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1 **Non-uniform splitting of a single mantle plume by double cratonic roots:**

2 **Insight into the origin of the central and southern East African Rift System**

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15

16 **Abstract**

17 Using numerical thermo-mechanical experiments we analyze the role of active mantle
18 plume and pre-existing lithospheric thickness differences in the structural development of the
19 central and southern East African Rift system. The plume-lithosphere interaction model setup
20 captures the essential features of the studied area: two cratonic bodies embedded into surrounding
21 lithosphere of normal thickness. Results of numerical experiments suggest that localization of rift
22 branches in the crust is mainly defined by the initial position of the mantle plume relative to the
23 cratons. We demonstrate that development of the Eastern branch, the Western branch and the
24 Malawi rift can be the result of non-uniform splitting of the Kenyan plume that has been rising
25 underneath the southern part of the Tanzanian craton. Major features associated with Cenozoic
26 rifting can thus be reproduced in a relatively simple context of the interaction between a single
27 mantle plume and pre-stressed continental lithosphere with double cratonic roots.

28

29 **Introduction**

30 The East African Rift system (EARS) (McConnell, 1972; Chorowicz, 2005; Braile et al.,
31 2006; Ring, 2014) has two branches, the Eastern branch and Western branch. The N-S oriented
32 Eastern branch (Baker et al., 1972; Williams, 1982; Baker, 1987; Keller et al., 1991; Smith, 1994;
33 Mechie et al., 1997; Ebinger, 2005) extends over 2000 km from the Afar Triple Junction (Mohr,
34 1970; McClusky et al., 2003) in the north to the North Tanzanian Divergence Zone (Dawson,
35 1992; Le Gall et al., 2004, 2008; Isola et al., 2014) in the south (Figure 1) and consists of the
36 relatively narrow Main Ethiopian rift (Keranen et al., 2009), the wide rift in the Turkana
37 depression (Morley et al., 1992) and a narrow rift at the Tanzanian craton margin in the Kenya
38 rift (Zeyen et al., 1997). In the present study, however, we consider only the southern part of the
39 Eastern branch from Northern Kenya to North-Eastern Tanzania. The Western branch (Ebinger,

40 1989; Pasteels et al., 1989; Daly et al., 1989; Morley et al., 1999; Bauer et al., 2013) is composed
41 by the Albert-Edward, Kivu and Tanganyika-Rukwa rifts stretched in NE-SW, N-S and NW-SE
42 directions, respectively, depicting an arcuate map-trace along the western side of the Tanzanian
43 craton (Figure 1). The southern prolongation of the Western branch is represented by the Malawi
44 rift (Ring et al., 1992; Laó-Dávila et al., 2015) that is aligned on a N-S trend extending from the
45 Rungwe volcanic province (southern Tanzania) to the Urema graben (Mozambique). Geological
46 estimates indicate a higher degree of total lithospheric extension in the Kenya rift as compared to
47 the Western branch and the North Tanzanian Divergence Zone (Ring, 2014 and references
48 therein).

49 The Western and Eastern branches are separated by the Archaean (2500-3000 Ma)
50 Tanzanian craton (Bell and Dodson, 1981; Chesley et al., 1999; Manya, 2011) characterized by a
51 strong and cold lithosphere with a 150-300-km-thick keel (Artemieva, 2006; Adams et al., 2012;
52 Mulibo and Nyblade, 2013a,b). The 175-km-thick (Artemieva, 2006) Archaean-Paleoproterozoic
53 Bangweulu block (Andersen and Unrug, 1984; De Waele et al., 2006) that is stable since 1750
54 Ma (Lenoir et al., 1994), lies to the south-west of the Tanzanian craton. The Tanzanian and
55 Bangweulu cratons are both surrounded by Proterozoic orogenic belts (Cahen et al., 1984; Begg
56 et al., 2009) with a relatively thin (≤ 150 km) thermal lithosphere (Artemieva, 2006; Koptev and
57 Ershov, 2011).

58 Geophysical (e.g., Ritsema et al., 1998; Nyblade et al., 2000; Nolet et al., 2006) and
59 geochemical (e.g., Rooney et al., 2012; Armitage et al., 2015) observations indicate the presence
60 of mantle plumes under the EARS, possibly rooted into a common deep mantle anomaly
61 (Ritsema et al., 1999) corresponding to the African superplume. However, the actual number of
62 plumes and their relative position within this broad upwelling region remain contentious (e.g.,
63 Weeraratne et al., 2003; Chang and Van der Lee, 2011; Chang et al., 2015; Civiero et al., 2016).

64 It is commonly assumed that the Cenozoic rifts have avoided the cratons and follow the
65 mobile belts (McConnell, 1972; Mohr, 1982) which serve as the weakest pathways for rift
66 propagation. Structural control of the pre-existing heterogeneities within the Proterozoic belts at
67 the scale of individual faults or rifts has been demonstrated as well (Versfelt and Rosendahl,
68 1989; Ring, 1994; Theunissen et al., 1996; Corti et al., 2007; Morley, 2010; Katumwehe et al.,
69 2015; Smets et al., 2016).

