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2 **Main Manuscript for**

3 **Oxygen and Magnesium Mass Independent Isotopic Fractionation**
4 **Induced by Chemical Reactions in Plasma**

5

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15 **Author Contributions:**

16 F.R. designed the project and run the experiments. F.R. and M.C. wrote the paper and
17 interpretation. S.M. and A.G. analyzed the isotopic composition with the NanoSIMS.

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20

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23 Abstract – Significance Statement - Main Text – Materials and Methods – Acknowledgments –
24 References – Figure 1-3 with captions (embedded in the Text) – Schemes 1-3 (embedded in the
25 Text). Supplementary Informations.

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29

30 **Abstract**

31

32 Enrichment or depletion ranging from -40 to +100% in the major isotopes ^{16}O and ^{24}Mg were
33 observed experimentally in solids condensed from carbonaceous plasma composed of
34 $\text{CO}_2/\text{MgCl}_2/\text{Pentanol}$ or $\text{N}_2\text{O-Pentanol}$ for O and $\text{MgCl}_2/\text{Pentanol}$ for Mg. In NanoSims imaging,
35 isotope effects appear as micrometer size hotspots embedded in a carbonaceous matrix showing
36 no isotope fractionation. For Mg, these hotspots are localized in carbonaceous grains which show
37 positive and negative isotopic effects so that the whole grain has a standard isotope composition.
38 For O, no specific structure was observed at hotspot locations.

39 These results suggest that MIF (Mass Independent Fractionation) effects can be induced by
40 chemical reactions taking place in plasma. The close agreement between the slopes of the linear
41 correlations observed between $\delta^{25}\text{Mg}$ vs. $\delta^{26}\text{Mg}$ and between $\delta^{17}\text{O}$ vs. $\delta^{18}\text{O}$ and the slopes
42 calculated using the empirical MIF factor η discovered in ozone^(1,2) attests to the ubiquity of this
43 process.

44 Although the chemical reactants used in the present experiments cannot be directly transposed
45 to the protosolar nebula, a similar MIF mechanism is proposed for oxygen isotopes: at high
46 temperature, at the surface of grains, a mass independent isotope exchange could have taken
47 place between condensing oxides and oxygen atoms originated from the dissociation of CO or H_2O
48 gas.

49 **Significance Statement**

50 *Paste your significance statement here. Please note that it should not exceed 120 words, but should be at*
51 *least 50 words in length. It should not include any references.*

52 Both the physical effect and the chemical conditions at the origin of the oxygen isotope variations
53 in the solar system have been a puzzling question for 50 years. The data reported here with a new
54 experimental protocol bring the MIF effect (Mass Independent Fractionation) originally identified on
55 ozone back to the center of the debate. Similarly to Titanium isotopes, we observe that the MIF
56 effect for O and Mg is triggered by redox reactions in plasma. These observations reinforce the
57 idea of a universal mechanism observable in photochemical reactions when molecular collisions
58 involving indistinguishable isotopes yield a symmetrical complex stabilized as a chemical product.

59

60

61 **Main Text**

62

63 **Introduction**

64

65 Since their discovery in 1973 in the Calcium-Aluminium rich inclusions of the carbonaceous
66 chondrites⁽³⁾, it has been shown that large enrichments and depletions in ¹⁶O were ubiquitous in
67 the solar system, among meteorites, terrestrial planets and the Sun⁽⁴⁻⁶⁾ and a prominent feature of
68 atmospheric chemistry⁽⁷⁻¹⁰⁾. They were evidenced from the fact that, in a three oxygen isotopes
69 diagram, nearly all solar system samples have isotopic compositions (reported in ‰ variations as
70 $\delta^{17}\text{O}$ versus $\delta^{18}\text{O}$ values, $\delta^m\text{O}_{\text{sample}} = [(R_{\text{sample}}/R_{\text{standard}}) - 1] \times 1000$; with $R = {}^m\text{O}/{}^{16}\text{O}$ and $m = 17$ or
71 18) defining a linear correlation of slope close to 1 (hereafter referred to as 1:1 CL, the 1 to 1
72 Correlation Line), instead of a slope 0.52 for the “classical” mass dependent isotopic fractionations
73 (MDF) known to occur during physical and chemical processes⁽¹¹⁾. The question of the origin of the
74 1:1 CL is central to the formation of solids in the early solar system. It has been successively
75 proposed to result from: (i) the injection in the protosolar nebula (PSN) of pure ¹⁶O of supernovae
76 origin⁽³⁾, (ii) a mass independent fractionation (MIF) effect analogous to those observed
77 experimentally during the synthesis of ozone^(1,7), and (iii) a self-shielding effect on the solar (or non-
78 solar) UV light by CO⁽¹²⁻¹⁴⁾. While the lack of presolar grains enriched in ¹⁶O makes the first proposal
79 unlikely⁽¹⁵⁾, the two other ones have gained some recent theoretical^(14,16) or experimental^(17,18)
80 support. In the present paper, we address experimentally several issues raised by the ozone
81 experiment to explore whether this 1:1 CL could be due to MIFs reactions having taken place in the
82 PSN.

83 The formation of ozone results from a three-body reaction. However, at high temperature in a
84 PSN dominated by H₂, the low concentration of elements heavier than H mean that three-body
85 reactions cannot play an important role in the gas phase⁽¹⁶⁾. In order to overcome this difficulty, the
86 surfaces of growing grains were proposed as possible catalysts for the reactions leading to a MIF
87 effect⁽¹⁶⁾. Experimental evidence of this effect were reported for SiO/O₂/H₂ mixtures⁽¹⁷⁾ but with
88 variations of smaller magnitude than in solar system materials. The present paper is an additional
89 test of this theoretical proposal.

90 **Experimental**

91

92

93 Because of the difficulties to carry out controlled condensation experiments in hot plasmas, our
94 approach was not intended to mimic the conditions of the PSN but to answer specific questions
95 having key implications for cosmochemistry. Are oxygen MIFs linked to a precise class of chemical
96 reactions? Are they restricted to gas phase reactions or, as suggested by⁽¹⁸⁾, can they take place
97 during condensation of solids from a gas or be transferred from the gas to condensing solids? Can

98 the conditions of the appearance of MIFs in the laboratory be reasonably extended to the conditions
99 prevailing in the PSN? To this aim, we investigated reactions involving the isotopes of the two major
100 elements constituting the telluric planets, O and Mg. The comparison between isotopic effects on
101 O and Mg will bring important constraints because, at variance with O, all Mg isotopic fractionations
102 in meteorites are considered to result from MDF processes such as evaporation, condensation and
103 diffusion⁽¹⁹⁻²²⁾.

104 We have explored different reactions taking place in plasmas between gaseous species leading
105 to the condensation of O and Mg bearing compounds. A summary of these experiments is shown
106 in Table S1. Here we report only the results of experiments that have yielded MIF effects. The
107 experimental protocol is comparable to that previously described for H or Ti^(18, 23) and is detailed in
108 Material and Methods.

109 In microwave plasma, the molecules are dissociated by electron impact producing highly
110 reactive radicals. The high temperature of the gas enhances the rates of chemical or isotopic
111 reactions. In addition, this type of plasma allows the condensation of enough material for isotopic
112 analyses. The aim of these plasma experiments was to produce isotopic exchanges combined with
113 chemical reactions. For oxygen isotopes we studied the reactions between Pentanol and N₂O or
114 between a solution of Pentanol-MgCl₂ and CO₂. The radical O atoms, produced by the dissociation
115 of CO₂ or N₂O, react with their parent molecules and an isotopic exchange between O and CO₂ or
116 between O and N₂O can take place. These isotopic exchanges are mediated by the transient
117 formation of the activated complexes [CO₃]* or [N₂O₂]*. For magnesium isotopes we studied the
118 reaction between Pentanol and MgCl₂ for which the isotopic exchange can take place between Mg
119 and MgCl₂ through the transient formation of the complex [Mg₂Cl₂]*. The purpose of the introduction
120 of Pentanol along with the O and Mg gaseous carriers, is to produce carbonaceous C_xH_y[•]
121 radicals^(23,24) (dot designates radicals) that can react with the activated complexes before their
122 dissociation as O+CO₂, O+N₂O or Mg+MgCl₂. The reaction of the activated complex with a
123 carbonaceous radical should lead to the retrieval of O and Mg from the gas by condensation via
124 the polymerization of organic C_xH_y[•]-O or C_xH_y[•]-Mg macromolecules. A MIF effect is predicted in
125 such a case, when an activated complex is involved in two chemical reactions (here dissociation or
126 reaction with C_xH_y[•] radicals) at the same time because the rate of a chemical reaction involving a
127 complex formed by identical isotopes is not the same than that of a complex formed by non-identical
128 isotopes⁽²⁵⁾. Note that there is no scientific consensus on this interpretation⁽²⁶⁾.

129 **Results**

130
131 The Mg and O isotope compositions were measured using NanoSIMS at the Museum National
132 d'Histoire Naturelle in Paris (cf. Table S2 for instrumental setup). A selection of the data is reported

133 in Table S3 Tables and Figures. NanoSIMS analyses of the majority of the plasma deposited
134 carbonaceous material (PDCM) exhibit small isotopic variability, on the order of 10-30‰, likely due
135 to analytical effects caused by surface sample roughness. In the following, PDCM will refer to this
136 large fraction of condensed organic macromolecules that does not show any significant MIF effect.
137 The measured isotopic variations are thus expressed in δ units using the average isotopic
138 compositions (for O and Mg) of the PDCM as reference values. This procedure allows to minimize
139 analytical matrix effects and to express the isotopic fractionation relative to the non-fractionated
140 compounds produced from the same gas. Analyses of terrestrial samples with no MIF (San Carlos
141 olivine and terrestrial kerogen) and of the PDCM show an external reproducibility (measured in a
142 $10 \mu\text{m}^2$ area) of $\pm\approx 20\%$ (2σ) for $\delta^{25-26}\text{Mg}$ and $\pm\approx 30\%$ (2σ) for $\delta^{17-18}\text{O}$ (cf. Table S3.1, Figure S3.1
143 and Table S3.3 Figure S3.3).

144 The Mg and O isotopic results (Figures 1 and 2 constructed from Tables S3.2, S3.4 and S3.5)
145 show four major features. (i) Large MIFs distributing along a line of slope close to 1 are observed
146 both for O and Mg in the experimental condensates. For Mg (Figure 1a) the observed slope
147 (1.127 ± 0.034) is close to the calculated one (0.98 if mass dependent effects are ignored cf. S4)
148 using the model developed for MIF in ozone^(18,25). Similarly, for O, the observed slopes (1.086
149 ± 0.048 for $\delta^{17,18}\text{O} < 0\%$ and 1.066 ± 0.079 for $\delta^{17,18}\text{O} > 0\%$, ignoring data with $\delta^{18}\text{O} > 250\%$) are close
150 to the calculated one (1.00). Note that no free parameter is involved in the calculation of these
151 slopes. (ii) MIFs appear specific to chemical reactions. In Figure 2a, the negative $\delta^{17}\text{O}-\delta^{18}\text{O}$ data
152 correspond to the oxidation of Pentanol by N_2O (see also Figure S5), while the positive ones
153 correspond to the oxidation of the Pentanol/ MgCl_2 solution by CO_2 . In the Pentanol/ $\text{MgCl}_2/\text{CO}_2$
154 experiment, one area on the Si wafer shows variations distributed along a slope ≈ 0.5 but with an
155 enormous range of variations in $\delta^{18}\text{O}$ (from $\approx +700$ to $\approx +1500\%$; Figure 2a) impossible to explain
156 by a "classical MDF" (iii) Oxygen isotopic variations along the 1:1 CL are not correlated to the ^{16}O
157 concentration (i.e. to the ion count rate; cf. Table S3.5), so that they cannot be explained by mixing
158 between 2 components but requires instead variations in the magnitude of the MIF factor producing
159 multiple end-members. (iv) At the scale of 200-500 nm, MIFs are not systematically associated with
160 elemental O hotspots, i.e. with an increase in the ionic emissivity of O compared to that from the
161 amorphous PDCM where no MIF effect is observed. Note that, in the case of the Pentanol/ N_2O
162 experiment, no hotspot in emissivity was observed at the location of negative MIF effects.

