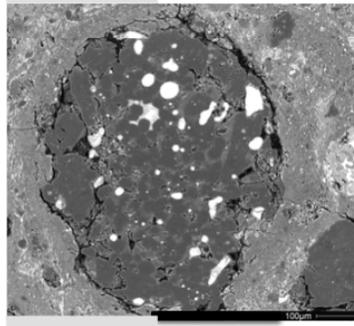
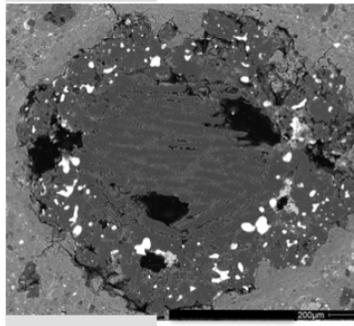
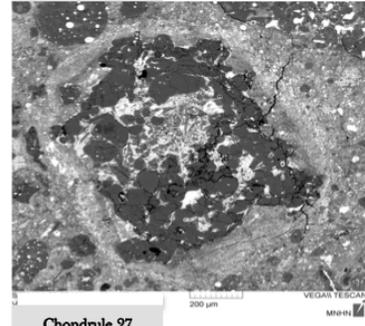
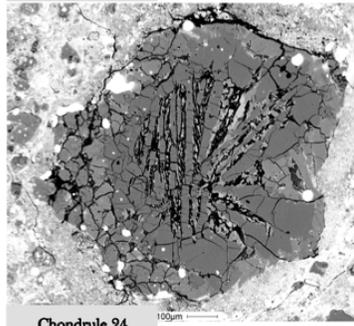
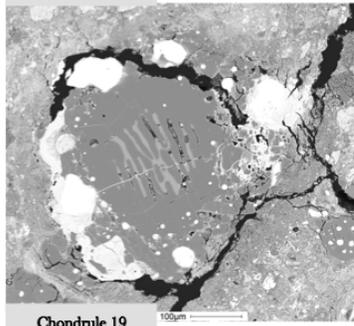
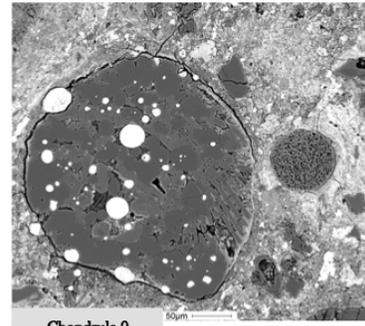
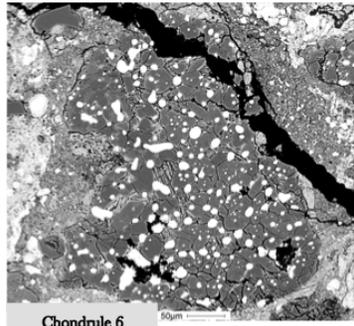
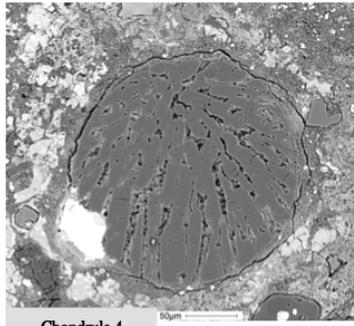
	mesostasis	-212 ± 27	-
	pyroxene	20 ± 27	-
Chondrule 9	pyroxene	-41 ± 36	-
	mesostasis	64 ± 40	-
	pyroxene	22 ± 41	-
	pyroxene	155 ± 53	-
	surrounding matrix	-77 ± 75	-
<hr/>			
Paris			
Session II			
Chondrule 1	olivine	-93 ± 28	-
	mesostasis	140 ± 36	-