70 However, as shown by Koptev et al. (2015, 2016), the formation of two rift zones on
71 opposite sides of a thick lithosphere segment can be explained without appealing to pre-imposed
72 heterogeneities at the crustal level. These models have provided a unified physical framework to
73 understand simultaneous development of the Western and Eastern branches around a thicker
74 Tanzanian craton (Roberts et al., 2012) as a result of the interaction between pre-stressed
75 continental lithosphere and single mantle plume anomaly corresponding to the Kenyan plume
76 (George et al., 1998; Pik et al., 2006; Chang and Van der Lee, 2011). Yet, the southern
77 prolongation of the Western rift by the Malawi rift has not been reproduced in any of these “one-
78 craton” experiments (Koptev et al., 2015, 2016). In order to overcome this discrepancy, we
79 follow-up on our previous studies with a series of laterally widened thermo-mechanical models
80 characterized by a presence of a second zone of lithospheric thickening that roughly mimics the
81 isometric (i.e. having equal horizontal dimensions) Bangweulu block situated south-west of the
82 Tanzanian craton and by a single mantle plume seeded underneath the southern part of the
83 Tanzanian craton (e.g., Hansen et al., 2012; Bagley and Nyblade, 2013).

84

85 **Model and experiments**

86 Our modeling is based on the thermo-mechanical viscous-plastic 3DELVIS code (Gerya
87 and Yuen, 2007) that combines the finite difference method with a marker-in-cell technique (see
88 Gerya (2010) for more details).

89 The 3D model box encompasses the entire 650-km-deep upper mantle, with large
90 horizontal scales (2000×2000 km) and offers “lithospheric-scale” resolution (~ 5 km \times 5 km \times 5
91 km per grid cell). The initial model setup corresponds to onset of a stratified three-layer
92 (upper/lower crust and lithospheric mantle) continental lithosphere which is underlain by
93 asthenospheric upper mantle. The initial geotherm is piece-wise linear, with 0°C at the surface,
94 700°C at the Moho, 1300°C at the lithosphere-asthenosphere boundary and 1630°C at the bottom
95 of the model box. The mantle plume is initiated by a temperature anomaly of $+370^\circ\text{C}$ at the base
96 of the upper mantle. In all presented experiments we have applied a constant extension in E-W
97 direction with a half rate of 3 mm/year, a typical value for pre-break-up continental rifts
98 including the Nubia-Somalia plate system (Stamps et al., 2008; Saria et al., 2014).

99 The general model series (Table 1; models **1-13**) is characterized by lateral homogeneity
100 of the crustal composition and by presence of two zones of 250 km-thick cratons embedded into
101 surrounding “normal” (150 km-thick) lithosphere. The first cratonic block is elongated in a N-S
102 direction (horizontal dimensions are 400×800 km) roughly mimicking the configuration of the
103 Tanzanian craton whereas the second one is small and isometric (horizontal dimensions are $400 \times$
104 400 km) corresponding to the Bangweulu block. The initial plume location represents a key
105 controlling parameter of our study (Figure 2). In all performed experiments (except for “no-
106 plume” model **1**; Figure 3) the mantle plume has been shifted to the south up to 230 km with
107 respect to the center of the Tanzanian craton. We have studied the impact of small lateral
108 variations in its initial position: a southward shift varies from 210 km to 230 km whereas
109 latitudinal displacement is from 10 km to the west to 30 km to the east; different configurations

110 of the Tanzanian craton have been tested as well (Table 1; Figures 4-5). In order to investigate
111 the potential role of second-order structural heterogeneities we have performed several
112 complementary experiments (models **14-17**, Figure 6) including stronger (plagioclase flow law
113 instead of wet quartzite flow law) lower crust within the cratonic blocks (models **16-17**) and/or
114 third zone of lithospheric thickening situated to the west of the Tanzanian craton and roughly
115 mimicking the size and configuration of the Masai block (models **14-15, 17**).

116

117 **Results and discussions**

118 Similarly to previous 3D experiments (Burov and Gerya, 2014; Koptev et al., 2015,
119 2016), the models presented here predict a rapid mantle ascent as the mantle plume reaches the
120 lithospheric bottom after 0.5 Myr. The common point of the performed models is a separation of
121 the upwelling plume head into three parts by the lithosphere of the Tanzanian and Bangweulu
122 blocks. After being divided, the buoyant plume material ponds at the base of “normal”
123 lithosphere adjacent to the western, eastern and southern sides of the Tanzanian craton (Figure 4).

124 In absence of active mantle upwelling (“no-plume” model **1**), ultra-slow tectonic
125 extension can only result in broadly distributed small-offset parallel faults (Figure 3). On the
126 contrary, in most of the other experiments (Figure 4), the continental crust above the hot plume
127 material is subjected to localized brittle deformation forming three linear, 100-500 km-long
128 rifting centers stretched in the direction perpendicular to external E-W extension. As already
129 shown by Burov and Gerya (2014), such large-scale linear normal faults are triggered and
130 maintained by mantle flow that impacts the bottom of continental lithosphere.