163 As reported in Figure 1 in the case of the Pentanol/ MgCl_2 experiments, isotopic Mg hotspots
164 were clearly associated with the rim of carbonaceous grains (cf.⁽²⁷⁾ for the formation of nanoparticles
165 from plasma). As observed for Ti isotopes⁽¹⁸⁾, the Mg MIF effect varies within the grains in such a
166 way that the isotopic mass balance is achieved at the scale of the bulk grain. For experiments
167 involving oxygen isotopes, the morphology of the grains carrying the MIF effect could not be

168 determined and it cannot be excluded that the MIF carriers are part of the molecular structure of
169 the PDCM.

170

171

172 Discussion

173

174 Several theoretical origins of the MIF effect (i.e. the origin of the MIF factor noted η) have been
175 proposed^(26, 28-30). Here we take for granted that the MIF effect appears when the activated complex
176 responsible for the isotopic exchange is involved in a chemical reaction^(6, 16, 25).

177 For the CO₂/MgCl₂/Pentanol experiment, the complex is likely [CO₃]^{*}⁽³¹⁾. Because (i) no oxygen
178 MIF effect is observed in the Pentanol/CO₂ experiment (cf. Table S1) and because (ii) at the
179 location of oxygen MIF, chlorine is observed by SEM as ≈200 nm spots while Mg is absent, the MIF
180 effect is attributed to the reaction of [CO₃]^{*} with an organochlorine radical (C_xH_y-Cl)^{*}. The complex
181 [CO₃]^{*} can either decompose and return to the gas (Channel 1 of the following reaction 1) or be
182 stabilized (Channel 2) as (C_xH_y-ClO). In this model both Channels lead to a MIF effect with
183 fractionation opposite in sign, but only the isotopic effect in the condensed phase can be preserved
184 in the reaction products since isotopic re-equilibration will take place in the gas.

185

Reaction 1.

186 The lack of measurable MIF effects in the PDCM indicates that C_xH_y-O can also directly
187 condense without reacting with the complex [CO₃]^{*}. We thus now distinguish between two types of
188 condensation: a complex-mediated condensation and a direct condensation (i.e. a two-body
189 reaction). A similar scheme can be proposed to explain the Pentanol/N₂O experiment where the
190 complex would be [N₂O₂]^{*}.

191 The present experimental observations can be tentatively extended to the origin of oxygen
192 isotopic variations in the PSN. A possible scenario - alternative to the self-shielding ones⁽¹²⁻¹⁴⁾ - can
193 be proposed (Reaction 2; Figures. 3a, b):

194

Reaction 2

195 In this scenario, E is a chemical element (E ≡ Si, Al, Mg, Ca etc..) forming an oxide EO that
196 condenses from the gas phase. O_{Ads} and EO_{Ads} correspond to species adsorbed at the surface of
197 a solid (subscript Ads in Reaction 2) during its condensation, so that the formation rate of [EO-
198 O]_{Ads}^{*} is enhanced by many orders of magnitude relative to the rate in gas phase reaction. Note
199 that, contrary to the self-shielding model, the dissociation of H₂O_{Gas} or CO_{Gas} does not produce MIF
200 fractionated oxygen atoms (O_{Gas}). The oxygen isotope exchange takes place in [EO-O]_{Ads}^{*} and
201 produce the MIF effect during its decomposition. The Channel 1 corresponds to the stabilization of
202 a fraction of the complex [EO-O]_{Ads}^{*} in the solid (subscript Sol) by unknown reactions. The dashed

203 arrows for the Channel 2 indicate that the other fraction of the complex $[\text{EO-O}]_{\text{Ads}}^*$ can be either
204 incorporated in the solid (Channel 2a) as observed here for Mg and for $\text{Ti}^{(32)}$ isotopes or can
205 spontaneously dissociate to the gas phase. (Channel 2b). The O returned to the gas by this
206 destabilization of a fraction of the complex, re-equilibrates isotopically with the major oxygen
207 bearing molecules of the PSN (H_2O and CO). The magnitude and the sign of MIF in the two fractions
208 of the complex stabilized in the solid are expected to depend on the type of chemical reaction taking
209 place, as observed in the present experiments with oxidation of pentanol by either N_2O or CO_2
210 producing either ^{16}O excess or ^{16}O depletion, respectively (Figure 2). In other words, in the PSN, if
211 this process occurs, then different oxides should fractionate differently oxygen isotopes but,
212 possibly in a mass independent manner.

213 This catalytic property of grain surfaces for enhancing the rate of reaction between adsorbed
214 species, overcomes one of the major difficulties for having MIF effects in the PSN. Indeed, in the
215 PSN, in a large excess of H_2 , the O atoms produced by the photodissociation of H_2O or $\text{CO}^{(33)}$ are
216 likely to react with H_2 to form $\text{OH} + \text{H}$ much faster than being involved in a gaseous three body
217 reaction similar to that which leads to ozone⁽¹⁶⁾. However, the chemical nature of the adsorbed
218 activated complex taking part to condensation reactions in the PSN is, at this stage, impossible to
219 predict. This is exemplified by the fact that no MIF is observed in meteorites for Mg, while large
220 MIFs are produced in our experiments. This may indicate that the possible activated complexes
221 involving two Mg atoms are unstable (such as $[\text{Mg}_2\text{O}]^*$) in the chemical conditions that prevailed
222 during the condensation of Mg-rich silicates (forsterite could for instance form by direct
223 condensation and not by a complex-mediated condensation).

224 Thus, assuming that the gas from which the first oxides and silicates condense in the PSN has
225 the isotopic composition measured for the Solar wind ($\delta^{17}\text{O} \approx \delta^{18}\text{O} \approx -60\text{‰}$)^(4,5), only the grains
226 condensed by a complex-mediated condensation would develop MIF effects with $\delta^{17}\text{O} \approx \delta^{18}\text{O}$ either
227 $> -60\text{‰}$ ⁽³⁴⁾ or $< -60\text{‰}$ ^(35,36). All the other condensates, formed by direct condensation, would have
228 the oxygen isotopic composition of the PSN (Figure 3b). Note however that, due to mass balance
229 effects, a gaseous reservoir previously fractionated by complex-mediated condensation could
230 produce, by direct condensation, solids with MIF effects.

231 The simultaneous occurrence of these two types of condensation reactions, having different
232 consequences as far as oxygen isotopes are concerned, may account for part of the large diversity
233 of isotopic compositions observed in CAIs and chondrules from primitive meteorites⁽³⁴⁾. Note
234 however that the large $\delta^{17}\text{O} \approx \delta^{18}\text{O}$ variations observed in the present experiments (several hundreds
235 of ‰) compared with the much smaller enrichments in $^{17,18}\text{O}$ relative to the Sun observed in most
236 solar system solids (50-70‰) implies that their oxygen isotopic composition contains only a small
237 MIF contribution.

238 As postulated in the Reaction 2 and observed during the formation of $\text{SiO}_2^{(17)}$, complexes⁽³⁷⁾
239 resulting from the dissociation of H_2O could be responsible for the extreme ^{16}O depletion ($\delta^{17,18}\text{O}$
240 (‰) up to +180‰) observed⁽³⁸⁾ in some rare minerals whose synthesis involves the reaction
241 between Fe (or FeS) and $\text{H}_2\text{O}^{(38)}$ (cf. Figure3b).

242 Although the self-shielding model remains a viable possibility in the PSN to produce reservoirs
243 variously depleted or enriched in $^{16}\text{O}^{(12, 13)}$, the presently proposed MIF constitutes an additional
244 potential source of ^{16}O enrichments and depletion. This scenario is testable in laboratory through
245 dedicated experiments of high temperature oxide condensation.

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250 **Materials and Methods**

251

252

253 1- Experimental protocol.

254

255 All experiments (cf. the schematic drawing in Figure 4) were performed in a glass line where the
256 pressure reaches 10^{-4} Torr. A gas vector (N_2 for the Mg experiment and N_2O or CO_2 for the oxygen
257 experiments) is injected into the line and passes through a cylinder glass tube ($\varnothing = 1$ cm, $l = 10$ cm)
258 where the pressure is maintained at ≈ 1 Torr by dynamical pumping. Glass reservoirs containing
259 organic liquids (Pentanol or $MgCl_2$ dissolved in Pentanol) were connected to the line by stopcocks.
260 The vapor pressure of the liquids was injected in the gas vector by adjusting the leak of these
261 stopcocks. The gas to organic pressure ratio was of the order of 1:1. The characteristic lifetime of
262 the gas in the tube was 1/10 sec. The plasma was produced in the center of the glass tube by an
263 HF discharge at 2450 MHz. In such conditions, the molecular temperature is estimated to be around
264 1000K. The plasma is not thermal i.e. the temperature of the electrons is higher than that of the
265 molecules.

266 Silicon wafers were deposited 5 cm downhill from the discharge. Carbonaceous matter was
267 deposited on the glass walls around the discharge and on the silicon wafers. It was mechanically
268 collected as a powder for isotopic analyses. The molecular organic structure of this carbonaceous
269 matter is described elsewhere⁽³⁹⁾. Except for experiments involving N_2O , MIF effects were found on
270 the wafers. Except for Mg, no structural grain corresponding to the size and distribution of isotopic
271 hotspots was identified by SEM.

272 We report in this paper only experiments where MIF effects were identified (other experiments
273 are listed in Table S1). It should be kept in mind that MIF-bearing grains are rare in the organic
274 matter (named PDCM for Plasma Deposited Carbonaceous Matter for the organic matter with no
275 significant MIF). So, the fact that no MIF region was observed in these samples is not definitive
276 proof that they are absent.

277 2. Analytical protocol.

278 The Mg and O isotope compositions were measured using NanoSIMS at the Museum National
279 d'Histoire Naturelle in Paris. The isotopic compositions of individual grains (\varnothing 0.5 to 2 μm) were
280 analyzed by image analysis using the *Image* software package (L. Nittler, Carnegie Institution of
281 Washington, Washington, DC). For NanoSIMS analyses, carbonaceous powder collected on the
282 glass walls (region 4 in Figure 4) were pressed on gold, gold coated and mounted on 1 inch holders.
283 Silicon wafers are gold coated without any further preparation. The instrumental parameters of the
284 NanoSims are reported in Table S2.

285 In order to have comparable topographic and matrix effects for the isotopic reference and for
286 the samples, all data are reported using the PDCM as the reference value. In others terms, the
287 isotopic fractionation is expressed relative to the carbonaceous deposits that do not show MIF
288 effects in excess of $\approx \pm 2\sigma$ relative to their average value. When analyses were performed randomly,
289 each area ($\approx 20 \times 20 \mu\text{m}$) was divided in 9 regions (i.e. ROI $\approx 6 \times 6 \mu\text{m}$; noted “*Random Area*” in
290 Tables S3.3 and S3.5). Data reported in Tables S3.2, 3.4, and 3.5 are used to construct the Figures
291 in the text. Two examples of the distribution of the Mg and O data recorded on the PDCM are shown
292 in Figures S3.1 and S3.3.