	mesostasis	115 ± 36	-
	olivine	-66 ± 48	-
	pyroxene	-62 ± 32	-
	mesostasis	37 ± 30	-
	olivine	-213 ± 31	-
	pyroxene	21 ± 45	-
Chondrule 7	mesostasis	-96 ± 26	-
	olivine	-116 ± 30	-
	mesostasis	-132 ± 24	-
	olivine	-36 ± 30	-
	pyroxene	-113 ± 50	-
<hr/>			
Paris			
Session III			
Chondrule A	olivine	-374 ± 30	603 ± 145
	olivine	-279 ± 29	758 ± 182
	olivine	-334 ± 31	961 ± 231
	olivine	-179 ± 45	949 ± 228
	mesostasis	-17 ± 25	30720 ± 1516
	mesostasis	-274 ± 18	6175 ± 305
	olivine	-398 ± 23	961 ± 231
	mesostasis	266 ± 119	6470 ± 319
	olivine	76 ± 36	781 ± 188
	pyroxene	-179 ± 25	503 ± 121
	pyroxene	-342 ± 23	985 ± 237
	mesostasis	-175 ± 22	5630 ± 278
Chondrule B	mesostasis		66419 ± 3277
	pyroxene		860 ± 207
	olivine		844 ± 203
	pyroxene		900 ± 216
	olivine		859 ± 206

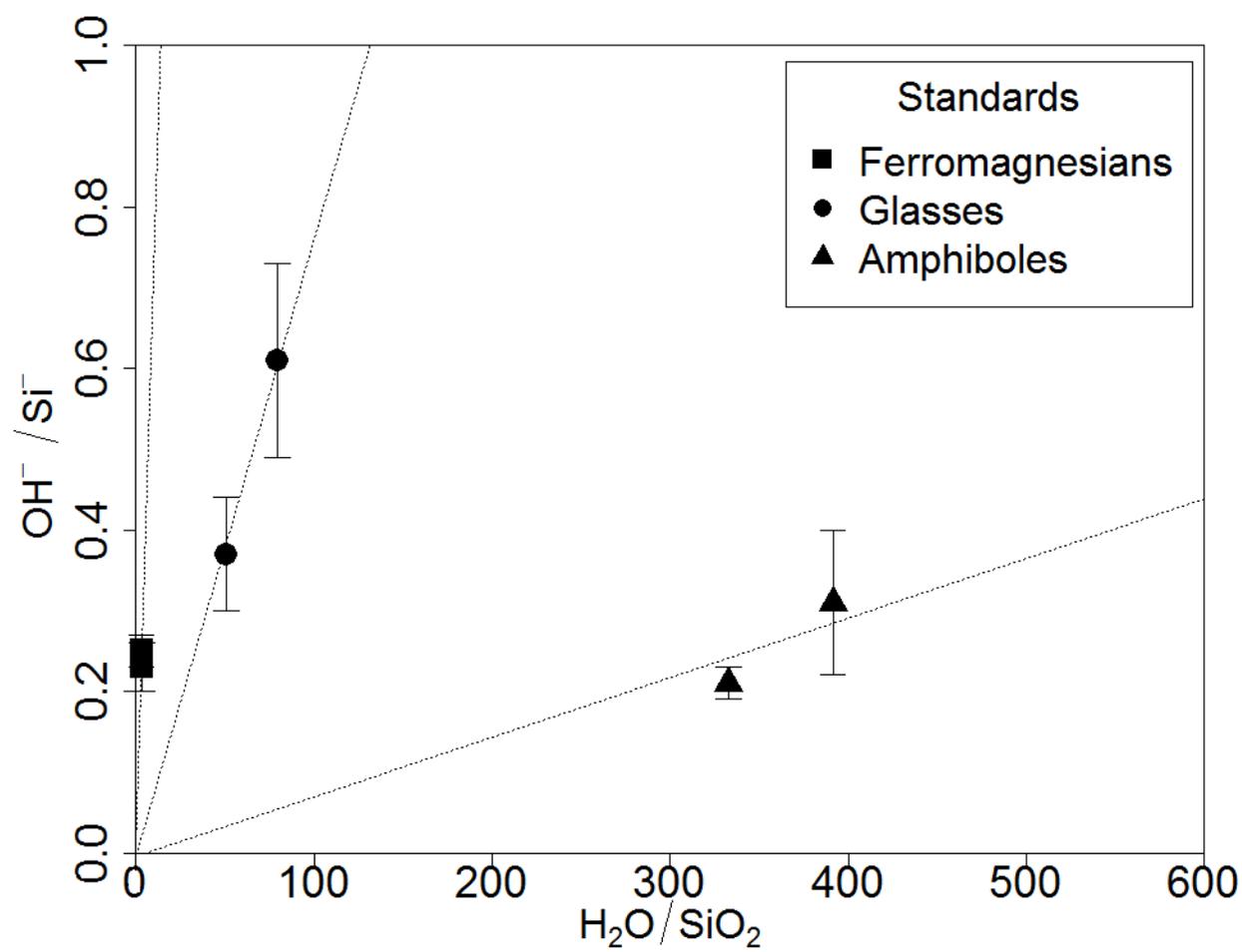
	olivine		998 ± 240
	mesostasis	101 ± 34	33462 ± 1651
	pyroxene	-225 ± 21	698 ± 168
	pyroxene	-172 ± 22	1174 ± 282
Chondrule D	mesostasis	-102 ± 29	8931 ± 441
	olivine	-225 ± 21	799 ± 182
	mesostasis	-270 ± 21	7940 ± 392
	pyroxene	-337 ± 20	595 ± 143
Chondrule E	mesostasis	191 ± 29	87466 ± 4317
	pyroxene	83 ± 27	615 ± 148