131 The degree of development (in terms of modeled strain rates) of each of these branches is
132 directly controlled by the relative amount of hot plume material ponding underneath the
133 corresponding lithosphere segment (Figure 4). In certain cases, however, this amount appears to

134 be too small to localize any visible deformation in the crust. For example, the plume's eastward
135 shift of 10 km in combination with the simplified shape of the Tanzanian craton results in the
136 absence of a discernible western rift (model **4**, Figure 4c). Similarly, the eastern branch is not
137 reproduced in the experiments assuming a more realistic Tanzanian craton and the mantle plume
138 with initial latitudinal displacement of 0 km or 10 km to the west (models **6** and **5**, respectively;
139 Figure 4e and Figure 4d). Only the plume's eastward shift up to 20 km (models **8-10**; Figure 4g-i)
140 can provide a symmetrical rifting on both (eastern and western) sides of the Tanzanian craton.
141 Further plume displacement to the east (up to 30 km) leads to a more developed eastern branch
142 comparing to the western one (models **11-13**; Figure 4j-l). The southern rift zone situated east of
143 the Bangweulu block is expressed clearly in all models. Note, however, that it becomes less
144 pronounced when the plume's southward shift decreases from 230 km to 210 km (compare
145 models **8** (Figure 4g) and **10** (Figure 4i) or models **11** (Figure 4j) and **13** (Figure 4l)). The
146 asymmetrical distribution of hot mantle material on the opposite sides of the craton (e.g., models
147 **11-12**) can provoke contrasted magmatic activity as observed in the central EARS, explaining a
148 magma-rich Eastern branch and magma-poor Western branch (e.g., Chorowicz et al., 2005; Ring,
149 2014).

150 We have identified model **9** (Figure 4h; Figure 5) as the "best-fit" experiment of the
151 general model series (models **1-13**) because it reproduces all three rift zones that are clearly
152 reflected not only in strain rate field but also in the surface topography (Figure 5c). Several other
153 models (such as models **3** or **8**), however, could also be considered as the "best-fit" (Figure 4b,g)
154 because they are able to reproduce a synchronous growth of the Eastern branch, the Western
155 branch and the Malawi rift as well. Note that modeled accumulated deformation in the localized
156 rift basins amounts to several tens km of total extension, that is in accord with geological
157 estimates for the central part of the EARS (e.g., Ring, 2014 and references therein). A striking

158 feature is the consistency between modeled and observed rift distribution in case of the Kenyan
159 plume seeded underneath the southern part of the Tanzanian craton.

160 A principal shortcoming of this “best-fit” experiment (model **9**; Figure 5) is that it does
161 not reproduce the NW-SE oriented Rukwa rift segment of the Western branch (Figure 1).
162 Similarly to most of the other models, the southern end of the western branch penetrates into the
163 area underlain by the Bangweulu block. In this case, thick lithosphere appears to be not resistant
164 to lateral propagation of localized deformation given homogenous crustal composition adopted in
165 the general model series when crustal rheology of the cratonic blocks does not differ from
166 surrounding “normal” lithosphere. The situation changes completely when not only thicker
167 lithospheric mantle but also stronger rheology of the lower crust is considered for the cratonic
168 areas (complementary model series: models **16-17**, Figure 6c,d): in this case, localized strain
169 tends to avoid the strong cratons keeping them almost undeformed. In particular, this leads to the
170 change of orientation from N-S to NW-SE for the southernmost segment of the western branch
171 (corresponding to the Rukwa rift) thus providing a framework much more consistent with the
172 observed data. Introduction of the Masai block, for its part, improves the relative location of
173 widening of the rift zone in the southern segment of the eastern branch (models **15, 17**; Figure
174 6b,d) that would correspond to the observed transition between the narrow Kenya rift to a
175 considerably wider zone (300-400 km) of block faulting in northern Tanzania (Dawson, 1992;
176 Ebinger et al., 1997; Foster et al., 1997; Le Gall, 2008; Corti et al., 2013).

177 It should be noted that given the better fit with the observed data in case of
178 complementary models containing more predefined complexities (see Figure 6d), the analyzed
179 plume-induced multi-branch continental rifting dynamics need further numerical investigation
180 and quantitative analysis of the controlling factors including complex 3D structure of the central

181 EARS in terms of inherited compositional and rheological heterogeneities of the crust and
182 lithospheric mantle (e.g., Sippel et al., 2017).

183

184 **Conclusions**

185 The fully-coupled thermo-mechanical models presented here start from relatively simple
186 initial conditions: a single mantle plume anomaly seeded underneath the continental lithosphere
187 embedding two cratonic blocks. These experiments evolve over time to create a complex system
188 characterized by asymmetrical splitting of the plume head into three parts. The resulting relative
189 distribution of hot plume material ponding below “normal” lithosphere is a controlling parameter
190 for deformation localization at the crustal level. Very small variations in initial plume position
191 with respect to cratonic bodies (up to several tens km only) appear to be able to change the
192 relation between these three segments of separated plume head that, in turn, alters the degree of
193 development of corresponding rift branches. We argue, thus, that the resulting rifting pattern is
194 largely controlled by relative position of the initial mantle plume anomaly with respect to first-
195 order lithospheric thickness differences rather than by second-order crustal and/or lithospheric
196 compositional heterogeneities as commonly assumed before (Theunissen et al., 1996; Corti et al.,
197 2007; Katumwehe et al., 2015; Smets et al., 2016).

198 The performed analysis permits to identify an initial model configuration that results in a
199 strain distribution that bears strong similarities with the central and southern EARS showing
200 simultaneous development of the Eastern branch, the Western branch and its southern
201 prolongation by the Malawi rift. The number and relative positions of rift branches with respect
202 to the cratons within the studied area can be explained by an impact of single mantle anomaly on
203 pre-stressed continental lithosphere that does not contain any pre-defined heterogeneities other
204 than the well-known cratonic blocks.

