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1 **Response of surface wind divergence to mesoscale SST anomalies under**  
2 **different wind conditions**

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## ABSTRACT

9 The response of the atmospheric boundary layer to mesoscale sea surface  
10 temperature (SST) is often characterized by a link between wind stress di-  
11 vergence and downwind SST gradients. In this study, an idealized simula-  
12 tion representative of a storm track above a prescribed stationary SST field  
13 is examined in order to determine in which background wind conditions that  
14 relationship occurs. The SST field is composed of a mid-latitude large-scale  
15 frontal zone and mesoscale SST anomalies. It is shown that the divergence  
16 of the surface wind can either correlate with the Laplacian of the atmospheric  
17 boundary layer temperature or with the downwind SST gradient. The first case  
18 corresponds to background situations of weak winds or of unstable boundary  
19 layers and the response is in agreement with an Ekman balanced adjustment  
20 in the boundary layer. The second case corresponds to background situations  
21 of stable boundary layers and the response is in agreement with downward  
22 mixing of momentum. Concerning the divergence of the wind stress, it gen-  
23 erally resembles downwind SST gradients for stable and unstable boundary  
24 layers, in agreement with past studies. For weak winds, a correlation with the  
25 temperature Laplacian is still found to some extent. In conclusion, our study  
26 reveals the importance of the large-scale wind conditions in modulating the  
27 surface atmospheric response with different responses in the divergences of  
28 surface wind and wind stress.

## 29 **1. Introduction**

30 Satellite measurements have shown evidence of a local response of the atmospheric boundary  
31 layer to oceanic mesoscale structures (ranging from tens to hundreds of km). It takes the form  
32 of a positive correlation between wind stress and sea surface temperature (SST) anomalies at all  
33 latitudes (Xie 2004). Equivalent relationships exist with correlation of divergence of the wind  
34 stress with along-wind SST gradient, or wind stress curl and across-wind SST gradient (Chelton  
35 et al. 2001, 2004; O'Neill et al. 2003). It was also revealed through the signature of ocean eddies  
36 in turbulent air-sea fluxes of sensible and latent heat (Bourras et al. 2004), or in cloud cover and  
37 rain rates (Frenger et al. 2013).

38 The coupling between the atmosphere and narrow oceanic structures has been explored through  
39 various analyses of the horizontal-momentum budget in the boundary layer based on theoretical  
40 models (Samelson et al. 2006; Schneider and Qiu 2015) or idealized simulations (Spall 2007;  
41 Kilpatrick et al. 2014, 2016). The general setting of these analyses was a large-scale wind blowing  
42 across (or along) an SST gradient, potentially leading to a change in the stability of the boundary  
43 layer. In locally unstable conditions (i.e. winds blowing from cold to warm waters), an increase  
44 of the downward transfer of momentum explains the correlation of wind or wind stress with SST  
45 anomalies (Wallace et al. 1989; Hayes et al. 1989). The mechanism of downward momentum  
46 mixing (hereafter DMM) was proposed to explain the relation between the divergence of wind  
47 stress and downwind SST gradients (e.g. Chelton et al. 2001; O'Neill et al. 2003).

48 Another mechanism that is considered in the literature is related to surface pressure variations  
49 induced by SST structures. It was initially proposed as an important source of coupling at tropical  
50 latitudes (Lindzen and Nigam 1987), and more recently as an important forcing for surface-wind  
51 convergence over mid-latitude SST fronts (Feliks et al. 2004; Minobe et al. 2008). The mechanism

52 is based on a thermal adjustment of the boundary layer to the underlying SST, which creates local  
53 variations of the hydrostatic pressure. Through a mechanism in terms of Ekman balance mass  
54 adjustment (hereafter EBMA), the divergence of the surface wind correlates with the Laplacian of  
55 sea level pressure. The latter is itself very close to the Laplacian of the atmospheric temperature if  
56 the boundary layer has adjusted to the underlying SST, which is more likely for low winds (Brachet  
57 et al. 2012; Lambaerts et al. 2013).

58 At mid-latitudes the importance of the pressure term compared to vertical mixing still remains  
59 unclear, largely depending on the spatial scales (Small et al. 2008) but also on the region of interest  
60 (Shimada and Minobe 2011) or on the season that is considered (Takatama et al. 2015). Moreover  
61 the two mechanisms can be active together to force a surface divergence response. For instance,  
62 in the Kuroshio Extension region, Putrasahan et al. (2013) show that the divergence of wind stress  
63 correlates with downwind SST gradients (see their Fig. 4). At the same time, divergence of surface  
64 wind correlates with the Laplacian of SST (see their Fig. 7)

65 Most past studies have examined the time-average response (at least weekly averages) or the  
66 transient response (a few hours) of the atmospheric boundary layer to SST anomalies. As pointed  
67 out by Liu and Zhang (2013), O'Neill et al. (2017) or Plougonven et al. (2018), the responses  
68 differ when considering averaged or transients fields. Here our goal is to determine the nature  
69 of the surface divergence response to mesoscale SST perturbations separating between classes  
70 of different large-scale wind conditions. For that purpose we use an idealized simulation of an  
71 atmospheric storm track above a frontal SST zone including a variety of oceanic structures of  
72 horizontal scales from 40 to 400 km.

73 Section 2 presents the configuration of the model with a brief description of the simulated storm  
74 track. We then document in section 3 the surface divergence response at the oceanic eddy scale  
75 by a composite analysis and we show that the simulations are consistent with observational results

76 such as those of Frenger et al. (2013). Section 4 describes the spatial organization of the boundary  
77 layer response, investigating how the response mechanisms change for different synoptic-wind  
78 configurations. Differences between the responses in wind divergence and wind stress divergence  
79 are also investigated. Section 5 summarizes the results of the previous sections and compares them  
80 with previous studies.

## 81 **2. Model description**

### 82 *a. General configuration*

83 The 3.6.1 version of the WRF model (Skamarock et al. 2008) is used to simulate a characteristic  
84 mid-latitude storm track above a prescribed SST field. The model integrates the nonhydrostatic  
85 compressible moist Euler equations. Microphysics is represented with the Kessler (1969) scheme,  
86 and convection with the Kain and Fritsch (1993) scheme. The model uses the Yonsei University  
87 (YSU) parametrization (Hong et al. 2006) for the atmospheric boundary layer in conjunction with a  
88 Monin-Obukhov parametrization for surface layers (MM5 scheme). We do not include the effect of  
89 ocean surface currents in the wind stress calculation although it is known to affect the atmospheric  
90 boundary layer above oceanic eddies (Renault et al. 2016; Takatama and Schneider 2017).

91 The Cartesian domain, periodic in the zonal direction  $x$ , is of size  $L_x \times L_y = 9216 \times 9216$  km.  
92 Horizontal resolution is set to 18 km, and fifty  $\eta$  levels are used for the hydrostatic-pressure  
93 vertical coordinate, equally spaced in pressure. Top pressure is set to 36 hPa, corresponding to  
94 an altitude of approximately 20 km and 13 levels are below 2 km of altitude. Free-slip boundary  
95 conditions are used at the poleward and equatorial walls of the domain, and  $y = 0$  corresponds to  
96 the equatorial side of the domain. A spatially varying Coriolis parameter is used with a largest

97  $\beta$  effect in the center of the domain. Typical values of these parameters correspond to 40°N (see  
 98 Appendix A).

99 The model is forced by using a gray-radiation scheme with an atmosphere transparent to water  
 100 vapor and clouds, as proposed by Frierson et al. (2006). This forcing allows to mimic simple  
 101 relaxation forcings on dry variables (e.g. Held and Suarez 1994), but with the sole dependence on  
 102 the SST field. The details of the radiative scheme are described in Appendix B.

### 103 *b. Oceanic forcing*

104 We prescribe the sea surface temperature field, stationary in time and composed of a large-scale  
 105 meridional gradient and an eddying component:

$$SST(x, y) = \overline{SST}(y) + \underbrace{e^{-(y-y_{sst})^2/l_0^2} \times F(x, y)}_{SST_{eddy}(x, y)} \quad (1)$$

106 The large-scale front is described by

$$\overline{SST}(y) = SST_{eq} - \frac{SST_{eq} - SST_{pol}}{2} \left( 1 + \tanh \left( \frac{y - y_{sst}}{l_{sst}} \right) \right) \quad (2)$$

107 with parameters defined in Table 1. SST ranges from  $SST_{pol} = 275$  K to  $SST_{eq} = 295$  K and is  
 108 characterized by a smooth transition between warm and cold waters with an SST gradient of the  
 109 order of 1 K/100 km.

110 The eddying component  $SST_{eddy}(x, y)$  is obtained from a snapshot of a 2D turbulent field  $F(x, y)$   
 111 of a surface quasi-geostrophic (SQG) model (Lapeyre and Klein 2006) run for a domain size  
 112 of  $L_x/2 \times L_y/2$  and extended by periodicity to the full domain. The SQG model was shown  
 113 to adequately represent the upper ocean dynamics at mesoscale (see review of Lapeyre 2017).  
 114  $F(x, y)$  is normalized to get a standard deviation of 1.1 K and its zonal average is set to zero. Then  
 115 it is multiplied with a Gaussian envelope to obtain the field  $SST_{eddy}$  located where the meridional  
 116 large-scale gradient of SST is the most intense.

117 Figure 1 shows the total SST field and the corresponding SST anomalies. The maximum value  
118 of  $|SST_{eddy}|$  is 5.0 K but mesoscale SST anomalies have a relatively moderate signature in the total  
119 SST field which is characterized by a frontal region between  $y \approx 3000$  and 6000 km (Fig. 1a).  
120 These anomalies display a variety of structures with mesoscale eddies of various diameters, as  
121 well as long and thin filaments of  $\sim 50$  km width attached to them (Fig. 1b, see also Fig. 5).

### 122 *c. Mean state of the troposphere*

123 A first simulation using  $\overline{SST}$  as surface boundary condition was run for 4 years. Starting from its  
124 final state, a new simulation was then integrated over 8 years using the SST defined in (1). Outputs  
125 are saved twice a day, and the first three months are discarded as a spin-up period when computing  
126 statistics. The dynamical equilibrium obtained by taking a time and zonal average is presented on  
127 Fig. 2.