293 Analyses on terrestrial kerogen powders give reproducibility of ± 37 and ± 23 on $\delta^{17}\text{O}$ (‰) and
294 $\delta^{18}\text{O}$ (‰), respectively (2σ), similar to those measured for the PDCM. Under the same conditions,
295 the reproducibility on the polished San Carlos olivine standard is $\approx \pm 5$ and $\approx \pm 8$ on $\delta^{17}\text{O}$ (‰) and
296 $\delta^{18}\text{O}$ (‰), respectively. The poor reproducibility on the PDCM is caused by the topography of the
297 organic powder pressed in gold, bearing in mind that the presence of non-detected MIF bearing
298 grains cannot be totally excluded.

299 For magnesium, a synthetic sample (Methylmagnesium chloride evaporated in the air) was used
300 to verify the MDF relations between the electron multipliers. For oxygen, these relations are
301 checked routinely for rock sample analyses.

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303
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307 F.R. and M.C. write the paper and interpretation. S.M. and A.G. analyzed the isotopic
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311 **References**

312

313

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- 410

411

412 Figure Legends

413 **Figure 1:** (a) Magnesium isotopic compositions of Mg-bearing carbonaceous grains reported as
414 $\delta^{25}\text{Mg}$ vs. $\delta^{26}\text{Mg}$ variations relative to the PDCM ($0:0 \pm \approx 20\%$). Data were collected in 4 grains either
415 as surface or volume variations (cf. Table S3.2). Error bars include statistical errors (ion counting
416 statistic) and the reproducibility on the standard ($\pm 2\sigma$). The 1:1 dashed line is drawn for reference
417 but does not stand for the best fit line to the data.

418 (b) $\delta^{26}\text{Mg}$ and $\delta^{25}\text{Mg}$ variations of the rim of the 5 μm size grain shown in Figure 1c are reported
419 as a function of the analytical sputtering time expressed in sec (data also reported in Figure 1a).
420 The total sputtering duration (6600 sec) indicates that the rim of the grain is ≈ 200 nm thick. The
421 PDCM ($\delta^{26}\text{Mg} \approx \delta^{25}\text{Mg} \approx 0\%$) embedding the grain reappears after complete sputtering. Note the
422 two-outlier data (which are not analytical errors) observed during the flipping from negative to
423 positive δ values (also obvious in Figure 1a). Cf. sample *Robert-Juillet-2019_23.im* in Table S3.2.

424 (c) Ionic image of the distribution of $\delta^{26}\text{Mg}$ shown in Figure 1b. This image is made by the
425 summation of the 3 scans collected between 3600 and 4800 sec (cf. Figure 1b). The core of the
426 grain (in blue) has a homogeneous isotopic composition ($\delta^{26}\text{Mg} \approx \delta^{25}\text{Mg} \approx 0\%$) while the rim exhibits
427 marked enrichments in ^{26}Mg (^{25}Mg not reported here) localized in spots that do not exceed 200 nm
428 in size. The PDCM where the $\delta^{25}\text{Mg}$ and $\delta^{26}\text{Mg}$ cannot be defined because of the too low counting
429 rates on ^{25}Mg and ^{26}Mg appears in black.

430 (d) SEM image of another Mg-bearing carbonaceous grains (not analyzed for isotopic analyses)
431 deposited on the silicon wafers. The structure suggests that the extreme isotopic variations
432 observed in the grain shown in (c) are concentrated in the fine-grained rim.

433

434

435 **Figure 2:** Figure 2a: Oxygen isotopic compositions of grains embedded in the PDCM are
436 reported as $\delta^{17}\text{O}$ vs. $\delta^{18}\text{O}$. The δ variations are expressed relative to PDCM ($0-0\% \pm \approx 30\%$; 2σ).
437 Error bars include statistical errors (ion counting statistic) and the reproducibility on the standard
438 ($\pm 2\sigma$). Open diamonds stand for the oxidation of Pentanol by N_2O . Open and black dots stand for
439 the oxidation of the MgCl_2 /Pentanol solution by CO_2 (open dots for grains concentrated in the same
440 $100 \times 100 \mu\text{m}$ area and defining a slope ≈ 0.5). The 1:1 dashed line is drawn for reference, not for
441 the best fit line. The correlation defined by the black dots does not intercept at $0-0\%$; this can be
442 an analytical matrix effect or a MDF contribution to the MIF effect.

443 Figure 2b: NanoSims ion image ($20 \times 20 \mu\text{m}$; sample *Robert-Juillet-2020_9* in Table S3.5)
444 showing four hotspots in oxygen intensities where MIF compositions were measured and reported
445 in Figure S6 (PDCM in black). The measured $\delta^{17}\text{O}$ are reported on the figure for 3 of the 4 hotspots.
446 In order to increase the contrast of the image the colors vary according to the logarithmic intensity
447 of the count rate (Cps for counts/sec).

448 Figure 2c. NanoSims ion image ($20 \times 20 \mu\text{m}$; sample *Robert-Juillet-2020_18* in Table S3.5)
449 showing isotopic hotspots on Si wafer (in $\delta^{18}\text{O}(\%)$ units).

450

451 **Figure 3:** (a) $\delta^{17}\text{O}$ vs. $\delta^{18}\text{O}$ diagram for the solar system. Selection of observational constraints.
452 The MDF line (slope 0.52) is shown for the Earth-Moon system (dashed line). Minerals from
453 refractory inclusions (CAIs), ameboid olivine aggregates (AOAs) and chondrules in carbonaceous
454 chondrites define the 1:1 CL. Sun: $\delta^{17,18}\text{O} \approx -60\%$ ⁽⁵⁾. Some rare CAIs, relict CAIs in chondrules and

455 chondrule from CH-CB chondrites are enriched in ^{16}O relative to the Sun^(35,36). Extreme ^{16}O
456 depletion was observed in the C3-Acfer 094 meteorite⁽³⁸⁾. Most meteorites (Chondrites H, L and
457 LL), the Earth, Moon and Mars have bulk $\delta^{17,18}\text{O}$ values around $0\pm 10\%$.

458 (b) Qualitative interpretation of (a). The PSN gas has the oxygen isotopic composition of the
459 Sun: $\delta^{17,18}\text{O} \approx -60\%$. The position of the grains on the 1:1 line is, at first order, dictated by the
460 chemical reaction responsible for the complex-mediated condensation of oxides and silicates
461 (noted $[\text{EO}_2]^*$ with E designating Si, Mg, Al etc.). Oxides would exhibit either ^{16}O depletion or
462 enrichment relative to the Sun. The extreme ^{16}O depletion observed in the C3-Acfer 094 meteorite
463 could result from a reaction involving one of the $[\text{O-H}_x\text{O}]^*$ complexes resulting from the dissociation
464 of H_2O ⁽³⁷⁾. Condensation taking place without going through an activated complex (noted direct
465 condensation) should yield grains having the oxygen isotopic composition of the ambient gas,
466 which can be different from the Sun if this gas was previously fractionated by complex-mediated
467 condensation.
468

469 **Figure 4:** (1) gas vector (N_2 or N_2O or CO_2) (2) reservoir of organic liquids (Pentanol or Pentanol-
470 MgCl_2) (3) leak of (2) in (1) (4) Microwave cavity (5) Silicon Wafer. Organic deposits are present on
471 the glass walls of region 4 and on the Wafer.

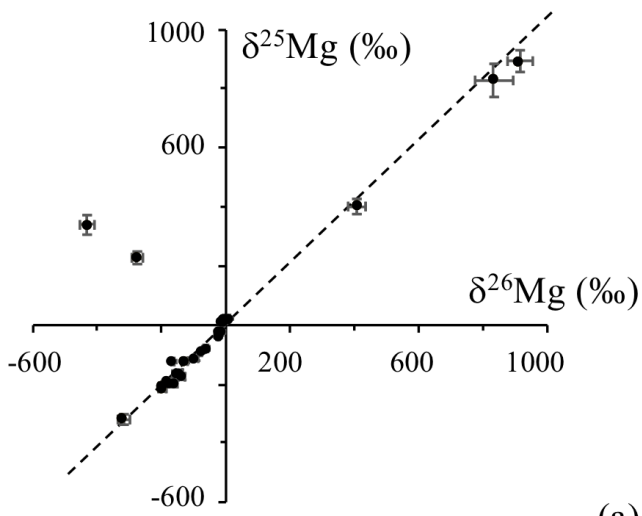
472 **Reaction Legends**

473 **Reaction 1:** Possible reaction at the origin of the oxygen MIF effect observed in the
474 $\text{CO}_2/\text{MgCl}_2/\text{Pentanol}$ plasma experiment.

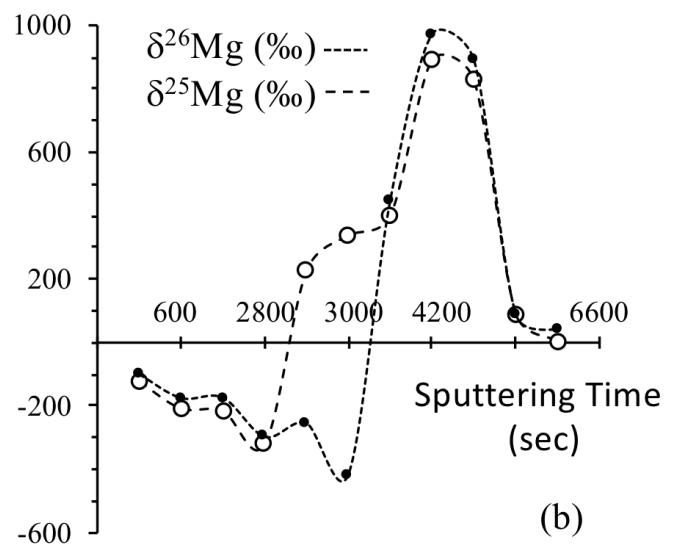
475 **Reaction 2:** Proposed chemical scheme at the origin of the mass independent oxygen isotopic
476 reaction observed in solar system materials.