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213

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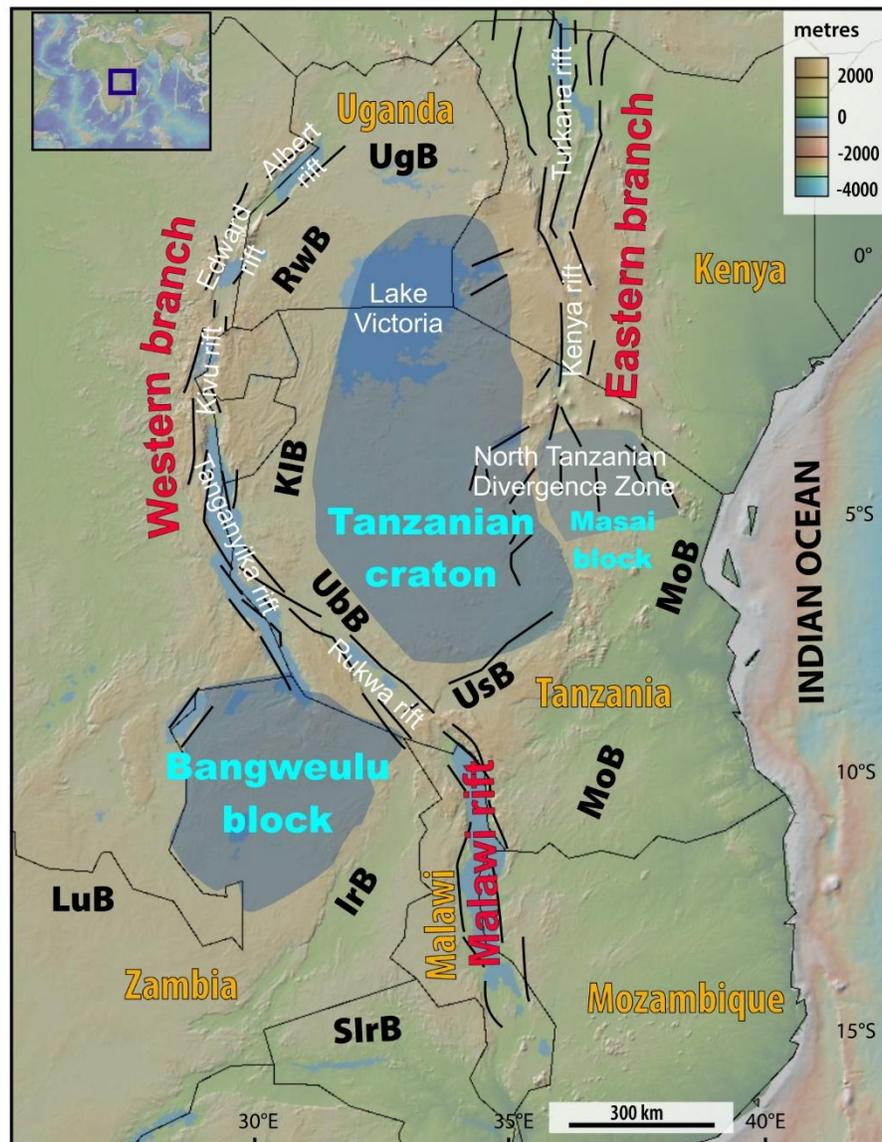
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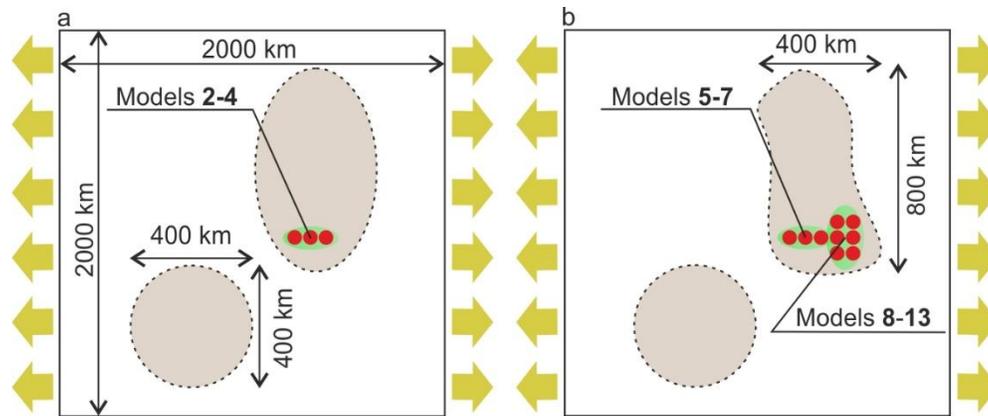
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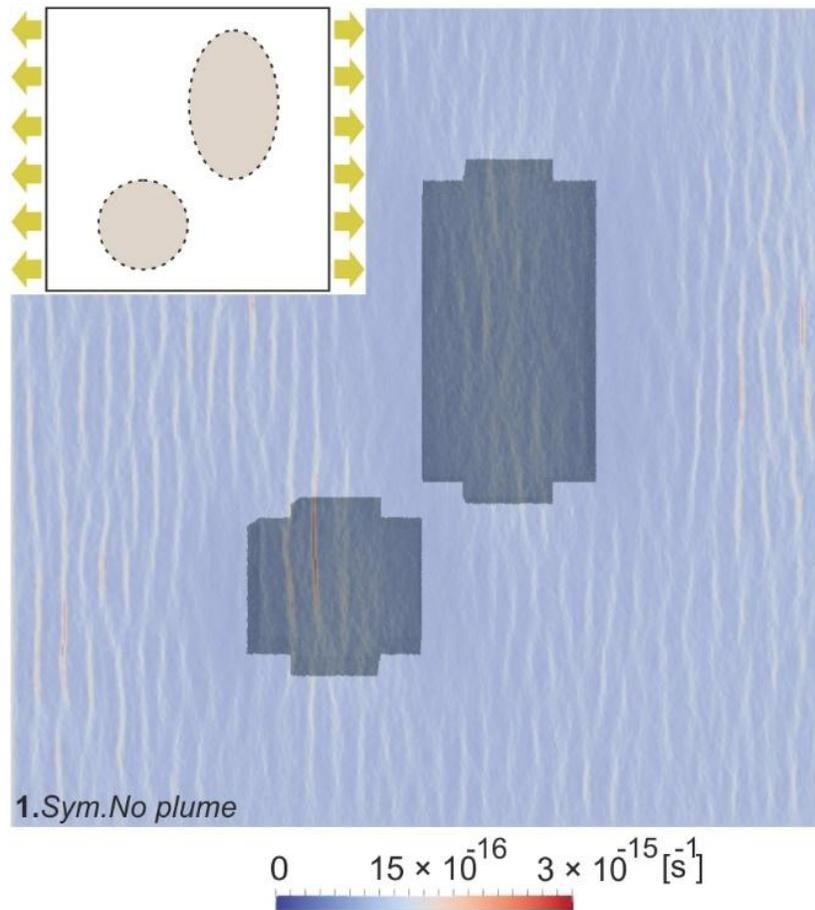
437 **Figure captions**