128 A typical storm track forms as a response to the large-scale forcing: a tropospheric jet is located  
129 around  $y = 6000$  km with a maximum speed larger than  $25 \text{ m s}^{-1}$  around  $p = 250$  hPa. The height  
130 of the tropopause changes from 200 hPa on the equatorial side of the domain down to 400 hPa on  
131 the poleward side (not shown). The eddy poleward heat flux is maximum in the free troposphere at  
132 the center of the domain between  $y = 4000$  km and  $y = 6000$  km, while the eddy kinetic energy has  
133 its maximum slightly poleward at  $y = 5500$  km (not shown). The simulated storm track is weaker  
134 than the southern-hemisphere storm track for which the zonal jet reaches values of  $35 \text{ m s}^{-1}$  but  
135 has realistic features of midlatitudes baroclinic zones. A more detailed analysis of the response of  
136 the storm track to the oceanic mesoscale SST field is carried out in Foussard et al. (2019).

### 137 **3. Composite analysis at the oceanic eddy scale**

138 In order to assess the consistency of our idealized simulations with the observed relation between  
139 surface variables and SST anomalies, we first discuss the main features of the response of the  
140 atmospheric boundary layer to a typical mesoscale eddy. To that end, composites for cold and  
141 warm eddies are computed in the line of Park et al. (2006) or Frenger et al. (2013). For the sole  
142 purpose of identifying the position of the eddies, we use a method based on a wavelet packet  
143 decomposition (see details in Lapeyre and Klein 2006; Doglioli et al. 2007). The procedure is  
144 to decompose  $SST_{eddy}$  in elementary wavelets of compact support (using the Haar basis). Then  
145 wavelet coefficients smaller than a given value are filtered out. The field that is recomposed with  
146 the remaining wavelets is such that it is zero at a given point if it does not belong to an eddy. This  
147 allows to determine the precise location of each structure in order to compute the composites. The  
148 amplitude of an eddy is defined as the spatial average of the SST anomaly over the set of grid  
149 points within the eddy. The coordinates of its center are defined as their averaged values and the  
150 eddy radius is defined as  $R_{eddy} = \sqrt{\mathcal{A}/\pi}$ , with  $\mathcal{A}$  the area of the eddy (defined as the set of points  
151 belonging to the specific eddy). Only eddies with amplitude larger than 2 K are retained. This  
152 results in 16 warm and 16 cold eddies, with radii ranging from 81 to 145 km (see Fig. 1b).

153 For each eddy and each instantaneous snapshot, the large-scale background wind is defined  
154 as the average of the 10-meter wind within a square box of width equal to 10 radii centered on  
155 each eddy. This yields a direction (used for the rotation of different quantities) and an amplitude  
156 (used to separate strong and weak-wind conditions). For presentation of the composites, all fields,  
157 including sea surface temperature, are rotated so that the large-scale wind blows towards  $x > 0$ , and  
158 are translated so that the eddy center is at  $(x, y) = (0, 0)$ . No spatial filtering has been applied to  
159 create the composites. Derivatives and Laplacian are computed using physical coordinates before

160 rotation and translation are made. At the end, spatial coordinates are rescaled in units of eddy radii  
161  $R_{eddy}$ . Composites of surface wind speed (hereafter 10 m-winds) and SST created through this  
162 procedure show the usual response with accelerated (compared to environment) winds over warm  
163 SST anomalies and decelerated winds over cold SST anomalies (Fig. 3). Note that the asymmetry  
164 between warm and cold eddies in terms of wind acceleration or deceleration cannot be interpreted  
165 since too few eddies served to create the composite fields.

166 We now turn to the analysis of surface wind divergence. Rather than computing the mean diver-  
167 gence, we choose to separate the response depending on the large-scale wind speed. To that end,  
168 we have selected conditions with large-scale winds larger than  $10 \text{ m s}^{-1}$  (to be called strong-wind  
169 conditions) and smaller than  $3 \text{ m s}^{-1}$  (to be called weak-wind conditions). These categories corre-  
170 spond to 33% and 7% of instantaneous snapshots respectively. In the following, we only consider  
171 the response to warm eddies as the results with cold eddies are qualitatively similar, but with an  
172 opposite sign (not shown). Finally, we have tested that changing the thresholds does not change  
173 qualitatively the results.

174 The divergence of the surface wind reveals significant differences between strong and weak-  
175 wind conditions (Fig. 4a and 4d). Strong-wind conditions (Fig. 4a) are characterized by a dipolar  
176 spatial pattern with a divergent wind field upwind of the eddy and a convergent wind field down-  
177 wind, with a typical amplitude of the order of  $10^{-5} \text{ s}^{-1}$ . This is consistent with accelerated wind  
178 speeds over warm eddies and is similar to observations (e.g. Park et al. 2006; Ma et al. 2015).  
179 Note also that the downwind convergence is twice as large as the upwind divergence, which is  
180 generally not observed when doing averages over all weather conditions (e.g. Frenger et al. 2013).  
181 For weak-wind conditions (Fig. 4d), the situation is different as a strong monopolar convergence  
182 pattern is located slightly downwind of the warm eddy.

183 To determine the importance of the DMM and EBMA, the surface divergence was compared  
184 with the downwind SST gradient and the Laplacian of atmospheric temperature in the boundary  
185 layer. The downwind SST gradient  $\mathbf{k} \cdot \nabla SST(x, y, t) = (\mathbf{U}_{10m}(x, y, t) / |\mathbf{U}_{10m}(x, y, t)|) \cdot \nabla SST(x, y)$   
186 was computed for each time output and grid point, then put in the new reference frame. Figures 4b  
187 and 4e show this quantity, for strong and weak winds in the case of warm eddies. Due to our  
188 specific definition, the downwind SST gradient is different in amplitude for strong and weak-  
189 wind conditions (Figs. 4b and 4e) but, in both cases, we recover the standard dipolar pattern. For  
190 strong-wind conditions, the shape of the downwind SST gradient is similar in part to the shape  
191 of the surface divergence (compare Figs. 4a and 4b), except for the surface convergence zone  
192 that extends further downstream (Fig. 4a). Another difference is that the downwind SST gradient  
193 has positive and negative poles with almost equal amplitude contrary to surface divergence for  
194 which some asymmetry is apparent. In weak-wind conditions, the downwind SST gradient differs  
195 from surface convergence with a monopolar shape for the later and a dipolar shape for the former  
196 (compare 4d and 4e).

197 For each wind condition, the Laplacian of boundary layer temperature  $\nabla^2 \theta$  was computed as the  
198 Laplacian of the temperature averaged between the surface and 500 m. It is represented in Figs. 4c  
199 and 4f for strong and weak winds. For strong winds, it is intensified and negative in the downwind  
200 side of the SST anomaly and is located close to the region of largest surface convergence (com-  
201 pare Figs. 4a and 4c). It thus seems that both the temperature Laplacian and the downwind SST  
202 gradient contribute in shaping the surface divergence pattern. This suggests that both DMM and  
203 EBMA may be important in setting the spatial variation of the surface divergence field. This result  
204 contrasts with the literature (e.g. Kilpatrick et al. 2016) as, in general, the downwind SST gradient  
205 seems the dominant parameter especially at high winds. A notable difference with these studies  
206 is that they only consider simplified configurations with quasi-unidirectional fronts, so that the

207 temperature Laplacian only comes from either the along or the cross-wind direction. On the con-  
208 trary, due to the geometry of oceanic eddies, the Laplacian can have variations in both directions.  
209 Indeed, in our simulation, it is found that about two thirds of the pressure Laplacian correspond to  
210 crosswind variations of pressure (not shown). Finally, in comparison to the temperature Laplacian,  
211 the SST Laplacian is centered over the oceanic eddy and is out of phase with the surface diver-  
212 gence (not shown). This is easily explained as the temperature anomaly that is generated above the  
213 warm oceanic eddy is advected downwind, so that SST and atmospheric temperature Laplacian do  
214 not correlate.

215 For weak-wind conditions, the temperature Laplacian is monopolar and negative above the SST  
216 anomaly because of weak temperature advection by the wind (Fig. 4f). Comparing Fig. 4d, 4e and  
217 4f, we see that, the surface divergence pattern is highly correlated with the temperature Laplacian  
218 , while it is not the case when compared to the downwind SST gradient. Actually, because of the  
219 weak temperature advection, the SST Laplacian is correlated with the temperature Laplacian as  
220 well as with the surface divergence (not shown). This is in agreement with the results of Lambaerts  
221 et al. (2013) who examined the fast adjustment of the boundary layer from rest to a turbulent eddy  
222 SST field. A possible interpretation of this result can rely on the EBMA mechanism: the warm  
223 SST anomaly creates a warm temperature anomaly in the boundary layer, which then creates a  
224 convergence field in the Ekman layer.

225 A last remark concerns moderate-wind conditions (i.e. winds between  $3 \text{ m s}^{-1}$  and  $10 \text{ m s}^{-1}$ ).  
226 In such conditions, it was found that the wind divergence response is between those for the two  
227 other wind conditions (not shown).

228 The difference in terms of the atmospheric response between weak and strong-wind conditions is  
229 reminiscent of results obtained by Chen et al. (2017) for eddies in the Kuroshio Extension region.  
230 In their study, they separated two different classes, one with a dipolar pattern in divergence of

231 surface wind (corresponding to 60% of the oceanic eddies that were observed) and one with a  
232 monopolar pattern (corresponding to 10% of the eddies). The first class was attributed to DMM  
233 while the second class to EBMA. An inspection of their Fig. 3c shows that, for the first class of  
234 eddies, the convergence maximum extends further downstream, a result consistent with our result  
235 for strong-wind conditions (Fig. 4a).

#### 236 4. Atmospheric response to a turbulent field of mesoscale eddies

237 The previous section has characterized the response of the wind field at the oceanic eddy scale  
238 in a simulation forced by the ocean with no ocean-atmosphere coupling. It showed that our simu-  
239 lation with fixed SST compares well with observations for strong-wind conditions. We now turn  
240 to examine the spatial organization of the atmospheric response in relation with the oceanic tur-  
241 bulent field, i.e. for scales smaller than 400 km. This contrasts with studies focusing on eddy  
242 composites or unidimensional fronts. To this end, we focus on a part of the spatial domain, of  
243 width  $1400 \times 1400$  km and centered at  $(x_0, y_0) = (5400, 4500)$  km, i.e. close to the center of the  
244 SST front. Results that are discussed hereafter apply for other spatial regions as well within the  
245 band where oceanic eddies are present.

246 In the following, we consider anomalies from the large-scale environment. These turbulent-scale  
247 anomalies, denoted as  $(\ )^*$ , are obtained (except for SST) by removing a large-scale component  
248 obtained by convoluting with a Gaussian kernel of radius  $r_{filter} = 200$  km. The SST anomaly  
249  $SST_{eddy}$  is given directly from the boundary condition through (1).