477

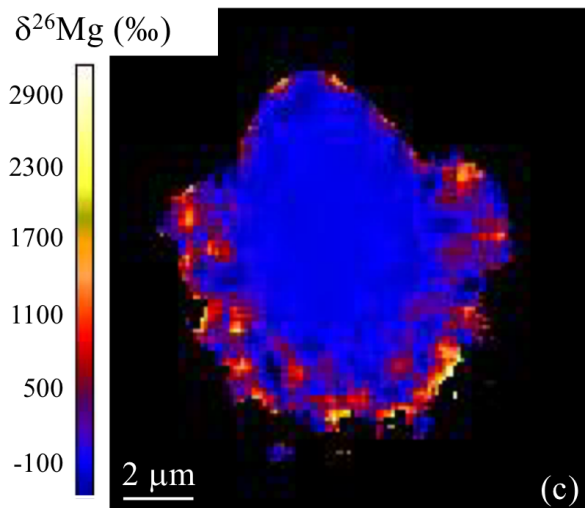
478



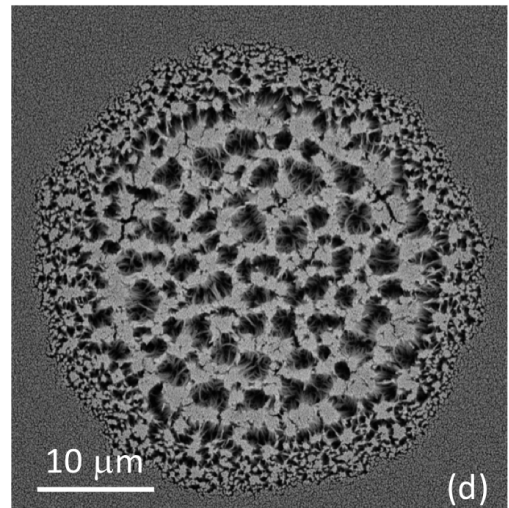
(a)



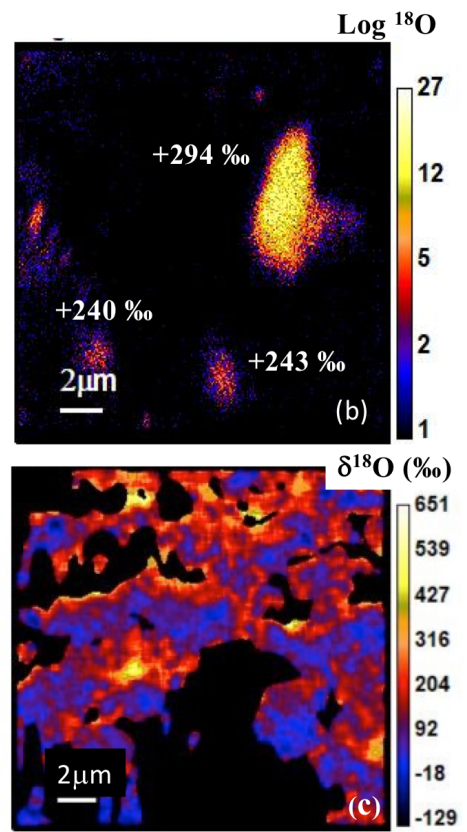
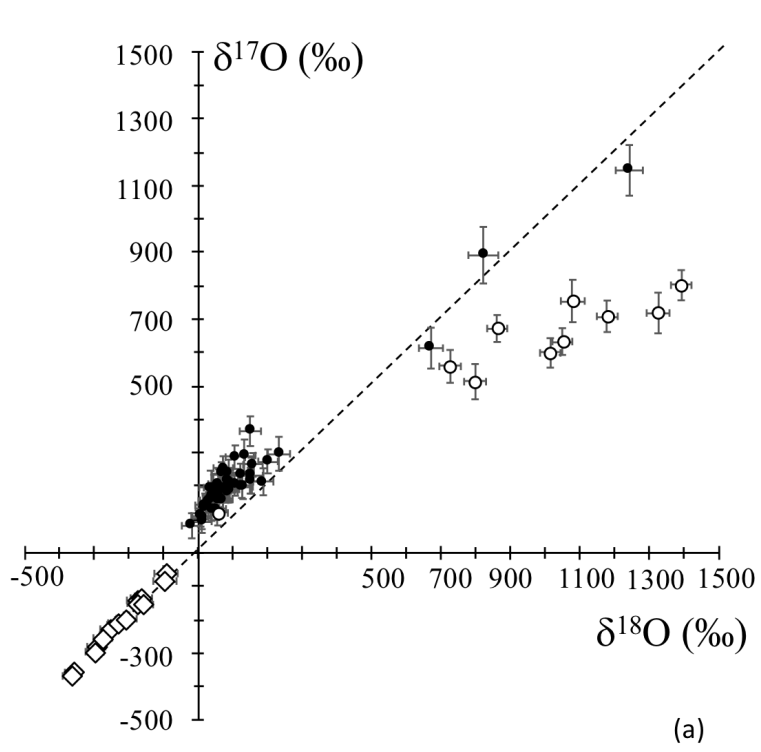
(b)

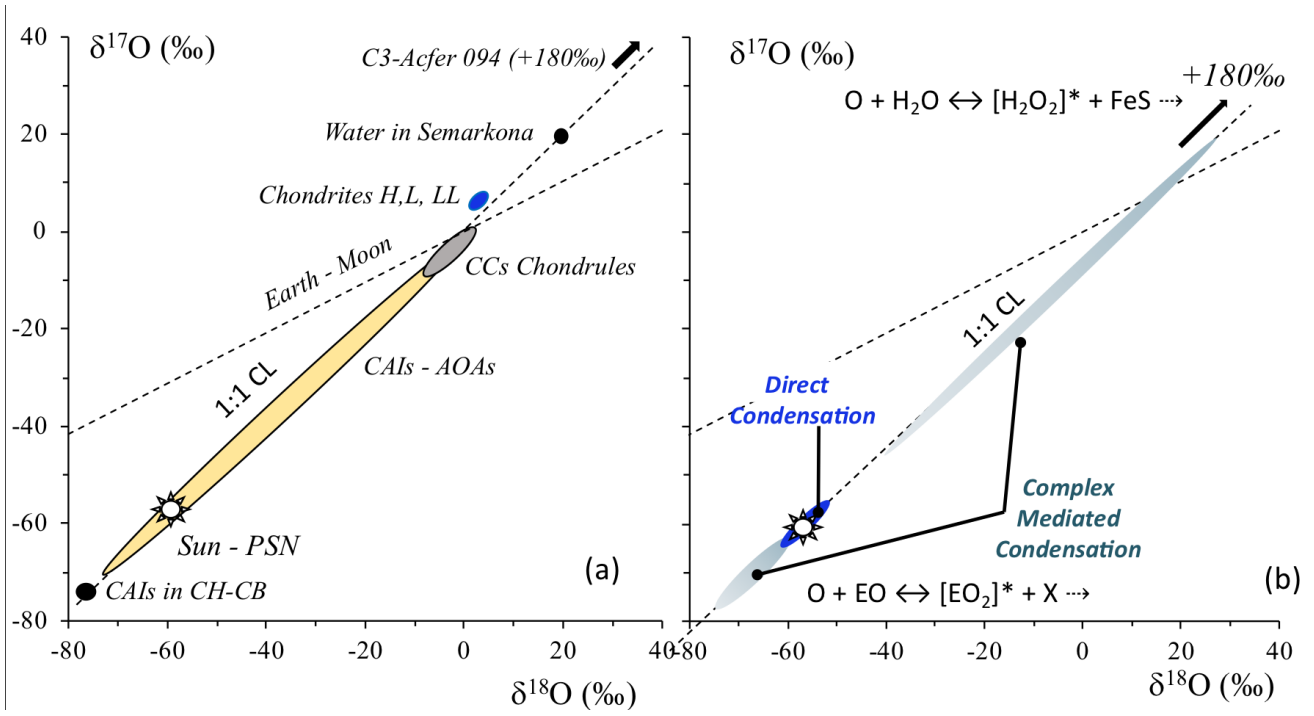


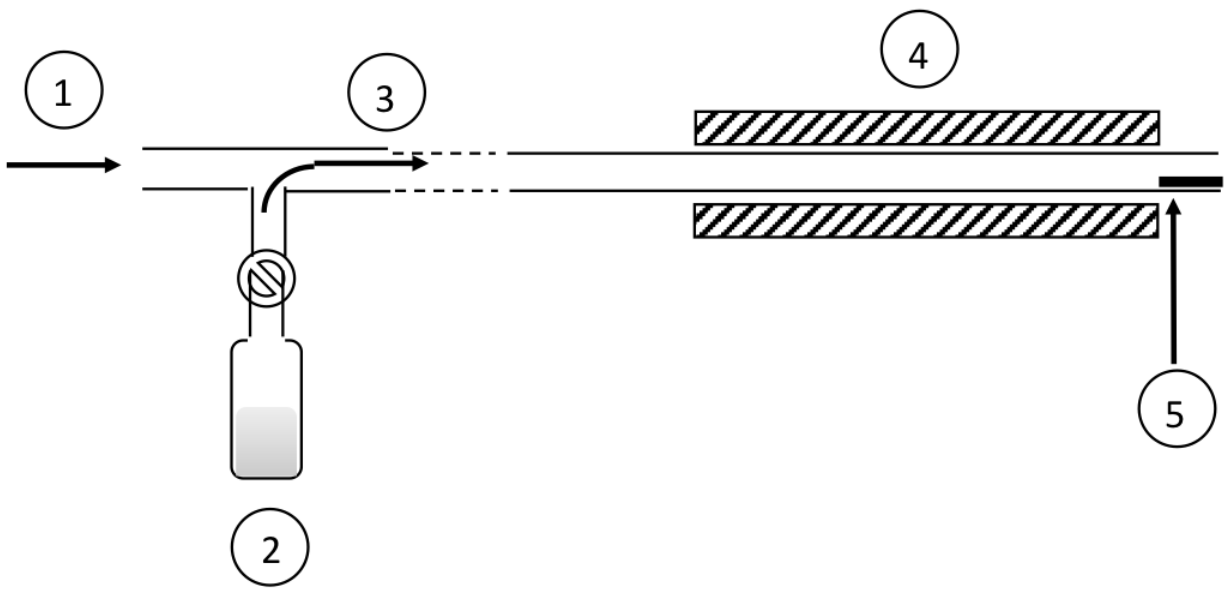
(c)

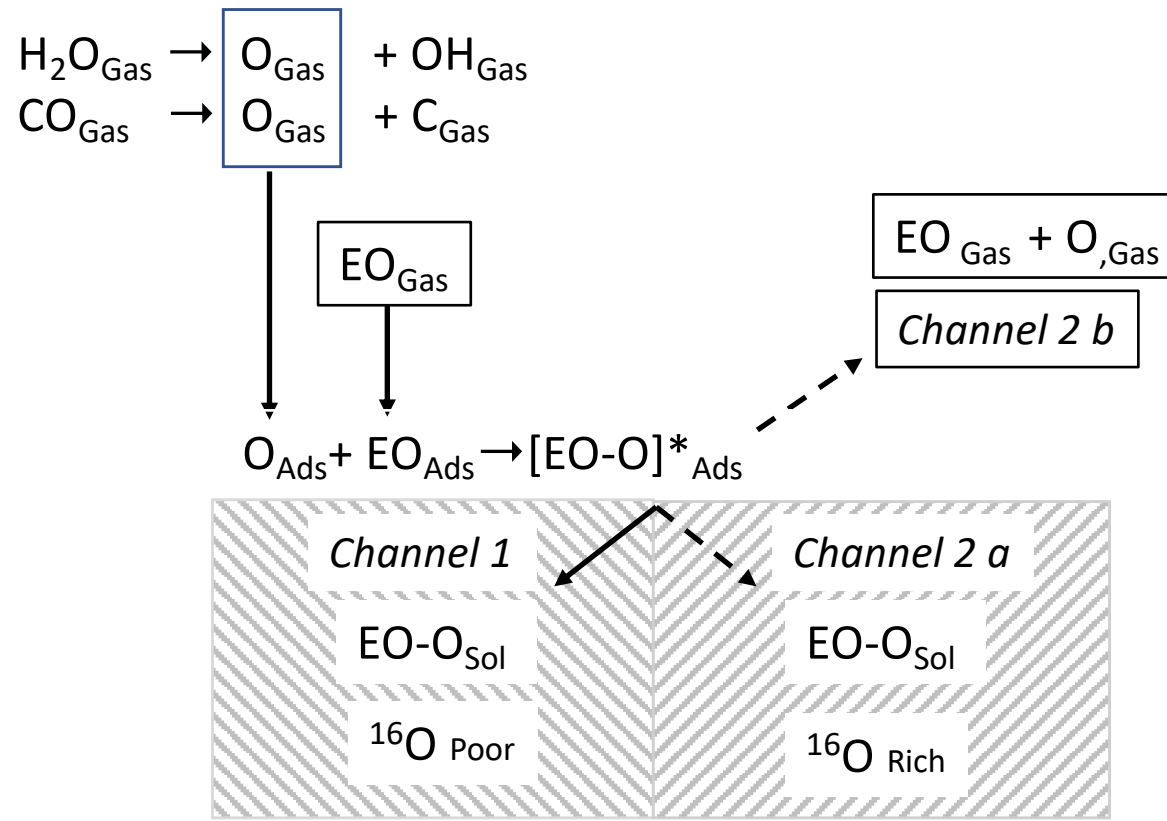


(d)











Oxygen and Magnesium Mass Independent Isotopic Fractionation
Induced by Chemical Reactions in Plasma

Supplementary Materials

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S1: MIF effects in different experiments.