438
439 **Figure 1.** Topographic map showing the tectonic setting of the central and southern EARS (after
440 Mulibo and Nyblade (2013b) and Corti et al. (2013)) that comprises the Tanzanian craton that
441 likely includes the Uganda Basement Complex (UgB), the Bangweulu block, the Masai block and
442 several Proterozoic orogenic belts: Rwenzori (RwB), Kibaran (KiB), Ubendian (UbB), Usagaran
443 (UsB), Mozambique (MoB), Iruminde (IrB), Southern Iruminde (SIrB), Lufilian (LuB). Black
444 lines show major faults (Corti et al., 2013). The inset indicates the location of the studied area.



445
 446 **Figure 2.** Model setups for the general model series shown on Figure 4: a) the models **2-4**,
 447 characterized by a simple quasi-rectangular shape of the Tanzanian craton as in previously
 448 published experiments (Koptev et al., 2015, 2016); b) the models **5-13**, with a more complex
 449 asymmetrical configuration of the Tanzanian block (based on its present-day surface outline).
 450 Initial plume positions are shown by red circles within tested areas (shaded green) with respect to
 451 cratonic blocks (gray ellipses). “No-plume” model **1** (see Figure 3) is characterized by simple
 452 symmetrical configuration of the Tanzanian craton and by absence of pre-imposed mantle plume
 453 anomaly. The initial model setup and geotherm have been adopted with respect to observation-
 454 based models of regional thermal and rheological structure of the continental lithosphere in East
 455 Africa (Artemieva et al., 2006; Albaric et al., 2009; Pérez-Gussinyé et al., 2009; Fishwick and
 456 Bastow, 2011). Rheological parameters have been chosen in consideration of extensive and
 457 successful experience obtained from heterogeneous continental rifting (e.g., Huisman and
 458 Beaumont, 2007; Wenker and Beaumont, 2016 and references therein) and plume-lithosphere
 459 interaction modelling (e.g., Burov and Guillou-Frottier, 2005; Burov et al., 2007; Burov and
 460 Cloetingh, 2010; Burov, 2011; Burov and Gerya, 2014; Koptev et al., 2017a,b; Beniest et al.,
 461 2017a,b) including our previous Africa-oriented experiments (Koptev et al., 2015, 2016) that
 462 have permitted to reproduce a number of key features of the central EARS as timing, surface
 463 velocity distribution, and large-scale topography.



464

465 **Figure 3.** “No-plume” model **1** at 5 Myr. Blue to red colors indicate crustal strain rate. The

466 cratons are the dark gray volumes. Unidirectional tectonic stretching of the continental

467 lithosphere in the absence of active mantle upwelling results in distributed closely spaced small-

468 offset parallel faults, which are not localized within any particular zone. Upper crustal distributed

469 deformation covers all model domains uniformly (including the cratonic areas) because of lateral

470 homogeneity of the crustal composition adopted in the general model series. Note that