250 The anomaly of time-mean surface wind speed  $\langle |\mathbf{U}| \rangle^*$  is presented in Fig. 5a. Here  $\langle \ \rangle$  is the  
251 time average for the whole analysis period. It bears striking similarities with  $SST_{eddy}$  with a cor-  
252 relation coefficient of  $r = 0.98$  and a regression coefficient of  $0.29 \text{ m s}^{-1} \text{ K}^{-1}$ . This is true for the  
253 anomalies associated with oceanic eddies, confirming results of the last section, but also for the

254 filamentary structures in between. The regression coefficient (also called coupling coefficient) is  
255 in the range of the usual values derived from observations (e.g. O’Neill et al. 2012) or from models  
256 (Song et al. 2009; Perlin et al. 2014).

257 The time-mean response in the surface winds generally reflects convergence above warm eddies  
258 (such as eddies B or D in Fig. 5b) and divergence above cold eddies (eddies A or C). The diver-  
259 gence field does not bear resemblance with the downwind SST gradients (Fig. 5c) while there is  
260 a high correlation with the Laplacian of temperature in the boundary layer (compare shadings in  
261 Figs. 5b and d). As discussed in the previous section, such a comparison is not helpful to reveal  
262 in which wind conditions EBMA or DMM are important. This is probably due to the fact that, in  
263 this region, the time-mean wind is weak (not shown).

264 We propose below to contrast conditions of strong and weak winds as well as different wind  
265 directions to better assess the role of the background wind and of the stability of the boundary layer.  
266 Several effects are anticipated: the wind speed will influence both how turbulent the boundary  
267 layer is, and how much advection decorrelates boundary layer temperature from SST. The direction  
268 of the wind will also play a role through the presence of a large-scale meridional SST gradient. For  
269 example, northerly winds will advect cold air above warm waters, inducing a larger temperature  
270 difference between ocean and atmosphere, and hence a more turbulent boundary layer.

### 271 *a. Method*

272 Composite atmospheric fields depending on large-scale wind conditions are built through the  
273 following steps. We consider the square box of size  $900 \times 900$  km, centered at  $(x_0, y_0)$  and  
274 located inside the previously used  $1400 \times 1400$  km domain. The chosen box is large enough to  
275 be free of local wind variations induced by the SST anomalies, but not too large in order to cover  
276 separate synoptic weather patterns. To double the statistics, we also used the box centered at

277  $(x'_0, y'_0) = (1208, 4500)$  km since the SST eddy field was duplicated in longitude. We introduce  
 278 the wind conditions as the couples  $\mathbf{U}_{ls} = (U_{ls}, V_{ls})$  for  $U_{ls} = -10, -5, 0, 5, 10$  m s<sup>-1</sup> and  $V_{ls} =$   
 279  $-10, -5, 0, 5, 10$  m s<sup>-1</sup> (*ls* for large-scale). Then, for each 12-hour output, the instantaneous wind  
 280 at 10 meters,  $\mathbf{U}_{10m}$ , is averaged over the (900 km × 900 km) box and is sorted out according to  
 281 which wind conditions  $(U_{ls}, V_{ls})$  it belongs (within  $\pm 2.5$  m s<sup>-1</sup>). Composite fields [ ] are finally  
 282 constructed by averaging over all outputs belonging to each large-scale wind condition  $(U_{ls}, V_{ls})$ .  
 283 In the following, we will consider composites for which more than 100 time outputs have been  
 284 averaged. Finally, we introduce  $\mathbf{k} = [ \mathbf{U}_{10m}(x, y, t) ] / [ |\mathbf{U}_{10m}(x, y, t)| ]$  as a composite vector in the  
 285 wind direction and  $\theta$  the average temperature from the surface to 500 m height.

### 286 *b. Surface wind divergence*

287 We now examine the differences of spatial structures in surface wind divergence for different  
 288 wind conditions. Three large-scale wind conditions are considered in details: northerly strong  
 289 winds ( $\mathbf{U}_{ls} = (0, -10)$  m s<sup>-1</sup>), weak winds ( $\mathbf{U}_{ls} = (0, 0)$  m s<sup>-1</sup>) and southerly strong winds ( $\mathbf{U}_{ls} =$   
 290  $(0, 10)$  m s<sup>-1</sup>).

291 Figures 6a and 6b present the surface divergence anomaly  $[\nabla \cdot \mathbf{U}_{10m}]^*$  (in colors) as well as  
 292  $\mathbf{k} \cdot \nabla SST_{eddy}$  and  $[\nabla^2 \theta]^*$  (in contours), for northerly wind conditions (i.e.  $\mathbf{U}_{ls} = (0, -10)$  m s<sup>-1</sup>). At  
 293 first glance, both  $[\nabla^2 \theta]^*$  and  $\mathbf{k} \cdot \nabla SST_{eddy}$  seem to correlate well with the surface wind divergence  
 294 (correlation coefficients of  $r = 0.81$  and  $0.63$ , respectively, see Table 2). However inspecting with  
 295 more attention Figs. 6a and 6b, we note that, at some particular locations, the spatial structures of  
 296 the downwind SST gradient and the temperature Laplacian are quite different. First, narrow SST  
 297 structures oriented parallel to the background wind such as the one to the South West of eddy A  
 298 produce patterns of wind convergence ( $[\nabla \cdot \mathbf{U}_{10m}]^* < 0$ ) while  $\mathbf{k} \cdot \nabla SST_{eddy}$  is almost zero (Fig. 6a).  
 299 At this location, surface wind convergence is collocated with negative values of SST Laplacian (not

300 shown) and with negative values of  $[\nabla^2\theta]^*$  (see Fig. 6b). Also, for the the small warm eddy D,  
301 only a monopolar pattern of convergence of surface winds is seen which differs from the dipolar  
302 pattern of  $\mathbf{k} \cdot \nabla SST_{eddy}$  (Fig. 6a). In fact, at this location, the convergence region is associated  
303 with large values of temperature Laplacian  $[\nabla^2\theta]^*$  (Fig. 6b). The significant correlation ( $r = 0.81$ )  
304 between surface wind divergence and temperature Laplacian and the similarity of spatial structures  
305 suggest that, for strong northerly winds, the surface wind divergence response is mostly due to  
306 EBMA. However because of temperature advection by the northerly wind, correlation of surface  
307 divergence with the SST Laplacian itself remains low, with a correlation coefficient of 0.18.

308 The weak-wind case is represented in Figs. 6c and 6d. The surface divergence is found to be  
309 generally weaker than for northerly winds (compare Figs. 6a and 6c). Looking at Fig. 6d, surface  
310 divergence and temperature Laplacian are well correlated (with a correlation coefficient of 0.63).  
311 Also, there is a fair correspondence between SST Laplacian and surface divergence (correlation  
312 coefficient of 0.39), because the temperature anomalies lie almost above the SST anomalies (not  
313 shown). At particular locations (near eddy A, or in some filamentary structures in the northern part  
314 of the domain), the surface divergence resembles the downwind SST gradient (Fig. 6c). However,  
315 in many other places (such as eddies B, C, D), the two fields do not coincide with each other. We  
316 conclude that, in these weak-wind conditions, there is a preferential response following EBMA.

317 The situation is different for a southerly wind (Figs. 6e and 6f) for which we see a clear corre-  
318 lation of the surface divergence with the downwind SST gradient (correlation coefficient of 0.83  
319 see Tab. 2). This manifests in similar spatial structures not only for eddies B, C and D but also  
320 for the filamentary structures between them. The response above eddies B and C shows a typical  
321 dipolar structure of convergence/divergence corresponding to a DMM response. On the contrary,  
322 the connection between the surface divergence and the temperature Laplacian is less obvious when

323 comparing the spatial structures of the two fields (Fig. 6f), although the spatial correlation is still  
 324 high (around 0.48).

325 We now explore more quantitatively and systematically the atmospheric response by computing  
 326 the correlation coefficients of surface divergence with either the downwind SST gradient (Fig. 7a)  
 327 or the temperature Laplacian (Fig. 7b) as a function of the wind conditions ( $U_{ls}, V_{ls}$ ). Two regimes  
 328 can be distinguished. The first one, for strong southerly ( $V_{ls} > 0$ ) or zonal winds, corresponds to a  
 329 better correlation of the surface divergence with the downwind SST gradient than with the temper-  
 330 ature Laplacian. On the contrary, for northerly ( $V_{ls} < 0$ ) or for weak winds, the surface divergence  
 331 better correlates with the temperature Laplacian. However, there is still some correlation with  
 332 the downwind SST gradient. This last result can be understood by the correlation that is found  
 333 between  $\mathbf{k} \cdot \nabla SST_{eddy}$  and  $[\nabla^2 \theta]^*$ , as shown in Fig. 7c. A simple explanation of this correlation  
 334 comes from the heat budget which can be approximated by

$$\mathbf{U}_{ls} \cdot \nabla [\theta]^* \approx \gamma (SST_{eddy} - [\theta]^*)$$

335 where the air-sea heat flux was replaced by a simple relaxation towards SST with a typical  
 336 timescale  $\gamma^{-1}$ . If  $\mathbf{U}_{ls}$  points towards  $x > 0$  with constant modulus, the quantity  $\mathbf{U}_{ls} \cdot \nabla$  can be  
 337 replaced by  $|\mathbf{U}_{ls}| \partial / \partial x$ . After some algebra, we have

$$\frac{1}{\gamma} |\mathbf{U}_{ls}| \frac{\partial^2 [\theta]^*}{\partial x^2} + \mathbf{k} \cdot \nabla [\theta]^* = \mathbf{k} \cdot \nabla SST_{eddy} \quad (3)$$

338 For a sufficiently strong wind (i.e.  $|\mathbf{U}_{ls}| \times |\partial_x SST_{eddy}| \gg \gamma |SST_{eddy}|$ ), the temperature anomaly  
 339 above the surface heating will be advected downstream and the first term in the l.h.s. of (3) will  
 340 dominate the second term. In this situation, we obtain a balance between downwind SST gradient  
 341 and temperature Laplacian which explains the correlation between the two quantities. This is  
 342 particularly true for strong winds with a southward component (Fig. 7c and table 2). Note however

343 that it involves the second derivative only along the wind direction, so that the total Laplacian may  
344 not be systematically related to the downwind SST gradient.