E	Chemical compounds	Inorganic E	Organic E (hetero-element)	Chemical reactant	Grains ⁽³⁾	Hotspots ⁽⁴⁾	MIF
Mg	Magnesium dichloride in Pentanol	MgCl ₂	---	C ₅ H ₁₁ OH / N ₂	Yes	Yes	Yes
Mg	Methylmagnesium chloride in THF ⁽¹⁾	---	CH ₃ MgCl	N ₂	No	No	No
Mg	Methylmagnesium chloride in THF ⁽¹⁾	N ₂ O	CH ₃ MgCl	---	No	No	No
O	Pentanol	N ₂ O	---	C ₅ H ₁₁ OH	No	No	Yes
O	Triacetoxy(vinyl)silane ⁽²⁾	---	(CH ₃ CO ₂) ₃ SiCH - CH ₂	N ₂	No	No	No
O	Triacetoxy(vinyl)silane ⁽²⁾	N ₂ O	(CH ₃ CO ₂) ₃ SiCH - CH ₂	---	No	No	No
O	Magnesium dichloride in Pentanol	N ₂ O	---	MgCl ₂ / C ₅ H ₁₁ OH	Yes	Yes	No
O	Magnesium dichloride in Pentanol	CO ₂	---	MgCl ₂ / C ₅ H ₁₁ OH	Yes	Yes	Yes

Notes: (1) Tetrahydrofuran. (2) Si has 3 oxygen bonds. (3) Optical grains on the silicon wafer (4) hospots designate an increase in the ionic intensity i.e. in a change in the ionic emissivity caused by a different chemical speciation of oxygen. These hotspots are revealed by the ion imaging mode of the NanoSims.

Table S1: The first column designates the element E (Mg or O) for which the isotopic compositions have been measured. In the last 3 columns, “Yes” indicates that grains (optical microscopy), hospots (NanoSims) or MIF were observed. No MIF effect was detected for Mg when using an organic bonded Mg-Cl instead of MgCl₂ in solution with Pentanol. MIF effects on O were observed for Pentanol or Pentanol/MgCl₂ in presence of N₂O and CO₂, respectively but not with Pentanol/MgCl₂ in presence of N₂O. No MIF effect was detected for O when using organic bonded O atoms. Optical grains are not systematically associated with MIF effects.

S2. NanoSims analytical protocols

Session	Magnesium	Oxygen	Oxygen
	March 2020	Nov. 2018	July 2020
Primary current	O ⁻	Cs ⁺	Cs ⁺
Beam Current (pA)	13	650	1.8
Diameter Aperture	D1-2	D1-1	D1-3
Presputtering time (min)	5	2	5
Presputtering current (pA)	2000	1500	180
Secondary currents	Mg ⁺	O ⁻	O ⁻
Eletron multipliers	^{24,25,26} Mg ⁺	^{17,18} O ⁻	^{16,17,18} O ⁻
Faraday cup		¹⁶ O ⁻	
Rastering	20x20 μm	No Image	20x20 μm
Pixels	256 x 256	-----	256 x 256
Dwell time (ms/pixel)	1000	-----	1000
Mass resolution	8000	8500	8000
Entrance slit	ES3	ES3	ES3
Aperure slit	AS2	AS3	AS2
Dead Time	44 ns	44 ns	44 ns
Reference for δ units	PDCM	PDCM	PDCM

Table S2: The instrumental parameters of the NanoSIMS.

S3.1

EM on 24Mg, 25Mg and 26Mg MgCl ₂ in Pentanol / N ₂			25Mg/24Mg Reference Std value		1234E-01		26Mg/24Mg Reference Std value		1.371E-01				
Sample	Comments	24Mg Tot.	24Mg cps	25Mg Tot.	25Mg cps	26Mg Tot.	26Mg cps	Delta 26Mg %	± (2 std.dev)	Delta 25Mg %	± (2 std.dev)		
Olivine San Carlos. Used as a standard : Delta 17, 18 = 0 ‰													
Rastering (2x2)	sple@1_1	Polished Sample	Random Area 1	1.15E+07	1.12E+05	1.43E+06	1.40E+04	1.57E+06	1.54E+04	-2.3	1.6	8.9	
Rastering (2x2)	sple@1_2	Polished Sample	Random Area 2	1.07E+07	1.05E+05	1.34E+06	1.31E+04	1.47E+06	1.43E+04	-4.5	1.7	8.8	
Rastering (2x2)	sple@1_3	Polished Sample	Random Area 3	1.10E+07	1.08E+05	1.37E+06	1.34E+04	1.51E+06	1.47E+04	-2.9	1.6	8.8	
Rastering (2x2)	sple@1_4	Polished Sample	Random Area 4	1.11E+07	1.09E+05	1.39E+06	1.35E+04	1.52E+06	1.48E+04	-4.5	1.6	9.8	
Rastering (2x2)	sple@1_5	Polished Sample	Random Area 5	1.11E+07	1.08E+05	1.37E+06	1.34E+04	1.51E+06	1.47E+04	-6.8	1.6	7.5	
Rastering (2x2)	sple@1_6	Polished Sample	Random Area 6	1.08E+07	1.06E+05	1.35E+06	1.32E+04	1.48E+06	1.44E+04	-3.6	1.6	10.6	
Rastering (2x2)	sple@1_7	Polished Sample	Random Area 7	1.11E+07	1.09E+05	1.38E+06	1.35E+04	1.52E+06	1.48E+04	-3.9	1.6	8.7	
Rastering (2x2)	sple@1_8	Polished Sample	Random Area 8	1.12E+07	1.10E+05	1.40E+06	1.36E+04	1.53E+06	1.49E+04	-7.7	1.6	5.9	
Rastering (2x2)	sple@1_9	Polished Sample	Random Area 9	1.09E+07	1.06E+05	1.35E+06	1.32E+04	1.47E+06	1.44E+04	-9.3	1.6	8.4	
Rastering (2x2)	sple@1_10	Polished Sample	Random Area 10	1.12E+07	1.09E+05	1.39E+06	1.35E+04	1.52E+06	1.48E+04	-8.6	1.6	5.3	
Rastering (2x2)	Robert-Juillet-2019_1	Polished Sample	Random Area 1	2.22E+06	1.11E+04	2.81E+05	1.40E+03	3.13E+05	1.56E+03	27.4	3.6	25.1	
Rastering (2x2)	Robert-Juillet-2019_2	Polished Sample	Random Area 2	2.26E+06	1.13E+04	2.86E+05	1.43E+03	3.16E+05	1.58E+03	19.8	3.6	24.3	
Rastering (2x2)	Robert-Juillet-2019_3	Polished Sample	Random Area 3	6.61E+06	3.30E+04	8.21E+05	4.11E+03	9.00E+05	4.50E+03	-6.2	2.1	7.4	
Rastering (2x2)	Robert-Juillet-2019_4	Polished Sample	Random Area 4	6.26E+06	3.13E+04	7.78E+05	3.89E+03	8.51E+05	4.25E+03	-9.4	2.2	7.1	
PDCM (used as a standard)													
FR20	Sample			24Mg cps	25Mg cps	26Mg cps	24Mg Tot.	25Mg Tot.	26Mg Tot.	Delta 26Mg %	± (2 std.dev)	Delta 25Mg %	± (2 std.dev)
Rastering (25x25)	Robert-April-2019_3.im	Hot-Spot 1 in Powder	Image 3	1.11E+04	1.37E+03	1.53E+03	3.32E+05	4.10E+04	4.57E+04	6.0	9.4	2.9	9.9
Rastering (25x25)	Robert-April-2019_4.im	Hot-Spot 1 in Powder	Image 4	4.55E+04	5.60E+03	6.16E+03	7.42E+05	9.14E+04	1.01E+05	-11.3	6.3	-1.8	6.6
Rastering (25x25)	Robert-April-2019_4.im	Hot-Spot 2 in Powder	Image 4	3.61E+04	4.50E+03	4.94E+03	3.07E+05	3.83E+04	4.20E+04	-1.5	9.8	11.4	10.2
Rastering (25x25)	Robert-April-2019_4.im	Hot-Spot 3 in Powder	Image 4	3.20E+04	3.97E+03	4.41E+03	2.80E+05	3.47E+04	3.85E+04	4.7	10.2	5.7	10.7
Rastering (25x25)	Robert-April-2019_4.im	Hot-Spot 4 in Powder	Image 4	2.54E+04	3.16E+03	3.44E+03	1.83E+05	2.27E+04	2.47E+04	-13.6	12.7	7.3	13.3
Rastering (25x25)	Robert-April-2019_4.im	Hot-Spot 5 in Powder	Image 4	1.55E+04	1.92E+03	2.11E+03	1.67E+05	2.08E+04	2.28E+04	-4.9	13.2	7.3	13.9
Rastering (25x25)	Robert-April-2019_5_mod.im	Hot-Spot 1 in Powder	Image 5	1.13E+05	1.40E+04	1.54E+04	1.78E+06	2.19E+05	2.41E+05	-12.5	4.1	-2.0	4.3
Rastering (25x25)	Robert-April-2019_6.im	Hot-Spot 1 in Powder	Image 6	1.13E+05	1.40E+04	1.54E+04	1.87E+06	2.31E+05	2.54E+05	-8.0	4.0	1.2	4.2
Rastering (25x25)	Robert-April-2019_6.im	Hot-Spot 2 in Powder	Image 6	6.98E+04	8.56E+03	9.50E+03	4.85E+05	5.95E+04	6.60E+04	-7.4	7.8	-6.3	8.2
Rastering (25x25)	Robert-April-2019_6.im	Hot-Spot 3 in Powder	Image 6	8.04E+04	9.87E+03	1.08E+04	1.51E+06	1.86E+05	2.03E+05	-21.2	4.4	-4.6	4.6
Rastering (25x25)	Robert-April-2019_7_mod.im	Hot-Spot 1 in Powder	Image 7	7.09E+04	8.70E+03	9.63E+03	1.99E+06	2.44E+05	2.70E+05	-9.8	3.8	-5.4	4.0
Rastering (25x25)	Robert-April-2019_7_mod.im	Hot-Spot 2 in Powder	Image 7	6.85E+04	8.38E+03	9.31E+03	8.59E+05	1.05E+05	1.17E+05	-8.5	5.9	-8.5	6.2
Rastering (25x25)	Robert-April-2019_7_mod.im	Hot-Spot 3 in Powder	Image 7	1.20E+05	1.46E+04	1.61E+04	2.86E+06	3.49E+05	3.84E+05	-21.8	3.2	-12.1	3.4
Rastering (25x25)	Robert-April-2019_7_mod.im	Hot-Spot 7 in Powder	Image 7	8.27E+04	1.01E+04	1.13E+04	1.13E+06	1.38E+05	1.55E+05	-4.1	5.1	-8.7	5.4
Rastering (25x25)	Robert-April-2019_7_mod.im	Hot-Spot 8 in Powder	Image 7	8.88E+04	1.08E+04	1.19E+04	2.22E+06	2.84E+05	3.12E+05	-20.7	3.6	-9.8	3.8
Rastering (25x25)	Robert-April-2019_8.im	Hot-Spot 1 on Si-Wafer	Image 8	6.19E+04	7.52E+03	8.41E+03	1.35E+06	1.64E+05	1.83E+05	-9.4	4.7	-15.9	4.9
Rastering (25x25)	Robert-April-2019_9.im	Hot-Spot 1 on Si-Wafer	Image 9	1.59E+05	1.97E+04	2.21E+04	1.44E+07	1.78E+06	2.00E+06	14.9	1.4	4.0	1.5
Rastering (25x25)	Robert-April-2019_11.im	Hot-Spot 1 on Si-Wafer	Image 11	3.39E+04	4.20E+03	4.64E+03	3.16E+06	3.91E+05	4.33E+05	-1.6	3.0	2.9	3.2
Rastering (25x25)	Robert-April-2019_11.im	Hot-Spot 2 on Si-Wafer	Image 11	1.99E+04	2.47E+03	2.75E+03	2.07E+06	2.56E+05	2.85E+05	5.8	3.7	4.3	3.9
Rastering (25x25)	Robert-April-2019_11.im	Hot-Spot 3 on Si-Wafer	Image 11	1.91E+04	2.38E+03	2.63E+03	2.74E+06	3.42E+05	3.78E+05	6.2	3.3	12.8	3.4
Rastering (25x25)	Robert-April-2019_11.im	Hot-Spot 4 on Si-Wafer	Image 11	8.36E+03	1.03E+03	1.13E+03	1.77E+05	2.18E+04	2.39E+04	-13.1	12.9	1.3	13.5