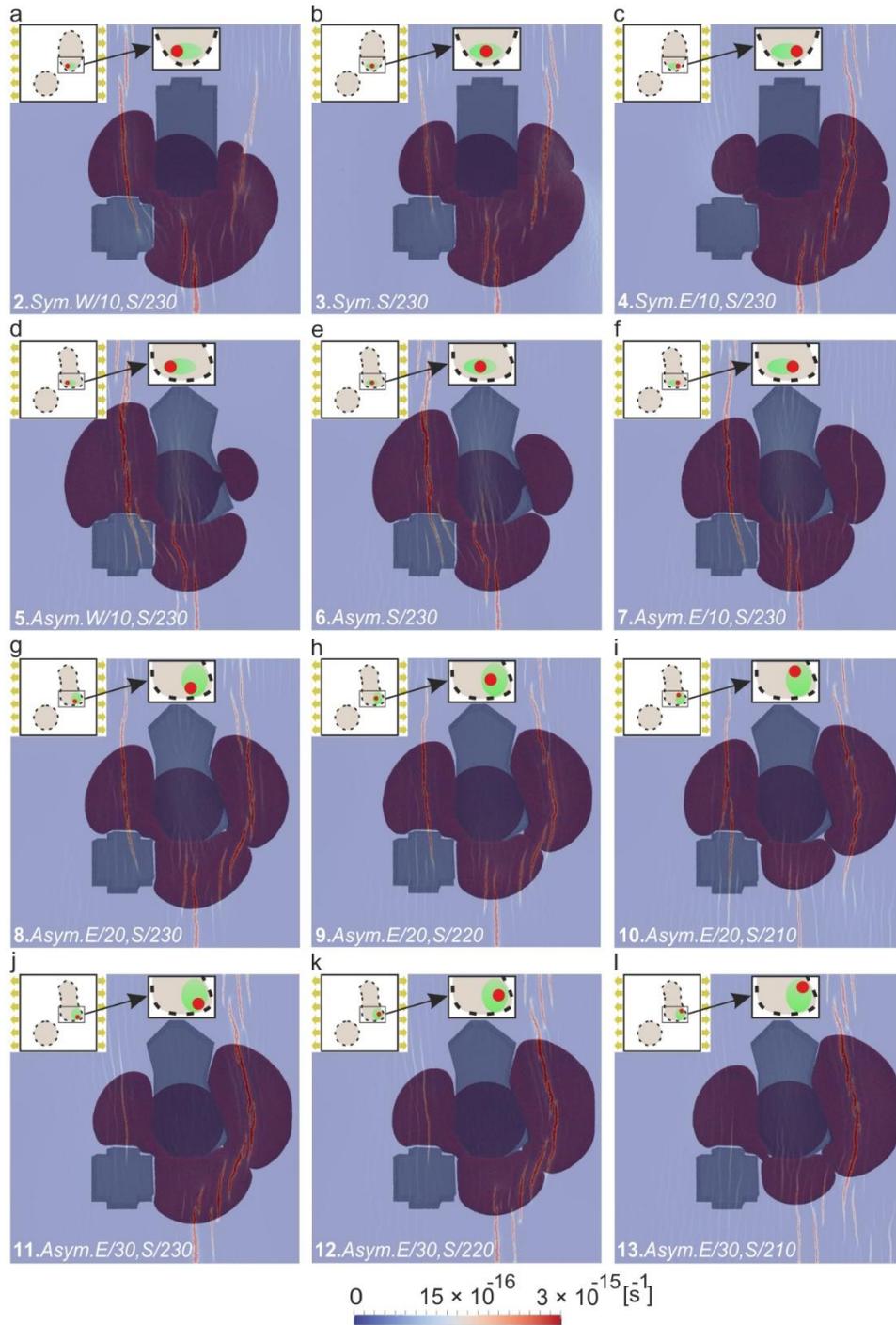
471 progressive focusing and amplification of localized non-axisymmetric deformation is generated

472 only by simultaneous presence of the hot plume material underneath the lithosphere basement

473 and passive horizontal extension while mantle plume impingement on non-pre-stressed