345 To understand why surface divergence correlates with downwind SST gradient in some situ-  
346 ations, and with the temperature Laplacian in others, we examine the dependence on the wind  
347 conditions of the boundary layer height and the air-sea temperature difference, both spatially aver-  
348 aged over the domain of Fig. 5. The result is displayed in Fig. 8 and is significant in the sense than  
349 the mean change of both quantities between different wind conditions is larger than their change  
350 across the SST front (for a given wind condition). Northerly winds tend to be associated with high  
351 boundary layers (Fig. 8a) and an atmospheric temperature much colder than the underlying SST  
352 (Fig. 8b). This can be explained by the advection of cold air from the North, tending to decrease  
353 stability over the region that is examined. This results in a typical situation of strong turbulence  
354 in the boundary layer associated with a deep boundary layer. Southerly winds are associated with  
355 warm air advected in the region creating a stable boundary layer (Fig. 8b), which results in shallow  
356 boundary layers (Fig. 8a). These differences can explain the different response in terms of wind  
357 divergence as the surface pressure anomaly and the surface divergence tend to be proportional to  
358 the height of the boundary layer (Feliks et al. 2004). Conditions with higher boundary layers will  
359 result in stronger EBMA. This can be confirmed by examining the coupling coefficient, computed  
360 as the regression coefficient between wind speed anomalies and SST anomalies as a function of  
361 the background wind ( $U_{ls}$ ,  $V_{ls}$ ). The coupling coefficient is the smallest for northerly winds cor-  
362 responding to large-scale unstable boundary layers (Fig. 8c). For zonal or southerly winds (corre-  
363 sponding to large-scale stable boundary layers), the coupling coefficient increases with the wind  
364 speed, in agreement with Byrne et al. (2015) in their simulation of the Southern Ocean. This con-  
365 firms that DMM is more efficient for southerlies, resulting in higher correlation between surface  
366 divergence and downwind SST gradient. As suggested by Skyllingstad et al. (2007) and Small

367 et al. (2008), the surface stability, rather than the boundary layer depth, is the more susceptible to  
368 explain this behavior.

369 Note that we repeated the analysis and compared the vertical velocity at 500 m ( $w^*$ ) with down-  
370 wind SST gradient or temperature Laplacian. For the weak-wind conditions (Fig. 9), as well as for  
371 the northerly case, the vertical velocity strongly correlates with the temperature Laplacian (with  
372 reversed sign) and not with the downwind SST gradient (Table 2). This is true at the scales of the  
373 main eddies, as well as at the scales of filaments (not shown). The correlation with the SST Lapla-  
374 cian remains small even in the weak wind case (correlation coefficient of 0.42 compared with 0.73  
375 for the temperature Laplacian). For southerly winds, it is difficult to determine whether vertical  
376 velocities are similar to temperature Laplacian or to SST gradients. Instantaneous snapshots are  
377 sometimes characterized by a cold front present in the domain, in general oriented S-SW towards  
378 N-NE. Despite averaging over different snapshots, these fronts leave a residual signature in the  
379 vertical velocity field. As a result, the vertical velocity field does not display any organization at  
380 the scales of oceanic eddies (not shown).

### 381 *c. Wind stress divergence*

382 Several observational studies (O'Neill et al. 2003; Chelton et al. 2004) pointed out a robust  
383 relation between wind stress divergence and downwind SST gradient. We now try to relate this  
384 result with the response of the surface divergence that we analyzed above.

385 If we neglect the role of surface oceanic currents, the wind stress vector  $\tau$  is written using bulk  
386 formula,

$$\tau = \rho_0 C_d |\mathbf{U}_{10m}| \mathbf{U}_{10m} \quad , \quad (4)$$

387 with  $C_d$  the drag coefficient (Stull 1989). Divergences of surface wind and wind stress are then  
 388 related by

$$\nabla \cdot \boldsymbol{\tau} = \rho_0 \mathbf{U}_{10m} \cdot \nabla (C_d |\mathbf{U}_{10m}|) + \rho_0 C_d |\mathbf{U}_{10m}| \nabla \cdot \mathbf{U}_{10m} \quad . \quad (5)$$

389 The first term on the r.h.s. of (5) describes the effect of spatial variations of stress-to-wind ratio (i.e.  
 390  $|\boldsymbol{\tau}|/|\mathbf{U}_{10m}| = C_d |\mathbf{U}_{10m}|$ ). Since both  $C_d$  and  $|\mathbf{U}_{10m}|$  vary with air-sea temperature difference, and  
 391 hence to some extent with SST, we expect  $\mathbf{U}_{10m} \cdot \nabla (C_d |\mathbf{U}_{10m}|)$  to be proportional to the downwind  
 392 SST gradient. The second term on the r.h.s. of (5) describes the direct effect of the spatial variation  
 393 of the wind direction, and more generally the divergence of the wind vector. As seen above, for an  
 394 unstable boundary layer or for weak winds,  $[\nabla \cdot \mathbf{U}_{10m}]^*$  is generally proportional to the temperature  
 395 Laplacian, and this should also be the case for the last term of (5) as well.

396 To examine the sensitivity of wind stress divergence to the wind direction, we approximate (5)  
 397 by

$$[\nabla \cdot \boldsymbol{\tau}] \approx \underbrace{\rho_0 [\mathbf{U}_{10m}] \cdot \nabla [C_d |\mathbf{U}_{10m}|]}_{E_{stab}} + \underbrace{\rho_0 [C_d |\mathbf{U}_{10m}|] \nabla \cdot [\mathbf{U}_{10m}]}_{E_{div}} \quad (6)$$

398 with a constant density  $\rho_0 = 1.2 \text{ kg m}^{-3}$ . In the following, we will consider the anomalies from  
 399 the large-scale environment, e.g.  $[\nabla \cdot \boldsymbol{\tau}]^*$ . Quantities  $E_{stab}$  and  $E_{div}$  will refer to their anomalies.  
 400 Relation (6) was assessed and revealed to be valid with a r.m.s. error of about 20% and a good  
 401 correlation between the wind stress and its approximation (6). The error rises to 38% for weak  
 402 winds (because the sum of the two terms underestimates  $[\nabla \cdot \boldsymbol{\tau}]$  by 20%).

403 Figure 10a presents the divergence of the wind stress for northerly winds ( $\mathbf{U}_{1s} = (0, -10) \text{ m s}^{-1}$ )  
 404 while Figs. 10b and 10c present its two components following decomposition (6). Comparing  
 405 Figs. 10b and 10c, term  $E_{stab}$  is in general larger than term  $E_{div}$  (r.m.s. ratio of 1.69, see Ta-  
 406 ble 4). The role of DMM in shaping the wind stress divergence can be understood by realizing  
 407 that  $E_{stab}$  is proportional to the downwind gradient of SST with a correlation coefficient of 0.97

408 (see table 3). This relation reflects the fact that variations of the drag coefficient  $C_d$  and the surface  
 409 wind speed are closely linked to SST variations. Indeed, the correlation coefficient of  $[\nabla \cdot \tau]^*$  with  
 410  $[\mathbf{k}] \cdot \nabla S S T_{eddy}$  is 0.91 (Tab. 3) which is in agreement with the role of vertical stability in explaining  
 411 the spatial patterns of wind stress divergence. Term  $E_{div}$  tends to reinforce the divergence close  
 412 to the eddy centers (e.g. eddies B and D) and also explains a significant part of  $[\nabla \cdot \tau]^*$  above  
 413 filamentary structures in SST between eddies C and D or to the South West of eddy A. This is in  
 414 agreement with a correlation coefficient of 0.73 between wind stress divergence and temperature  
 415 Laplacian.

416 For weak-winds conditions ( $\mathbf{U}_{Is} = (0, 0)$ ), the wind stress is smaller than for the northerly case  
 417 (ratio of r.m.s. of 0.38).  $E_{div}$  is of comparable magnitude with  $E_{stab}$  (r.m.s. of  $0.40 \times 10^{-7}$  against  
 418  $0.46 \times 10^{-7} \text{ N m}^{-3}$ ). As in the previous case,  $E_{stab}$  is found to be correlated with downwind SST  
 419 gradients (see Fig. 10e) with a correlation coefficient of 0.94 while  $E_{div}$  correlates with the temper-  
 420 ature Laplacian (Fig. 10f) with a correlation coefficient of 0.70. Both terms significantly contribute  
 421 to the wind stress spatial pattern: term  $E_{div}$  generally dominates close to eddy centers (e.g. eddies  
 422 B and D ) or far from the eddies (e.g. North of eddies C and D, Fig. 10f) while  $E_{stab}$  dominates  
 423 in smaller scale structures at the eddy peripheries such as near eddies A and C (Fig. 10e). Weak  
 424 wind conditions are therefore prone to a correlation between wind stress divergence and tempera-  
 425 ture Laplacian, when considering scales around 200 km, while both downwind SST gradient and  
 426 temperature Laplacian matter for smaller scales.

427 We now consider the case of southerly winds, i.e.  $\mathbf{U}_{Is} = (0, 10) \text{ m s}^{-1}$  (Fig. 10g-i). First, the  
 428 wind-stress divergence has an opposite sign with the case of northerly winds (compare Fig. 10a  
 429 and 10g). This is related to the high correlation of  $[\nabla \cdot \tau]^*$  with  $\mathbf{k} \cdot \nabla S S T_{eddy}$  (correlation coefficient  
 430 of 0.93). Two reasons can be invoked:  $E_{stab}$  is in general 61% larger than  $E_{div}$  (table 4);  $E_{div}$  (as  
 431 well as surface divergence) is better correlated with  $\mathbf{k} \cdot \nabla S S T_{eddy}$  than with  $[\nabla^2 \theta]^*$ .

432 More generally, for all wind conditions of  $\mathbf{U}_{ls}$  (except for weak winds), the correlation between  
433 wind stress divergence and downwind SST gradient is higher than 0.80 (Fig. 11a). The correlation  
434 with the temperature Laplacian is smaller (Fig. 11b) but still increases to values around 0.7 for  
435 northerly winds. The wind stress divergence response is clearly different from the surface wind  
436 divergence, as the correlation with the downwind SST gradient always dominates in the first case  
437 (Figs. 11a and 11b), while both downwind SST gradient and temperature Laplacian were important  
438 for the second case (Figs. 7a and 7b)

439 Figure 11c shows the value of the regression coefficient between wind stress divergence and  
440 downwind SST gradient for all wind conditions. Values are of the typical range of those found in  
441 the literature (Perlin et al. 2014). The first dependence of the regression coefficient is on the wind  
442 speed, consistent with observations (e.g. O'Neill et al. 2012). It is modulated by the direction of  
443 the large-scale background wind relative to the large-scale front, in agreement with the coupling  
444 coefficient between wind speed and SST (Fig. 8c).