	olivine	-172 ± 31	
	mesostasis	112 ± 29	10934 ± 539
	olivine	13 ± 29	1051 ± 253
	mesostasis	366 ± 35	31575 ± 1558
	olivine	-138 ± 27	933 ± 224
	mesostasis	-292 ± 32	15926 ± 786
	olivine	-208 ± 23	905 ± 218
	olivine	-225 ± 29	639 ± 154

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659 Table 3

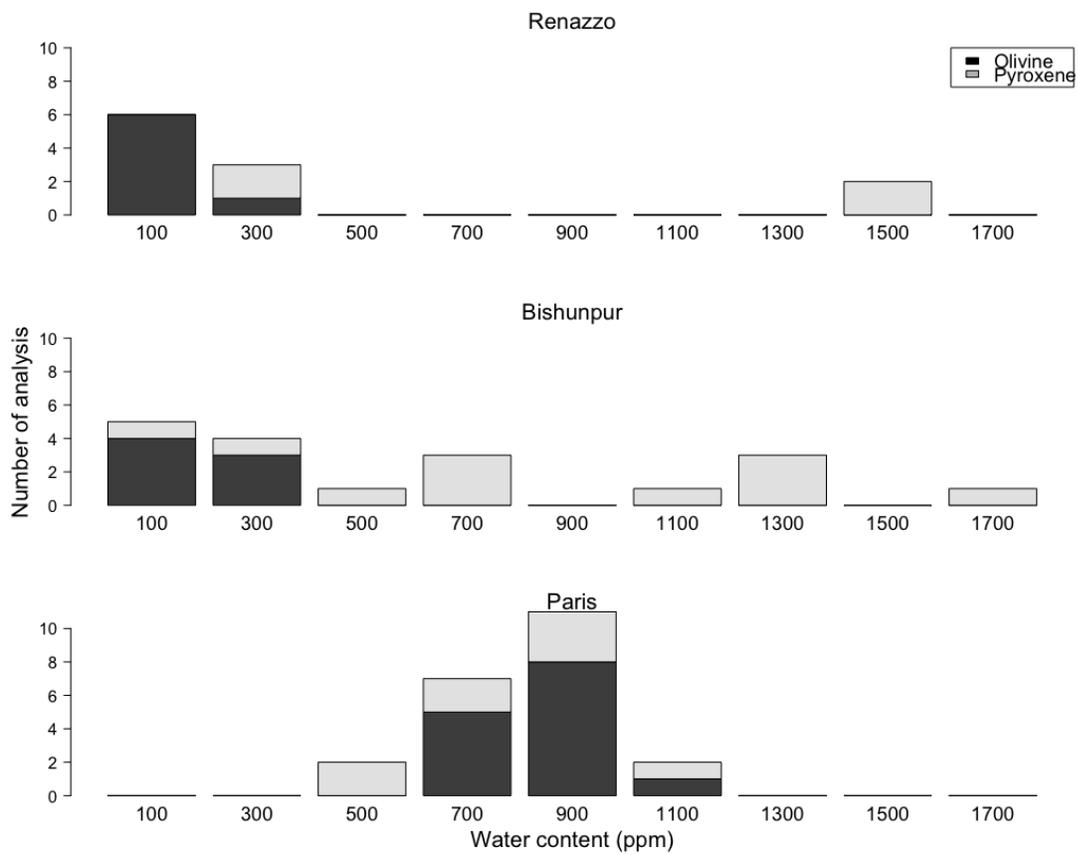
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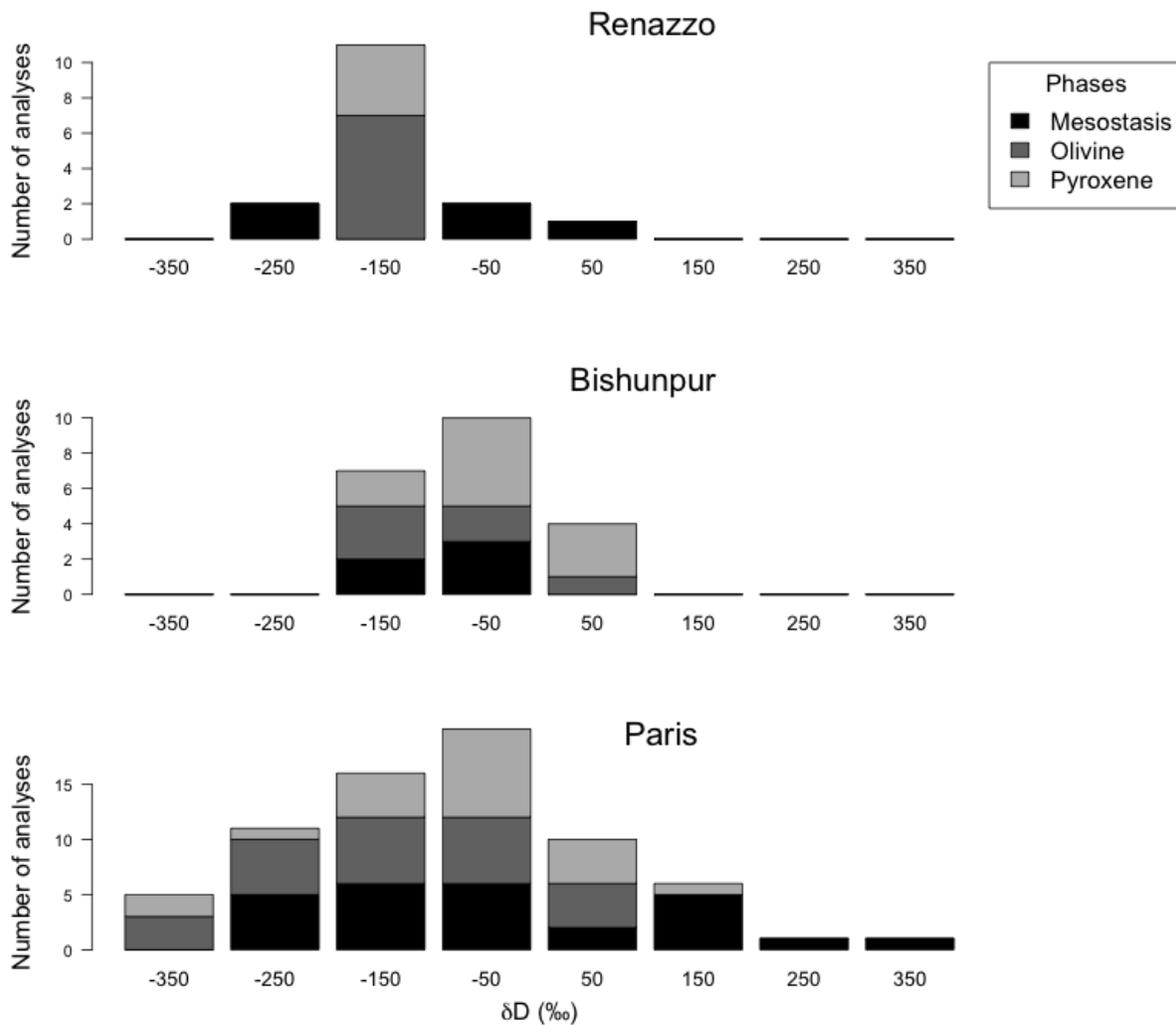
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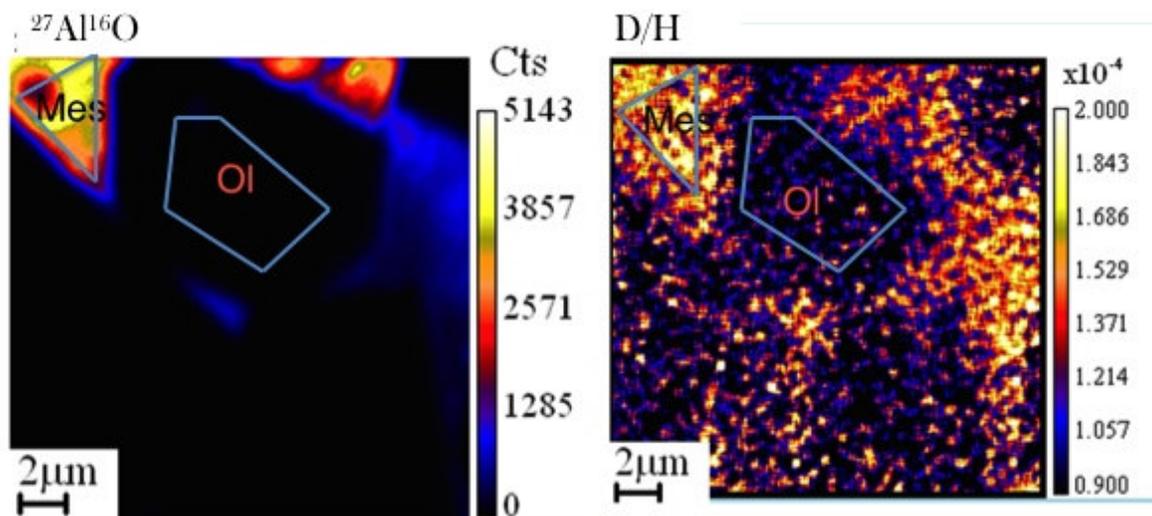
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