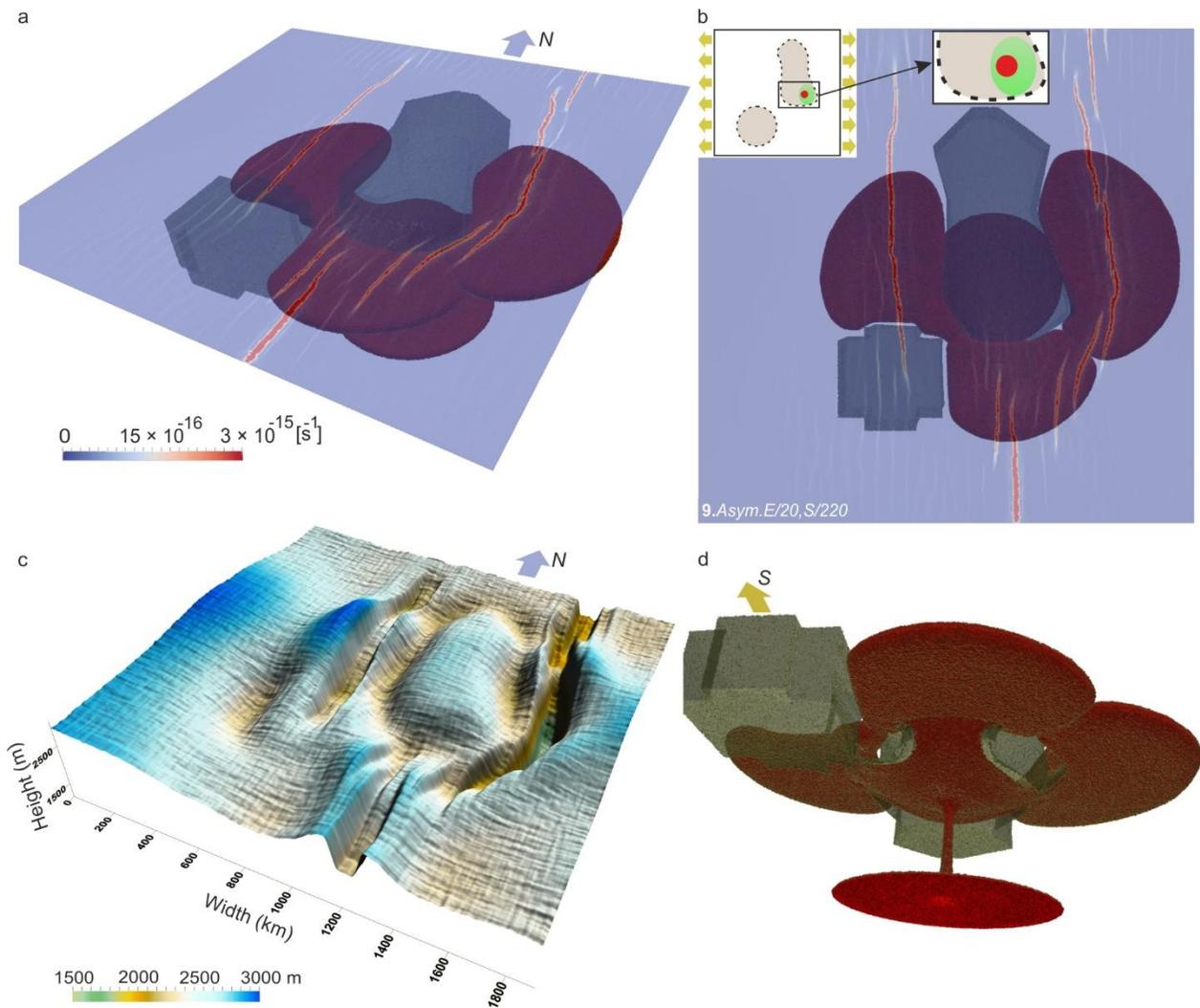
474 lithosphere can only result in axisymmetric domal-shaped features with multiple radiating rifts

475 (see Burov and Gerya, 2014 for more detail).