445 We conclude that the response in wind stress divergence to SST anomalies depends both on the  
446 magnitude of the mean surface wind and on the stability of the atmospheric boundary layer. In  
447 strong-wind conditions, we essentially find a wind stress divergence proportional to the downwind  
448 SST gradient. This is true for stable as well as for unstable boundary layers and is in agreement  
449 with the results of O'Neill et al. (2003) and Chelton et al. (2004). It can be understood as the  
450 addition of two effects in the wind stress. The first one comes from the variation of the drag  
451 coefficient and the surface wind speed due to SST (in relation with  $\nabla(C_d|\mathbf{U}_{10m}|)$ ) and generally  
452 dominates. It is responsible for a response related to the downwind SST gradient. The second  
453 effect comes from the direction and intensity of the wind (related to  $\nabla \cdot \mathbf{U}_{10m}$ ). Its contribution is  
454 large for unstable boundary layers in strong-wind conditions. In that case, the two effects add up

455 as they have similar spatial characteristics. For weak winds, the contribution of the second effect  
456 becomes as important as the first one, in particular at scales around 200 km.

## 457 **5. Conclusions**

458 In the present study, the response of surface winds to SST anomalies associated with oceanic  
459 eddies has been explored in an idealized simulation of an atmospheric storm track. Two mecha-  
460 nisms are generally invoked to explain the response in terms of divergence of surface wind and  
461 wind stress. A first one is related to pressure adjustment (EBMA mechanism, Lindzen and Nigam  
462 1987) while a second to downward momentum mixing (DMM mechanism, Wallace et al. 1989). It  
463 is expected that the surface wind divergence resembles the Laplacian of the atmospheric tempera-  
464 ture in the first case and the downwind SST gradient in the latter case. Our study has documented  
465 in which large-scale wind conditions one of the mechanisms is more active than the other. One  
466 advantage of our idealized simulation approach is that we could directly inspect the response in  
467 surface winds, contrary to other studies which considered equivalent neutral winds or wind stress.  
468 Also, using instantaneous winds averaged in composites (grouping together similar large-scale  
469 wind conditions) allows to separate the rapid response without a temporal filter, in a manner simi-  
470 lar to Byrne et al. (2015).

471 We first examined the response at the oceanic eddy scale through a composite analysis. It re-  
472 vealed that the surface wind divergence projects both onto the downwind SST gradient and onto  
473 the Laplacian of the atmospheric temperature in the boundary layer. For weak winds, the diver-  
474 gence of surface wind is proportional to the Laplacian of the boundary layer temperature. On  
475 the other hand, for strong winds, the surface divergence has a main pattern similar to the down-  
476 wind SST gradient, but with a downstream extension (related to the temperature Laplacian spatial  
477 extension).

478 The atmospheric response was then investigated over a large region including a field of  
479 mesoscale oceanic eddies and filaments of scales between 40 and 400 km. The analysis revealed  
480 a more complex response that depends on the wind conditions, and more generally on the mean  
481 stability of the boundary layer. For large-scale unstable conditions or for weak winds, the diver-  
482 gence of the surface wind is correlated with the temperature Laplacian (corresponding to EBMA),  
483 while for large-scale stable conditions, it is correlated with downwind SST gradient (correspond-  
484 ing to DMM). For strong winds, the correlation of the surface divergence with the SST Laplacian  
485 is found to be small, due to the effect of the mean-wind advection.

486 Concerning the response in terms of wind stress divergence, a different picture is obtained. For  
487 strong winds, the divergence of wind stress is proportional to downwind SST gradient, even in  
488 large-scale unstable conditions. For weak winds, wind stress divergence is proportional to some  
489 extent to the temperature Laplacian. These results are valid at the scales of oceanic eddies, as well  
490 as smaller filamentary scales. This discussion shows that wind stress and surface wind divergences  
491 may behave differently considering their response to SST anomalies. We point out that such  
492 a distinction is rarely made in the literature and should be given greater consideration. Actually  
493 wind stress is directly related to the stability of the boundary layer while horizontal velocities in the  
494 atmosphere are less so but have a strong dependence on gradients of boundary layer temperature.

495 Several studies have examined the relevant parameters that set the atmospheric response sensi-  
496 tivity to DMM (Spall 2007; Small et al. 2008; Schneider and Qiu 2015; Ayet and Redelsperger  
497 2019). The first one is related to the magnitude of the mean wind speed. Our study confirms that  
498 the relative importance of DMM increases with wind speed. This is shown by a better correlation  
499 of surface divergence with downwind SST gradient than with temperature Laplacian for strong  
500 winds, except in situations of winds blowing from cold to warm waters. A second important pa-  
501 rameter is the spatial scale of the SST field (Small et al. 2008). One would expect the smaller

502 the lengthscale, the larger the sensitivity to DMM. However, for strong winds blowing from cold  
503 to warm waters, we found that EBMA still dominates with surface divergence proportional to  
504 temperature Laplacian down to 40 km. The dominance of EBMA over DMM (in terms of re-  
505 lation between surface wind divergence and temperature Laplacian or downwind SST gradient)  
506 was found to be related to the large-scale stability of the boundary layer. For unstable and deep  
507 boundary layers, an EBMA response is found, while DMM prevails for large-scale stable condi-  
508 tions. This may be related to the dependence of the pressure Laplacian to the mean height of the  
509 boundary layer (which favors EBMA) and to the dependence of the coupling coefficient (between  
510 wind and SST anomalies) on the stability (which favors DMM).

511 As the focus of the paper concerns the boundary layer and surface dynamics, we did a sensi-  
512 tivity study to the boundary layer parametrization scheme, using the Mellor-Yamada-Nakanishi-  
513 Niino (MYNN) scheme (Nakanishi and Niino 2004). We obtained qualitatively similar results (see  
514 Fig. 12), but with different intensities in agreement with results of Lambaerts et al. (2013) and Per-  
515 lin et al. (2014). A sensitivity to the number of vertical levels within the first 1000 m showed a  
516 weak dependence of the results on vertical resolution as well.

517 The present study has different limitations. The first one is the stationarity in time of the oceanic  
518 anomalies especially for scales below 50 km. However, because of the fast variability of the  
519 atmosphere, conditions of given wind do not occur over long timescales compared to the SST  
520 variability of the ocean. Therefore the composite analysis focuses only on the rapid response of  
521 the atmosphere and not on its time average which is tightly linked to fixed SST. Another limitation  
522 is the fact that ocean-atmosphere coupling was not considered although different feedbacks are  
523 known to modify the surface wind response to oceanic mesoscale anomalies. In particular, our  
524 parametrization of the surface atmospheric layer does not take into account ocean currents mod-  
525 ulation on the wind work (Renault et al. 2016; Moulin and Wirth 2016; Takatama and Schneider

2017). Moreover the air-sea coupling tend to damp oceanic eddies through Ekman pumping at the scale of oceanic eddies (Stern 1965; Dewar and Flierl 1987) as well at the scale of a turbulent eddy field (Oerder et al. 2018). A full air-sea coupling could reduce the SST amplitude and modulate the atmospheric response. These different mechanisms need to be taken into account in future studies.

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## APPENDIX A

### Coriolis parameter

The Coriolis parameter  $f$  has the following dependence in  $y$ :

$$f(y) = f_0 + \beta_{max} l_\beta \tanh\left(\frac{y - y_{sst}}{l_\beta}\right) . \quad (A1)$$

This formula in addition to parameters in Table 1 allows to model a storm track with a Coriolis parameter that ranges between values at 27.6 and 55.6°N. The  $\beta$  effect above the oceanic front has a realistic value of  $\beta_{max} = 1.75 \times 10^{-11} \text{ m}^{-1} \text{ s}^{-1}$ , corresponding to its value at a latitude of 40°N. In this way, we obtain a strong planetary vorticity gradient, which helps to maintain the eddy-driven jet to its mean position. Using a linear function for  $f$  with the same value of  $\beta = \beta_{max}$  would lead to unrealistic values of  $f$  on the northern or southern part of the domain.

## APPENDIX B

546

## Radiative scheme

547 The radiative scheme that is used in our simulation is a gray radiation scheme following the  
 548 ideas of Frierson et al. (2006). We introduce  $T$  as the absolute temperature and  $D$  the optical depth  
 549 (with the convention  $D = 0$  at the top of the atmosphere and  $D = D_0$  at the surface). The equations  
 550 for upward ( $F^\uparrow$ ) and downward ( $F^\downarrow$ ) radiative energy fluxes are

$$\frac{dF^\uparrow}{dD} = F^\uparrow - \sigma T^4 \quad (\text{B1})$$

$$\frac{dF^\downarrow}{dD} = -F^\downarrow + \sigma T^4 \quad (\text{B2})$$

551 with  $\sigma$  the Stefan-Boltzmann constant. To close the system, the fluxes at the surface and at the top  
 552 of the atmosphere are such that

$$F^\uparrow(D = D_0) = \sigma SST^4 \quad (\text{B3})$$

$$F^\downarrow(D = 0) = 0 \quad (\text{B4})$$

553 This choice of boundary conditions is different from Frierson et al. (2006) and allows to constrain  
 554 the forcing to almost entirely depend on the SST field.

555 We prescribe total optical depth  $D$  to be dependent only on latitude  $y$  and pressure  $p$ . The surface  
 556 optical depth  $D_0(y, p = p_0) = D_0(y)$  is such that

$$D_0(y) = D_{eq} \cos^2\left(\frac{\pi y}{2 L_y}\right) + D_{pole} \sin^2\left(\frac{\pi y}{2 L_y}\right) \quad (\text{B5})$$

557 Then we separate optical depths in the troposphere and stratosphere by introducing  $D_T$  and  $D_S$   
 558 such that

$$D = \max(D_T, D_S) \quad (\text{B6})$$

559 with

$$D_S(y, p) = \frac{1}{4} \frac{p}{p_0} D_0(y) \quad , \quad (\text{B7})$$

$$D_T(y, p) = (1 + D_0) \left( \frac{p}{p_0} \right)^{4\kappa} \left( 1 - \frac{\Delta\theta}{\alpha \overline{SST}} \log \left( \frac{p}{p_0} \right) \right)^4 - 1 \quad . \quad (\text{B8})$$

560 Table 1 summarizes the values of the various parameters.

561 In (B8),  $\overline{SST}$  is the zonal average of SST and

$$\alpha = \left( \frac{1 + D_0}{2 + D_0} \right)^{1/4} \quad (\text{B9})$$

562  $D_S$  is larger than  $D_T$  in the highest atmospheric layers, and transition from one expression to the  
563 other roughly sets the height of the tropopause in our experiment.