Table S3.1: Magnesium isotopic composition expressed as $\delta^{25}\text{Mg}(\text{‰})$ and $\delta^{26}\text{Mg}(\text{‰})$. The ionic counting rates are expressed in counts per second (cps) while the total number of counts (Tot) depend on the duration of the measurement. The average value of the PDCM is used as the *Reference Std. value* for calculating the δ values of the hotspots exhibiting marked departure in their isotopic composition. The ionic ratios $^{25}\text{Mg}/^{24}\text{Mg}$ and $^{26}\text{Mg}/^{24}\text{Mg}$ used to calculate the δ values are indicated as *Reference Std value*. Error bars are the statistical precision calculated on the total number of counts ($\pm 2\sigma$). Two examples of the statistical distribution of the data is shown for (1) an olivine sample (San Carlos) and (2) 1 region of the sample obtained in the MgCl₂ - Pentanol - N₂ experiment (noted PDCM) showing a statistical distribution of isotope ratios. The PDCM isotopic distribution is reported in Figure S3.1 bellow.

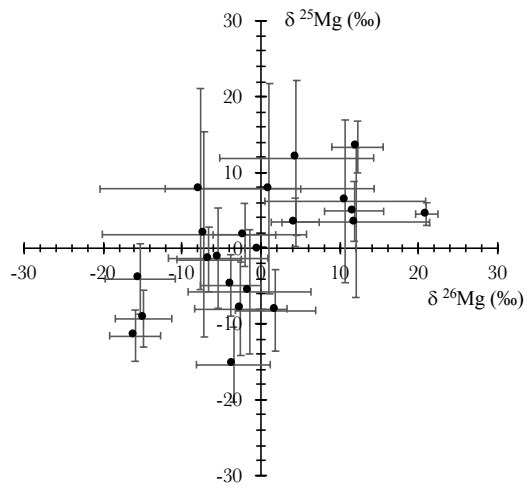


Figure S3.1: An example of the distribution of data recorded randomly on the PDCM is shown as $\delta^{25}\text{Mg}$ (‰) vs. $\delta^{26}\text{Mg}$ (‰). The reproducibility on this set of data is 15.6 and ± 19.8 (2σ) for $\delta^{25}\text{Mg}$ (‰) and $\delta^{26}\text{Mg}$ (‰), respectively. Error bars stand for the statistical analytical error.

Table S3.2:

25Mg/24Mg PDCM used as a reference 1.234E-01
 26Mg/24Mg PDCM used as a reference 1.371E-01

Sample	Comments	24Mg Tot.	24Mg cps	25Mg Tot.	25Mg cps	26Mg Tot.	26Mg cps	Delta 26 ‰	2 sigma ±	Delta 25 ‰	2 sigma ±	
FR20 - Individual grain on upstream SI Wafer												
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	6.47E+06	9.54E+02	7.25E+05	1.07E+02	8.14E+05	1.20E+02	-83.3	2.3	-92.5	2.2
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	8.56E+06	5.42E+04	1.08E+06	6.83E+03	1.18E+06	7.46E+03	3.1	1.9	21.5	1.8
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	1.27E+07	1.28E+05	1.60E+06	1.60E+04	1.74E+06	1.74E+04	-4.3	1.6	18.3	1.5
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	1.16E+07	2.26E+05	1.46E+06	2.84E+04	1.58E+06	3.08E+04	-6.4	1.7	19.8	1.6
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	1.81E+07	3.22E+05	2.27E+06	4.05E+04	2.46E+06	4.40E+04	-5.7	1.3	17.6	1.3
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	8.78E+06	3.82E+05	1.10E+06	4.80E+04	1.20E+06	5.21E+04	-5.6	1.9	19.4	1.8
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	6.66E+07	9.25E+03	8.29E+06	1.15E+03	9.03E+06	1.25E+03	-11.6	0.7	8.4	0.7
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	2.20E+06	3.54E+03	2.65E+05	4.27E+02	2.93E+05	4.71E+02	-30.2	3.7	-23.2	3.9
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	1.50E+06	2.41E+03	1.81E+05	2.91E+02	2.02E+05	3.24E+02	-20.1	4.5	-23.5	4.7
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	1.25E+06	2.02E+03	1.49E+05	2.39E+02	1.67E+05	2.68E+02	-30.1	4.9	-37.8	5.2
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	6.70E+05	1.08E+03	7.61E+04	1.22E+02	8.55E+04	1.38E+02	-68.3	6.8	-79.4	7.3
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	5.23E+05	8.40E+02	5.72E+04	9.20E+01	6.45E+04	1.04E+02	-99.8	7.9	-112.5	8.4
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	3.41E+05	5.49E+02	3.47E+04	5.58E+01	4.01E+04	6.45E+01	-142.9	10.0	-175.3	10.7
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	3.36E+05	5.40E+02	3.47E+04	5.58E+01	3.91E+04	6.29E+01	-150.7	10.1	-162.6	10.7
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	2.81E+05	4.52E+02	2.78E+04	4.46E+01	3.21E+04	5.15E+01	-167.7	11.2	-198.7	12.0
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	3.11E+05	4.99E+02	3.19E+04	5.13E+01	3.58E+04	5.75E+01	-160.5	10.6	-166.5	11.2
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	2.44E+05	3.92E+02	2.42E+04	3.89E+01	2.74E+04	4.40E+01	-181.1	12.1	-195.9	12.9
Rastering (25x25)	Robert-Juillet-2019_21.im	Image 21	2.40E+05	3.86E+02	2.39E+04	3.84E+01	2.69E+04	4.32E+01	-184.2	12.2	-194.4	12.9
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	1.83E+06	2.67E+02	1.97E+05	2.89E+01	2.08E+05	3.04E+01	-170.4	4.4	-123.8	4.5
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	4.35E+06	2.12E+04	5.74E+05	2.68E+03	6.26E+05	2.93E+03	4.0	2.5	22.0	2.6
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	1.13E+07	1.08E+05	1.41E+06	1.36E+04	1.54E+06	1.48E+04	-6.4	1.6	15.5	1.7
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	8.63E+06	2.21E+05	1.08E+06	2.76E+04	1.17E+06	2.99E+04	-11.7	1.8	13.5	1.9
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	4.66E+06	2.82E+05	5.84E+05	3.54E+04	6.32E+05	3.83E+04	-10.4	2.5	15.0	2.6
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	3.11E+07	4.32E+03	3.87E+06	5.37E+02	4.20E+06	5.82E+02	-16.6	1.0	7.6	1.0
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	2.66E+07	1.29E+05	3.33E+06	1.61E+04	3.62E+06	1.75E+04	-8.4	1.1	14.9	1.1
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	7.53E+05	1.21E+03	8.15E+04	1.31E+02	8.96E+04	1.44E+02	-131.6	7.0	-122.5	6.7
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	4.58E+05	7.37E+02	4.48E+04	7.21E+01	5.00E+04	8.05E+01	-203.6	9.4	-207.0	8.9
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	2.79E+05	4.49E+02	2.70E+04	4.35E+01	3.05E+04	4.91E+01	-202.6	12.2	-214.8	11.5
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	1.36E+05	2.18E+02	1.14E+04	1.83E+01	1.26E+04	2.03E+01	-321.7	18.8	-320.2	17.8
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	8.98E+04	1.45E+02	1.36E+04	2.19E+01	8.87E+03	1.43E+01	-280.2	17.1	-227.5	21.2
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	4.72E+04	7.60E+01	7.80E+03	1.26E+01	3.66E+03	5.89E+00	-435.5	22.6	-339.0	33.1
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	3.15E+04	5.07E+01	5.45E+03	8.77E+00	6.04E+03	9.72E+00	-398.9	27.1	-401.3	25.7
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	1.12E+04	1.81E+01	2.62E+03	4.22E+00	2.93E+03	4.71E+00	-903.9	39.1	-893.5	37.0
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	5.06E+03	8.14E+00	1.14E+03	1.84E+00	1.27E+03	2.04E+00	-823.6	59.2	-828.2	56.2
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	4.82E+03	7.76E+00	6.49E+02	1.04E+00	6.94E+02	1.12E+00	-50.5	78.5	91.9	75.9
Rastering (25x25)	Robert-Juillet-2019_23.im	Image 23	5.46E+04	4.49E+00	7.00E+03	5.76E-01	7.26E+03	5.98E-01	-29.3	23.5	40.0	23.9

Table S3.2: Magnesium isotopic composition expressed as $\delta^{25}\text{Mg}(\text{‰})$ and $\delta^{26}\text{Mg}(\text{‰})$. The plasma discharge is generated in a mixture MgCl_2 - Pentanol - N_2 . The ionic counting rates are expressed in counts per second (cps) while the total number of counts (Tot) depend on the duration of the measurement. Error bars are the statistical precision calculated on the total number of counts ($\pm 2\sigma$). The average value of the PDCM is used as the reference value (cf. previous Table S3.1) for calculating the δ values of the hotspots exhibiting marked departure in their isotopic composition. The average PDCM ionic ratios $^{25}\text{Mg}/^{24}\text{Mg}$ and $^{26}\text{Mg}/^{24}\text{Mg}$ are indicated as *Reference Std value*. Samples noted 2D and 3D indicates that the Mg isotopic compositions were measured (i) by image analysis as a cross section from the surface to the core of the hotspots (ii) by depth profiling until the complete disappearance of the hotspot caused by its sputtering. This Table was used to construct the Figure 1 of the text.