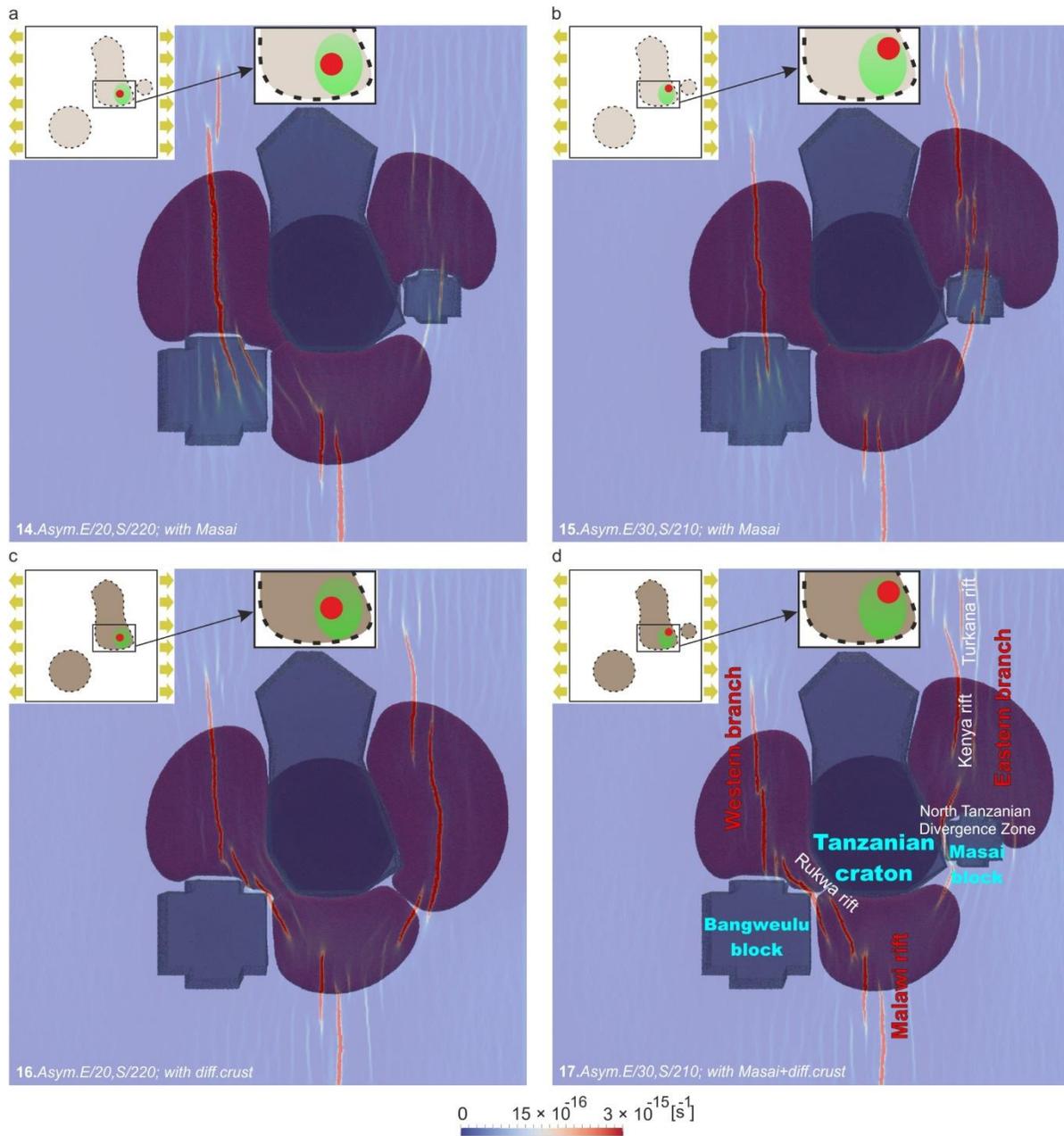


476

477 **Figure 4.** Top view of the results of 3D experiments 2-13 at 5-10 Myr. The plume material is
 478 shown in dark red. Left top insets schematically illustrate initial position of the mantle plume.
 479 Note that relative position of lithospheric heterogeneities and initial mantle plume anomaly
 480 appears to be the crucial factor for resulting rifting pattern.



481
 482 **Figure 5.** The “best-fit” experiment **9** (*Asym.E/20,S/220*) of the general model series: a) 3D view;
 483 b) top view; c) corresponding surface topography; d) bottom view. Note that strain distribution
 484 bears strong similarities with the central and southern EARS showing the “three-branches”
 485 pattern with simultaneous development of the Eastern branch, the Western branch and the
 486 Malawi rift.



487
 488 Figure 6. Top view of the complementary models **14-17** at 5-10 Myr. Complementary models
 489 differ from the general model series (models **1-13**; Figure 4-5) by: a-b, d) the presence of third
 490 area of lithospheric thickening corresponding to the Masai block (models **14-15, 17**) and/or by c-
 491 d) a stronger rheology for the lower crust within cratonic areas (models **16-17**). Note that these
 492 experiments containing more predefined complexities provide better fitting with the observed
 493 data than in case of the general model series. In particular, the model **17** reproduces not only

494 three first-order rift structures corresponding to the Eastern branch, the Western branch and the
495 Malawi rift but also such second-order features as NW-SE oriented Rukwa rift and along-axis
496 transition observed between the narrow Kenya rift and the broader Turkana depression to the
497 north and the multi-basin North Tanzania Divergence Zone to the south (compare Figure 1 and
498 Figure 6d).

499

500 **Table 1.** Controlling parameters of the models.

<i>Model number</i>	<i>Model series</i>	<i>Model title</i>	<i>Controlling parameters</i>				<i>Figure</i>
			<i>Tanzanian craton shape</i>	<i>Plume position (az/shift, km)</i>	<i>Presence of the Masai block</i>	<i>Stronger lower crust for cratons</i>	
1	General	<i>Sym.No plume</i>	Symmetrical	-	No	No	Fig.3
2	General	<i>Sym.W/10,S/230</i>	Symmetrical	W/10, S/230	No	No	Fig4a
3	General	<i>Sym.S/230</i>	Symmetrical	E/0, S/230	No	No	Fig4b
4	General	<i>Sym.E/10,S/230</i>	Symmetrical	E/10, S/230	No	No	Fig4c
5	General	<i>Asym.W/10,S/230</i>	Asymmetrical	W/10, S/230	No	No	Fig4d
6	General	<i>Asym.S/230</i>	Asymmetrical	E/0, S/230	No	No	Fig4e
7	General	<i>Asym.E/10,S/230</i>	Asymmetrical	E/10, S/230	No	No	Fig4f
8	General	<i>Asym.E/20,S/230</i>	Asymmetrical	E/20, S/230	No	No	Fig4g
9	General	<i>Asym.E/20,S/220</i>	Asymmetrical	E/20, S/220	No	No	Fig4h; Fig.5
10	General	<i>Asym.E/20,S/210</i>	Asymmetrical	E/20, S/210	No	No	Fig4i
11	General	<i>Asym.E/30,S/230</i>	Asymmetrical	E/30, S/230	No	No	Fig4j
12	General	<i>Asym.E/30,S/220</i>	Asymmetrical	E/30, S/220	No	No	Fig4k
13	General	<i>Asym.E/30,S/210</i>	Asymmetrical	E/30, S/210	No	No	Fig4l
14	Complementary	<i>Asym.E/20,S/220; with Masai</i>	Asymmetrical	E/20, S/220	Yes	No	Fig6a
15	Complementary	<i>Asym.E/30,S/210; with Masai</i>	Asymmetrical	E/30, S/210	Yes	No	Fig6b
16	Complementary	<i>Asym.E/20,S/220; with diff.crust</i>	Asymmetrical	E/20, S/220	No	Yes	Fig6c
17	Complementary	<i>Asym.E/30,S/210; with Masai+diff.crust</i>	Asymmetrical	E/30, S/210	Yes	Yes	Fig6d