564 In the model, the diabatic term due to radiative forcing is expressed in the temperature equation

565 as

$$R = \frac{1}{\rho C_p} \frac{\partial(F^\uparrow - F^\downarrow)}{\partial z} \quad (\text{B10})$$

566 To understand the nature of this forcing, we can compute the potential temperature at radiative  
567 equilibrium  $\theta_{Rad}$ , i.e. when  $R = 0$ . Below the tropopause,  $D(y, p) = D_T(y, p)$  and using (B1-B5)  
568 and (B8), we obtain

$$\theta_{Rad}(x, y) = \alpha SST(x, y) - \Delta\theta \log \left( \frac{p}{p_0} \right) \frac{SST(x, y)}{\overline{SST}(y)} \quad . \quad (\text{B11})$$

569 Since  $\alpha$  is weakly dependent on  $y$  (ranging between 0.92 and 0.96), the gray-radiation scheme  
570 relaxes temperature towards a profile whose meridional gradient is proportional to  $\partial\overline{SST}/\partial y$  in  
571 zonal mean. Such a profile is similar to the radiative equilibrium of Held and Suarez (1994).

572 To ensure that the boundary layer response does not depend on radiative parametrization choices,  
573 two other sensitivity runs were done. In the first one,  $\overline{SST}(y)$  was replaced by  $SST(x, y)$  in (B8)  
574 while in the second one,  $SST(x, y)$  was replaced by  $\overline{SST}(y)$  in (B3) While radiative fluxes act either  
575 as an additional heat source or sink at the scale of the oceanic eddies in each experiment, the impact

576 on the main heat budget remained small and no major differences were obtained concerning the  
577 results of this paper.

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705 Northerly, weak and southerly large-scale winds. . . . . 38

$(L_x, L_y)$ in km	(9216, 9216)
$(SST_{eq}, SST_{pole})$ in K	(295, 275)
$y_{sst}$ in km	4500
$(l_{sst}, l_0, l_\beta)$ in km	( 1000, 1500, 1500 )
$f_0$ in $s^{-1}$	$9.35 \times 10^{-5}$
$\beta_{max}$ in $m^{-1} s^{-1}$	$1.75 \times 10^{-11}$
$\Delta\theta$ in K	10
$(D_{eq}, D_{pole})$	(6, 1.5)
$p_0$ (in Pa)	$10^5$

TABLE 1. Common parameters.

	Northerlies	weak winds	Southerlies
$C([\nabla \cdot \mathbf{U}_{10m}]^*, [\nabla^2 \theta]^*)$	0.81	0.63	0.48
$C([\nabla \cdot \mathbf{U}_{10m}]^*, \mathbf{k} \cdot \nabla SST_{eddy})$	0.63	0.36	0.83
$C([\nabla \cdot \mathbf{U}_{10m}]^*, \nabla^2 SST_{eddies})$	0.18	0.39	0.08
$C([\nabla^2 \theta]^*, \mathbf{k} \cdot \nabla SST_{eddy})$	0.58	0.31	0.34
$C([w]^*, -[\nabla^2 \theta]^*)$	0.76	0.73	0.43
$C([w]^*, \mathbf{k} \cdot \nabla SST_{eddy})$	0.50	0.16	0.08
$C([w]^*, -\nabla^2 SST_{eddies})$	0.19	0.42	0.16

706 TABLE 2. Correlation coefficients between different parameters related to wind divergence for Northerly,  
707 weak and southerly large-scale winds. Each quantity was computed over the domain displayed in Fig. 5.

	Northerlies	weak winds	Southerlies
$C([\nabla \cdot \boldsymbol{\tau}]^*, [\nabla^2 \theta]^*)$	0.73	0.58	0.43
$C([\nabla \cdot \boldsymbol{\tau}]^*, \mathbf{k} \cdot \nabla S S T_{eddy})$	0.91	0.68	0.93
$C([\nabla \cdot \boldsymbol{\tau}]^*, \nabla^2 S S T_{eddies})$	0.07	0.27	0.05
$C(E_{stab}, \mathbf{k} \cdot \nabla S S T_{eddy})$	0.97	0.94	0.94
$C(E_{div}, \mathbf{k} \cdot \nabla S S T_{eddy})$	0.64	0.34	0.80
$C(E_{div}, \nabla^2 \theta^*)$	0.79	0.65	0.51

708 TABLE 3. Correlation coefficients between different parameters related to wind stress divergence for  
709 Northerly, weak and southerly large-scale winds.

	Northerlies	weak winds	Southerlies
$\text{rms}(E_{stab})$ in $10^{-7}$	1.44	0.46	1.78
$\text{rms}(E_{div})$ in $10^{-7}$	0.85	0.40	1.10
$\text{rms}([\nabla \cdot \tau]^*)$ in $10^{-7}$	2.09	0.80	2.96

710 TABLE 4. r.m.s. values of different parameters related to wind stress divergence for Northerly, weak and  
711 southerly large-scale winds.

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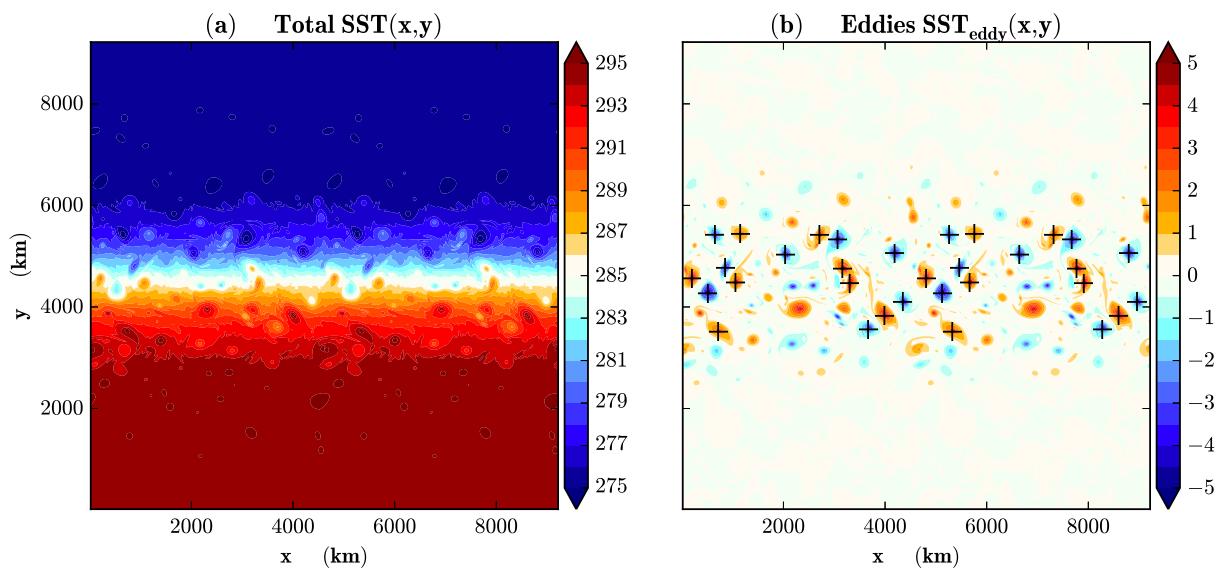
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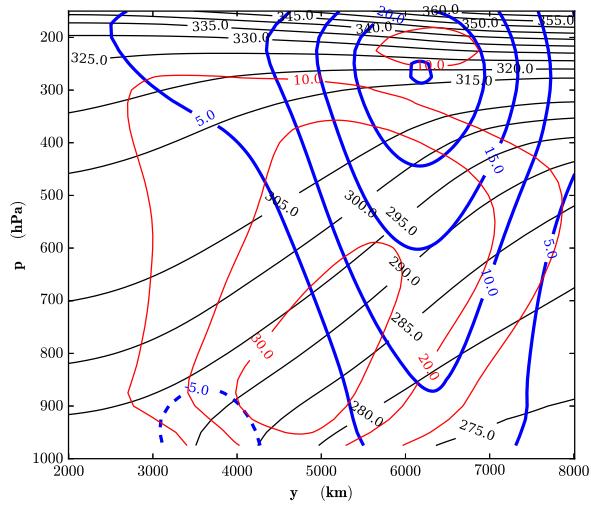
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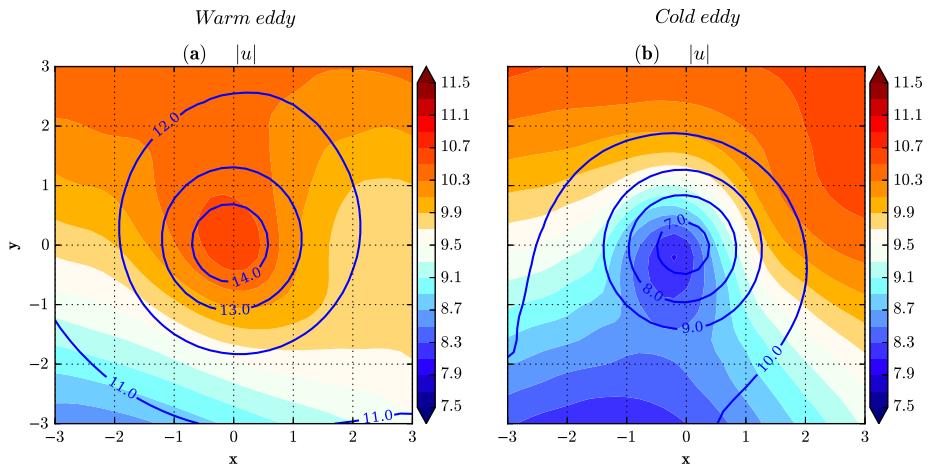
753 **Fig. 12.** Same as Fig. 7a and 7b in the simulation with the MYNN parameterization. . . . . 52