Faraday on 160, EM on 170 and 180
Pentanol / N2O

170/160 Reference value for delta 3.709E-04
180/160 Reference value for delta 1.972E-03

San Carlos Olivine.	Sample	Comments	160 Tot.	160 Cps	170 Tot.	170 Cps	180 Tot.	180 Cps	Delta 18 %	± 18 (2s) %	Delta 17 %	±17 (2s) %
Random Area 1	Robert-Oct-2019_1.im	Polished section	9.52E+07	1.16E+05	3.70E+04	4.52E+01	1.97E+05	2.41E+02	49.9	4.5	48.8	10.4
Random Area 3	Robert-Oct-2019_1.im	Polished section	8.41E+07	1.03E+05	3.30E+04	4.03E+01	1.74E+05	2.13E+02	51.1	4.8	57.9	11.0
Random Area 1	Robert-Oct-2019_2.im	Polished section	1.39E+08	1.05E+05	5.40E+04	4.05E+01	2.87E+05	2.16E+02	48.6	3.7	46.7	8.6
Random Area 2	Robert-Oct-2019_2.im	Polished section	1.15E+08	8.66E+04	4.52E+04	3.39E+01	2.37E+05	1.78E+02	47.3	4.1	58.5	9.4
Random Area 3	Robert-Oct-2019_2.im	Polished section	1.39E+08	1.04E+05	5.43E+04	4.08E+01	2.87E+05	2.16E+02	47.3	3.7	53.2	8.6
Random Area 4	Robert-Oct-2019_2.im	Polished section	1.32E+08	9.91E+04	5.18E+04	3.89E+01	2.74E+05	2.06E+02	52.9	3.8	57.5	8.8

D. Rumble's Kerogen	Sample	Comments	160 Tot.	160 Cps	170 Tot.	170 Cps	180 Tot.	180 Cps	Delta 18 %	± 18 (2s) %	Delta 17 %	±17 (2s) %
Random Area 1	Robert-Oct-2019_5.im	Organic Matter in Powder	2.20E+08	6.25E+04	8.62E+04	2.44E+01	4.55E+05	1.29E+02	48.0	6.8	56.4	3.0
Random Area 2	Robert-Oct-2019_5.im	Organic Matter in Powder	1.83E+08	5.17E+04	7.14E+04	2.02E+01	3.77E+05	1.07E+02	45.7	7.5	51.4	3.3
Random Area 3	Robert-Oct-2019_5.im	Organic Matter in Powder	2.40E+07	6.80E+03	9.32E+03	2.64E+00	4.88E+04	1.38E+01	32.2	20.7	46.8	9.0
Random Area 4	Robert-Oct-2019_5.im	Organic Matter in Powder	3.10E+08	8.80E+04	1.21E+05	3.43E+01	6.33E+05	1.79E+02	35.0	5.7	53.8	2.5
Random Area 5	Robert-Oct-2019_5.im	Organic Matter in Powder	1.47E+08	4.16E+04	5.76E+04	1.63E+01	3.03E+05	8.58E+01	43.9	8.3	56.3	3.6
Random Area 6	Robert-Oct-2019_5.im	Organic Matter in Powder	1.27E+08	3.61E+04	4.95E+04	1.40E+01	2.62E+05	7.44E+01	48.1	9.0	50.3	3.9
Random Area 7	Robert-Oct-2019_5.im	Organic Matter in Powder	2.30E+08	6.53E+04	9.02E+04	2.56E+01	4.75E+05	1.35E+02	47.3	6.7	57.5	2.9
Random Area 8	Robert-Oct-2019_5.im	Organic Matter in Powder	3.30E+07	9.34E+03	1.29E+04	3.65E+00	6.77E+04	1.92E+01	40.4	17.6	51.3	7.7
Random Area 9	Robert-Oct-2019_5.im	Organic Matter in Powder	1.80E+07	5.11E+03	7.04E+03	1.99E+00	3.69E+04	1.05E+01	40.9	23.8	53.9	10.4

FR9 PDCM (Used as Reference)	Sample	Comments	160 Tot.	160 Cps	170 Tot.	170 Cps	180 Tot.	180 Cps	Delta 18 %	± 18 (2s) %	Delta 17 %	±17 (2s) %
Rastering (15x15)	FR9@_1_1	Collected as a Powder	7.65E+08	2.34E+06	2.86E+05	8.72E+02	1.52E+06	4.65E+03	9.8	1.6	6.4	3.7
Rastering (15x15)	FR9@_1_2	Collected as a Powder	8.01E+08	2.44E+06	2.95E+05	8.99E+02	1.58E+06	4.82E+03	0.5	1.6	-7.9	3.7
Rastering (15x15)	FR9@_1_3	Collected as a Powder	7.78E+08	2.37E+06	2.89E+05	8.82E+02	1.53E+06	4.67E+03	-2.7	1.6	2.2	3.7
Rastering (15x15)	FR9@_1_4	Collected as a Powder	8.18E+08	2.50E+06	3.06E+05	9.32E+02	1.60E+06	4.89E+03	-5.9	1.6	6.6	3.6
Rastering (15x15)	FR9@_1_5	Collected as a Powder	7.47E+08	2.28E+06	2.80E+05	8.54E+02	1.51E+06	4.59E+03	22.7	1.6	10.2	3.8
Rastering (15x15)	FR9@_1_6	Collected as a Powder	7.87E+08	2.40E+06	2.92E+05	8.92E+02	1.57E+06	4.80E+03	12.9	1.6	0.4	3.7
Rastering (15x15)	FR9@_1_7	Collected as a Powder	7.66E+08	2.34E+06	2.83E+05	8.63E+02	1.54E+06	4.70E+03	20.5	1.6	-4.7	3.8
Rastering (15x15)	FR9@_1_8	Collected as a Powder	8.77E+08	2.68E+06	3.20E+05	9.76E+02	1.71E+06	5.23E+03	-8.9	1.5	-16.3	3.5
Rastering (15x15)	FR9@_1_9	Collected as a Powder	7.95E+08	2.43E+06	2.99E+05	9.12E+02	1.56E+06	4.75E+03	-7.1	1.6	14.1	3.7
Rastering (15x15)	FR9@_1_10	Collected as a Powder	8.16E+08	2.49E+06	3.04E+05	9.26E+02	1.63E+06	4.97E+03	11.5	1.6	3.0	3.6
Rastering (15x15)	FR9@_1_11	Collected as a Powder	8.13E+08	2.48E+06	3.05E+05	9.30E+02	1.64E+06	5.00E+03	21.9	1.6	10.7	3.6
Rastering (15x15)	FR9@_1_12	Collected as a Powder	8.33E+08	2.54E+06	3.05E+05	9.32E+02	1.65E+06	5.03E+03	2.8	1.6	-12.3	3.6
Rastering (15x15)	FR9@_1_13	Collected as a Powder	8.73E+08	2.66E+06	3.21E+05	9.79E+02	1.72E+06	5.24E+03	-2.5	1.5	-9.3	3.5
Rastering (15x15)	FR9@_1_14	Collected as a Powder	7.92E+08	2.42E+06	2.92E+05	8.91E+02	1.55E+06	4.73E+03	-6.4	1.6	-6.1	3.7
Rastering (15x15)	FR9@_1_15	Collected as a Powder	8.24E+08	2.52E+06	3.08E+05	9.39E+02	1.62E+06	4.94E+03	-4.7	1.6	6.9	3.6
Rastering (15x15)	FR9@_1_16	Collected as a Powder	8.19E+08	2.50E+06	3.04E+05	9.27E+02	1.60E+06	4.89E+03	-8.5	1.6	0.0	3.6
Rastering (15x15)	FR9@_1_17	Collected as a Powder	8.33E+08	2.54E+06	3.12E+05	9.51E+02	1.63E+06	4.99E+03	-5.1	1.6	8.4	3.6
Rastering (15x15)	FR9@_1_18	Collected as a Powder	8.27E+08	2.52E+06	3.10E+05	9.45E+02	1.62E+06	4.94E+03	-7.3	1.6	9.6	3.6
Rastering (15x15)	FR9@_1_19	Collected as a Powder	7.32E+08	2.23E+06	2.75E+05	8.38E+02	1.43E+06	4.35E+03	-11.7	1.7	12.1	3.5
Rastering (15x15)	FR9@_1_20	Collected as a Powder	8.87E+08	2.71E+06	3.28E+05	1.00E+03	1.78E+06	5.42E+03	16.6	1.5	-2.7	3.8
Rastering (15x15)	FR9@_1_21	Collected as a Powder	7.97E+08	2.43E+06	2.96E+05	9.04E+02	1.59E+06	4.87E+03	14.8	1.6	1.9	3.7
Rastering (15x15)	FR9@_1_22	Collected as a Powder	7.78E+08	2.37E+06	2.87E+05	8.75E+02	1.52E+06	4.64E+03	-8.6	1.6	-6.1	3.7
Rastering (15x15)	FR9@_1_23	Collected as a Powder	7.87E+08	2.40E+06	2.87E+05	8.77E+02	1.55E+06	4.72E+03	-3.7	1.6	-15.9	3.7
Rastering (15x15)	FR9@_1_24	Collected as a Powder	7.94E+08	2.42E+06	2.92E+05	8.90E+02	1.56E+06	4.75E+03	-5.0	1.6	-9.6	3.7
Rastering (15x15)	FR9@_1_25	Collected as a Powder	7.77E+08	2.37E+06	2.93E+05	8.93E+02	1.53E+06	4.67E+03	-0.6	1.6	15.6	3.7
Rastering (15x15)	FR9@_1_26	Collected as a Powder	8.33E+08	2.54E+06	3.12E+05	9.51E+02	1.66E+06	5.07E+03	12.0	1.6	8.8	3.6
Rastering (15x15)	FR9@_1_27	Collected as a Powder	8.27E+08	2.52E+06	3.08E+05	9.39E+02	1.64E+06	5.01E+03	7.5	1.6	3.0	3.6
Rastering (15x15)	FR9@_1_28	Collected as a Powder	7.71E+08	2.35E+06	2.81E+05	8.58E+02	1.52E+06	4.64E+03	-0.9	1.6	-17.2	3.8
Rastering (15x15)	FR9@_1_29	Collected as a Powder	8.27E+08	2.53E+06	3.10E+05	9.46E+02	1.65E+06	5.03E+03	10.9	1.6	9.8	3.6
Rastering (15x15)	FR9@_1_30	Collected as a Powder	8.18E+08	2.50E+06	3.03E+05	9.25E+02	1.64E+06	5.01E+03	18.4	1.6	-0.4	3.6
Rastering (15x15)	FR9@_1_31	Collected as a Powder	8.16E+08	2.49E+06	2.99E+05	9.13E+02	1.63E+06	4.98E+03	15.5	1.6	-11.0	3.7
Rastering (15x15)	FR9@_1_32	Collected as a Powder	8.16E+08	2.49E+06	3.04E+05	9.27E+02	1.63E+06	4.97E+03	11.8	1.6	3.2	3.6
Rastering (15x15)	FR9@_1_33	Collected as a Powder	7.81E+08	2.38E+06	2.93E+05	8.95E+02	1.54E+06	4.70E+03	-0.3	1.6	11.8	3.7
Rastering (15x15)	FR9@_1_34	Collected as a Powder	8.40E+08	2.56E+06	3.10E+05	9.46E+02	1.65E+06	5.02E+03	-6.1	1.6	-4.9	3.6
Rastering (15x15)	FR9@_1_35	Collected as a Powder	8.61E+08	2.63E+06	3.12E+05	9.52E+02	1.68E+06	5.14E+03	-8.5	1.5	-23.2	3.6
Rastering (15x15)	FR9@_1_36	Collected as a Powder	7.86E+08	2.40E+06	2.93E+05	8.94E+02	1.54E+06	4.71E+03	-4.5	1.6	5.1	3.7
Rastering (15x15)	FR9@_1_37	Collected as a Powder	8.03E+08	2.45E+06	3.00E+05	9.16E+02	1.58E+06	4.81E+03	-5.5	1.6	7.7	3.6
Rastering (15x15)	FR9@_1_38	Collected as a Powder	7.96E+08	2.43E+06	2.98E+05	9.11E+02	1.56E+06	4.77E+03	-4.6	1.6	10.5	3.7
Rastering (15x15)	FR9@_1_39	Collected as a Powder	7.99E+08	2.44E+06	2.95E+05	9.01E+02	1.56E+06	4.77E+03	-6.6	1.6	-3.8	3.7
Rastering (15x15)	FR9@_1_40	Collected as a Powder	8.00E+08	2.44E+06	2.90E+05	8.84E+02	1.56E+06	4.75E+03	-13.0	1.6	-24.3	3.7
Rastering (15x15)	FR9@_1_41	Collected as a Powder	7.53E+08	2.30E+06	2.81E+05	8.56E+02	1.46E+06	4.45E+03	-18.2	1.7	4.3	3.8
Rastering (15x15)	FR9@_1_42	Collected as a Powder	7.79E+08	2.38E+06	2.90E+05	8.87E+02	1.52E+06	4.63E+03	-11.3	1.6	5.4	3.7
Rastering (15x15)	FR9@_1_43	Collected as a Powder	7.58E+08	2.31E+06	2.81E+05	8.59E+02	1.48E+06	4.50E+03	-12.8	1.6	0.8	3.8
Rastering (15x15)	FR9@_1_44	Collected as a Powder	7.92E+08	2.42E+06	2.95E+05	9.00E+02	1.54E+06	4.69E+03	-16.0	1.6	3.6	3.7
Rastering (15x15)	FR9@_1_45	Collected as a Powder	7.85E+08	2.40E+06	2.92E+05	8.90E+02	1.53E+06	4.66E+03	-14.3	1.6	1.4	3.7
Rastering (15x15)	FR9@_1_46	Collected as a Powder	7.62E+08	2.32E+06	2.85E+05	8.70E+02	1.49E+06	4.54E+03	-10.2	1.6	8.5	3.7
Rastering (15x15)	FR9@_1_47	Collected as a Powder	7.44E+08	2.27E+06	2.75E+05	8.39E+02	1.45E+06	4.42E+03	-13.4	1.7	-4.5	3.8
Rastering (15x15)	FR9@_1_48	Collected as a Powder	8.03E+08	2.45E+06	2.98E+05	9.09E+02	1.56E+06	4.77E+03	-12.3	1.6	0.2	3.7
Rastering (15x15)	FR9@_1_49	Collected as a Powder	7.54E+08	2.30E+06	2.78E+05	8.47E+02	1.46E+06	4.45E+03	-19.6	1.7	-6.8	3.8
Rastering (15x15)	FR9@_1_50	Collected as a Powder	7.65E+08	2.33E+06	2.84E+05	8.67E+02	1.48E+06	4.51E+03	-19.8	1.6	1.3	3.8
Rastering (15x15)	FR9@_1_51	Collected as a Powder	7.39E+08	2.26E+06	2.74E+05	8.37E+02	1.45E+06	4.43E+03	-3.8	1.7	-0.2	3.8
Rastering (15x15)	FR9@_1_52	Collected as a Powder	6.88E+08	2.10E+06	2.59E+05	7.90E+02	1.37E+06	4.17E+03	6.5	1.7	13.3	3.9
Rastering (15x15)	FR9@_1_53	Collected as a Powder	7.21E+08	2.20E+06	2.64E+05	8.05E+02	1.42E+06	4.35E+03	1.9	1.7	-14.0	3.9
Rastering (15x15)	FR9@_1_54	Collected as a Powder	7.17E+08	2.19E+06	2.69E+05	8.20E+02	1.42E+06	4.32E+03	1.8	1.7	10.6	3.9
Rastering (15x15)	FR9@_1_55	Collected as a Powder	7.08E+08	2.16E+06	2.61E+05	7.96E+02	1.40E+06	4.26E+03	0.1	1.7	-7.3	3.9
Rastering (15x15)	FR9@_1_56	Collected as a Powder	6.97E+08	2.13E+06	2.58E+05	7.89E+02	1.37E+06	4.18E+03	-4.1	1.7	-0.7	3.9
Rastering (15x15)	FR9@_1_57	Collected as a Powder	7.36E+08	2.25E+06	2.74E+05	8.36E+02	1.44E+06	4.39E+03	-8.7	1.7	3.2	3.8
Rastering (15x15)	FR9@_1_59	Collected as a Powder	7.49E+08	2.29E+06	2.79E+05	8.52E+02	1.47E+06	4.49E+03	-4.5	1.6	5.4	3.8
Rastering (15x15)	FR9@_1_60	Collected as a Powder	7.10E+08	2.17E+06	2.60E+05	7.95E+02	1.41E+06	4.29E+03	5.0	1.7	-11.4	3.9
Rastering (15x15)	FR9@_1_61	Collected as a Powder	6.46E+08	1.97E+06	2.37E+05	7.24E+02	1.28E+06	3.89E+03	1.3	1.8	-10.8	4.1
Rastering (15x15)	FR9@_1_62	Collected as a Powder	6.51E+08	1.99E+								