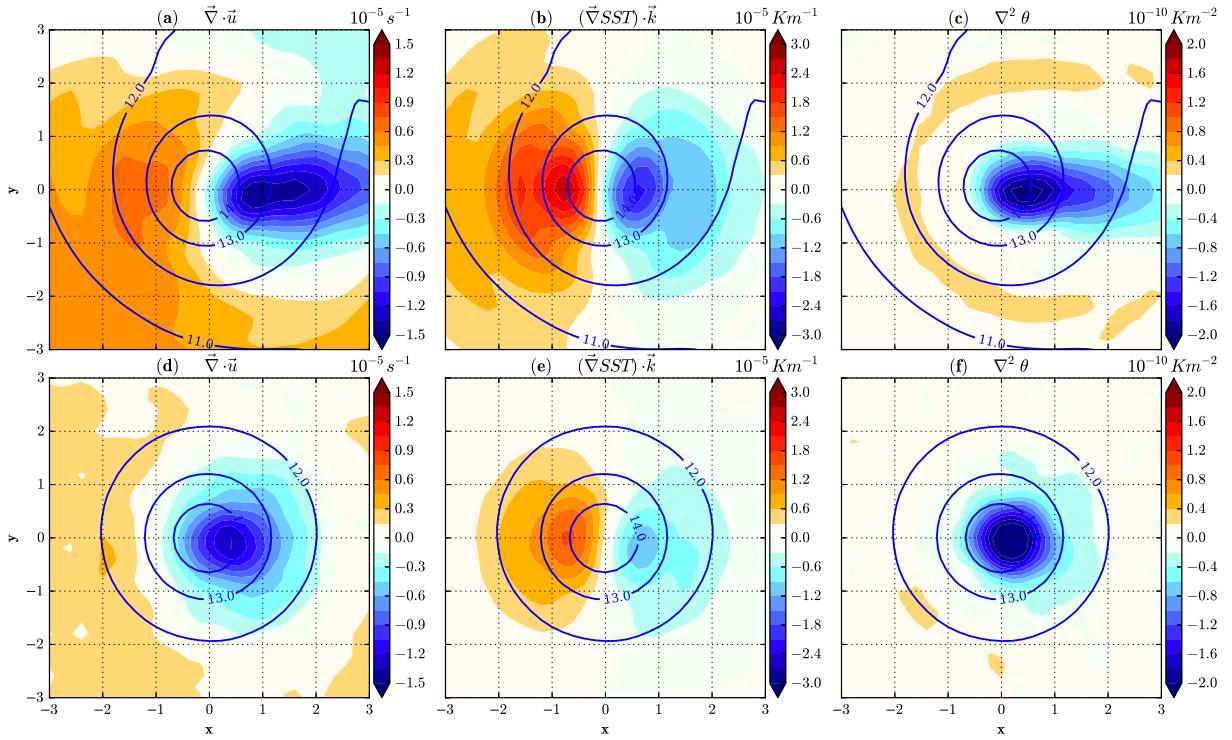
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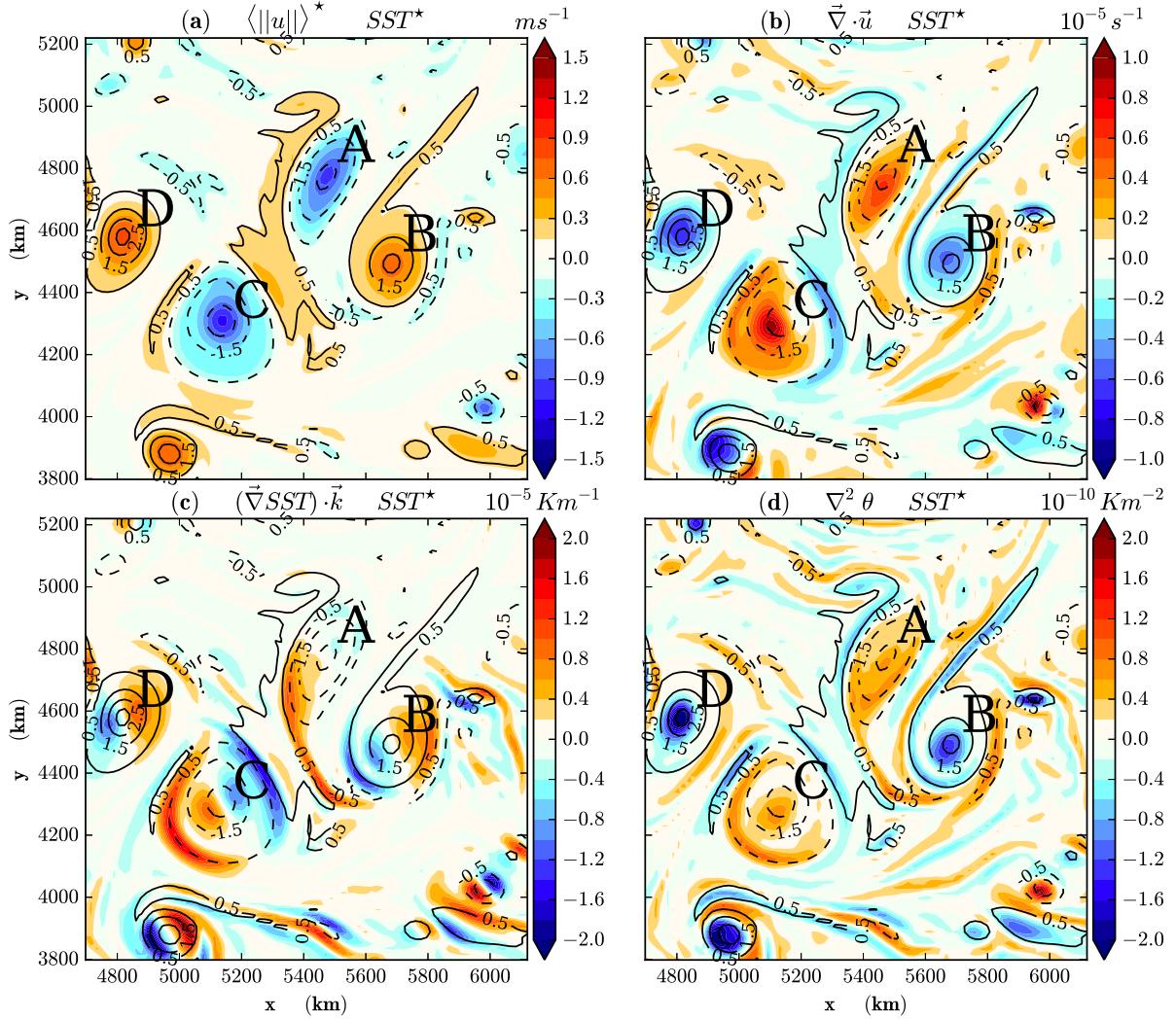
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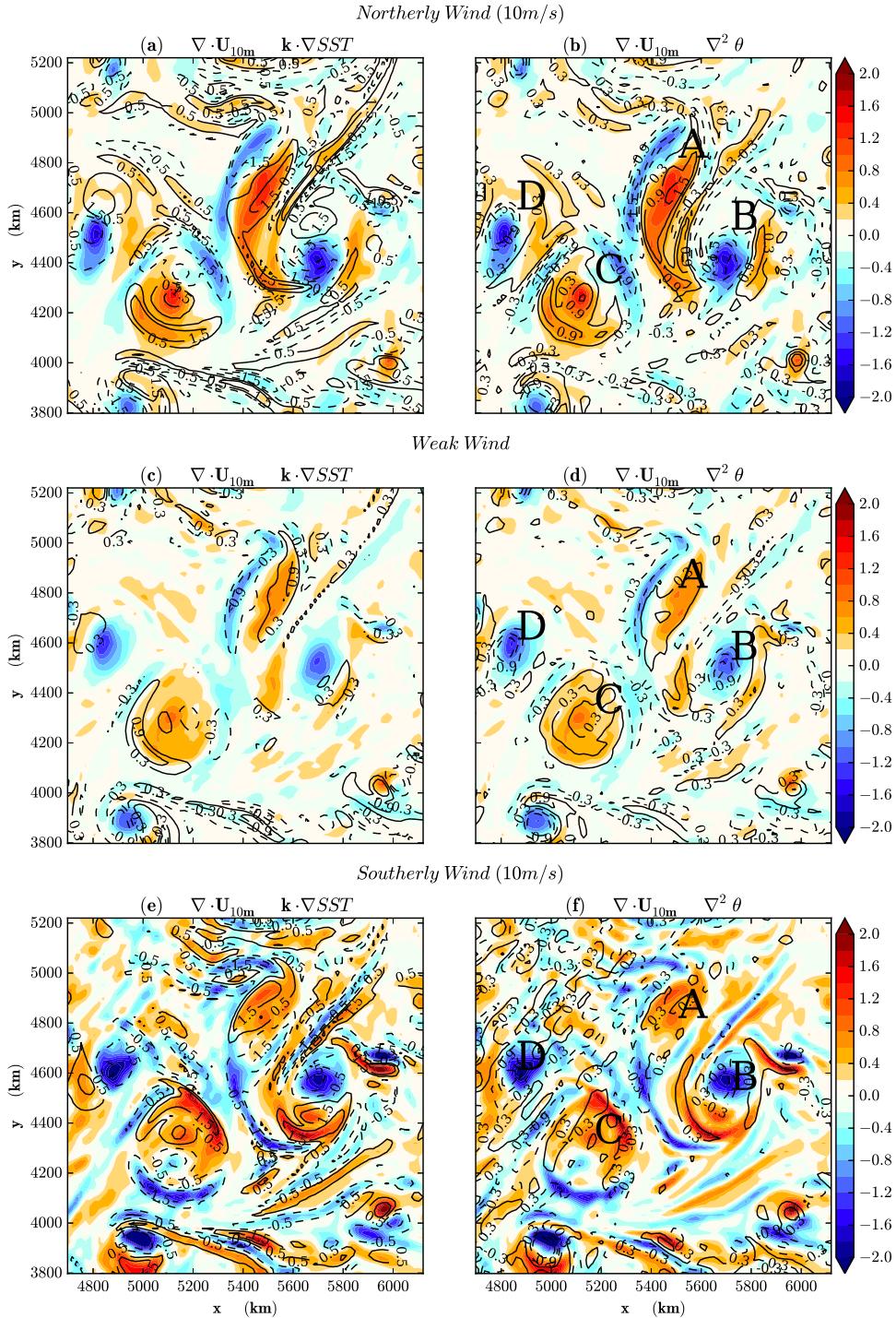
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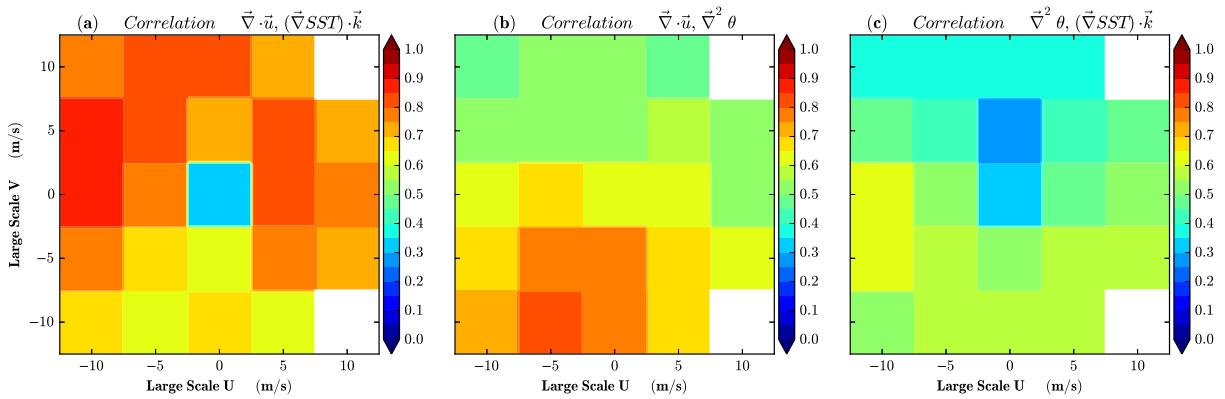
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 762 and (d-f) weak-wind conditions (wind speeds smaller than  $3 \text{ m s}^{-1}$ ). Contours correspond to SST (in K).



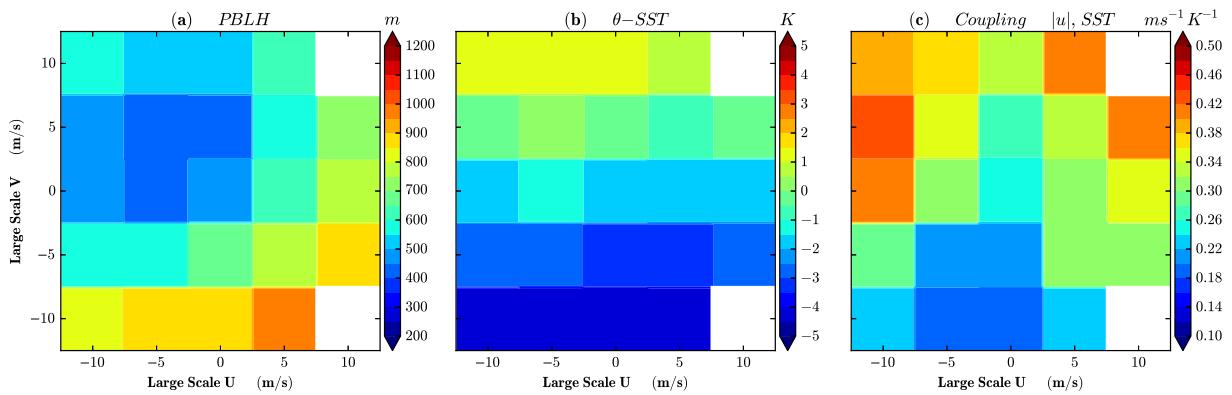
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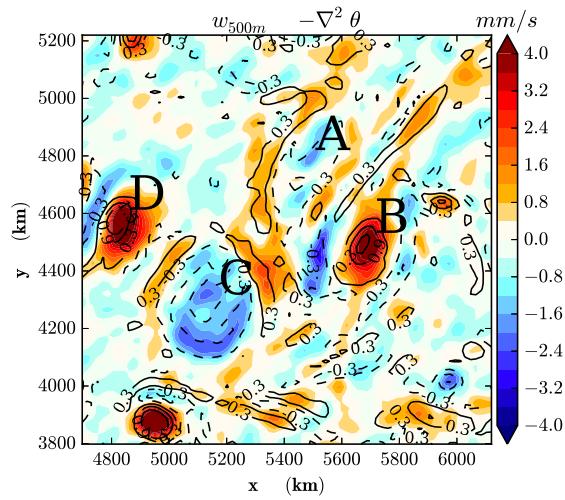
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 770  $10^{-10} \text{ K m}^{-2}$ ).



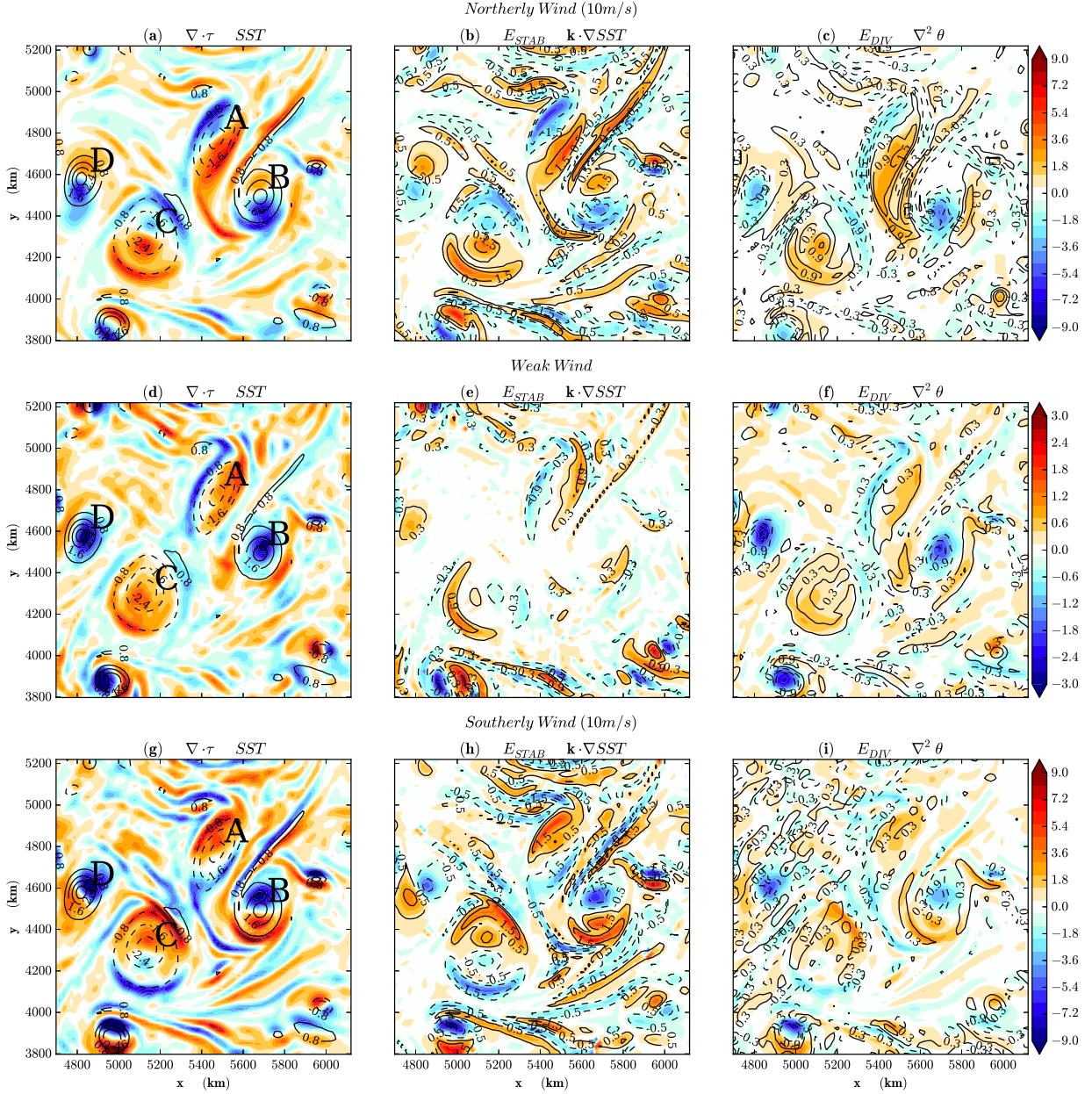
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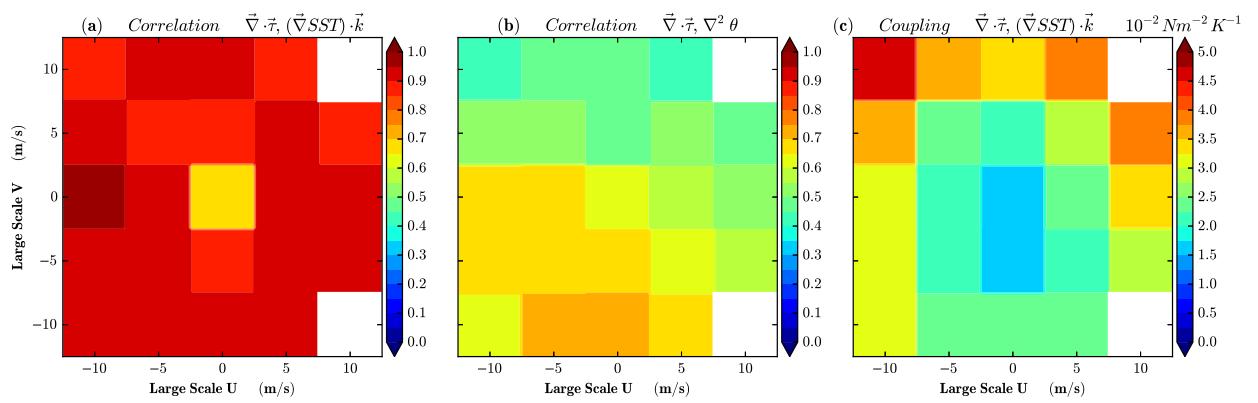
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781 FIG. 10. Composites of wind-stress divergence  $[\nabla \cdot \tau]^*$  (a, d, g), stability effect on wind-stress divergence  $E_{stab}$   
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 784 Contours in (c, f, i) correspond to  $[\nabla^2 \theta]^*$  (in  $10^{-10} \text{ K m}^{-2}$ ). (a, b, c) strong northerly winds; (d, e, f) weak  
 785 winds, (g, h, i) strong southerly winds.



786 FIG. 11. Correlation coefficient of wind stress divergence with (a) downwind SST gradient, (b) temperature  
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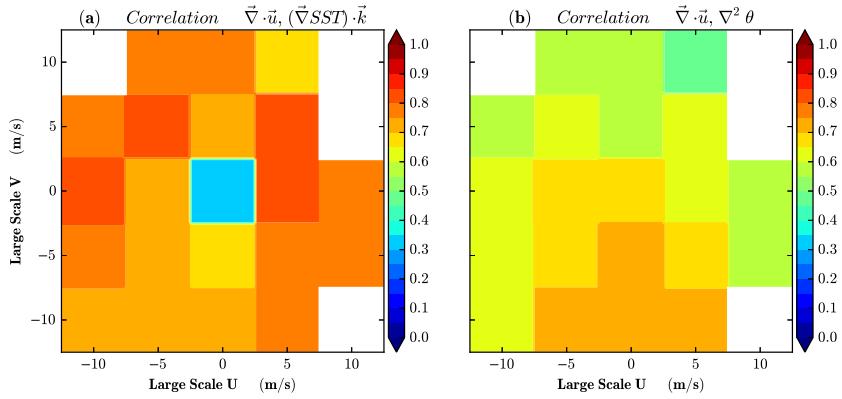


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