Table S3.3: Oxygen isotopic composition expressed as $\delta^{17}\text{O}(\text{‰})$ and $\delta^{18}\text{O}(\text{‰})$. The plasma discharge is generated in a mixture N_2O - Pentanol. The sample collected on the glass walls is noted *Powder*. The ionic counting rates are expressed in counts per second (cps) while the total number of counts (Tot) depend on the duration of the measurement. The average value of the PDCM is used as the reference value for calculating the δ values of the hotspots exhibiting marked departure in their isotopic composition. The ionic ratios $^{17}\text{O}/^{16}\text{O}$ and $^{18}\text{O}/^{16}\text{O}$ used to calculate the δ values are indicated as *Reference Std value*. They correspond to the average values determined on the PDCM noted FR9 in this Table. Error bars are the statistical precision calculated on the total number of counts ($\pm 2\sigma$). Three examples of the statistical distribution of the data is shown for (1) an olivine sample (San Carlos) (2) a terrestrial kerogen (noted D. Rumble) for which the $\delta^{17}\text{O}$ and $\delta^{18}\text{O}(\text{‰})$ were determined relative to SMOW ($\approx +0.5$ and $\approx +1.0$ ‰, respectively) (3) one region of the sample (noted PDCM) showing a statistical distribution of isotope ratios. The PDCM distribution is reported in Figure S3.3 below.

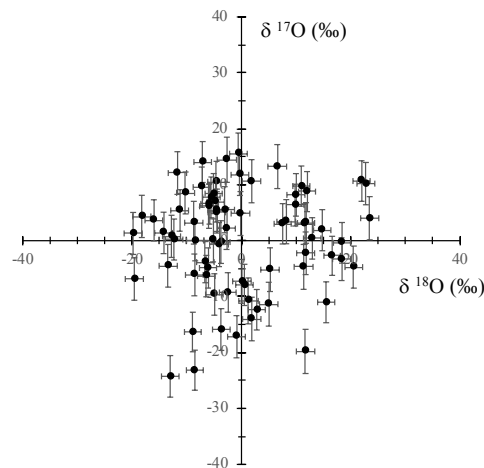


Figure S3.3: An example of the distribution of data recorded randomly on the PDCM is shown as $\delta^{17}\text{O}$ (‰) vs. $\delta^{18}\text{O}$ (‰). *Samples FR9@1* in Table S3.3. The reproducibility on the average is ± 21.7 and ± 18.4 (2σ) on $\delta^{17}\text{O}$ (‰) and $\delta^{18}\text{O}$ (‰), respectively. Error bars stand for the statistical analytical error. The statistical error bars depend only on the total number of ions detected. These error bars do not overlap 0-0‰ because the stochastic distribution of the data caused by topographic effects exceeds by far this statistical error.

S3.4:

Faraday on 16O, EM on 17O and 18O
Pentanol / N2O

17O/16O PDCM used as the reference = 3.709E-04

18O/16O PDCM used as the reference = 1.972E-03

	Sample	Comments	16O Tot.	16O Cps	17O Tot.	17O Cps	18O Tot.	18O Cps	Delta 18 ‰	± 18 (2σ) ‰	Delta 17 ‰	±17 (2σ) ‰
Depth Profiling	FR12@2_66	Area 1 in PDCM Powder	4.53E+08	1.38E+06	1.58E+05	4.81E+02	8.09E+05	2.47E+03	-94.0	2.2	-61.8	5.0
Depth Profiling	FR12@2_67	Area 1 in PDCM Powder	1.03E+09	3.13E+06	2.73E+05	8.34E+02	1.42E+06	4.35E+03	-294.9	1.7	-280.9	3.8
Depth Profiling	FR12@2_68	Area 1 in PDCM Powder	8.95E+08	2.73E+06	2.34E+05	7.13E+02	1.24E+06	3.77E+03	-299.8	1.8	-296.0	4.1
Depth Profiling	FR12@2_69	Area 1 in PDCM Powder	7.91E+08	2.41E+06	1.88E+05	5.75E+02	9.96E+05	3.04E+03	-361.6	2.0	-358.1	4.6
Depth Profiling	FR12@2_70	Area 1 in PDCM Powder	7.31E+08	2.23E+06	1.71E+05	5.22E+02	9.10E+05	2.78E+03	-368.2	2.1	-368.6	4.8
Depth Profiling	FR12@4_1	Area 2 in PDCM Powder	7.34E+08	2.24E+06	2.34E+05	7.13E+02	1.19E+06	3.64E+03	-176.2	1.8	-142.1	4.1
Depth Profiling	FR12@4_2	Area 2 in PDCM Powder	1.26E+09	3.85E+06	4.05E+05	1.24E+03	2.08E+06	6.34E+03	-165.1	1.4	-135.5	3.1
Depth Profiling	FR12@4_3	Area 2 in PDCM Powder	1.60E+09	4.90E+06	5.45E+05	1.66E+03	2.86E+06	8.71E+03	-97.2	1.2	-84.5	2.7
Depth Profiling	FR12@8_1	Area 4 in PDCM Powder	1.19E+09	3.64E+06	3.30E+05	1.01E+03	1.70E+06	5.19E+03	-276.4	1.5	-254.8	3.5
Depth Profiling	FR12@8_2	Area 4 in PDCM Powder	1.61E+09	4.91E+06	4.72E+05	1.44E+03	2.44E+06	7.44E+03	-232.2	1.3	-210.3	2.9
Depth Profiling	FR12@8_3	Area 4 in PDCM Powder	1.18E+09	3.62E+06	3.73E+05	1.14E+03	1.93E+06	5.88E+03	-174.7	1.4	-150.6	3.3
Depth Profiling	FR12@8_4	Area 4 in PDCM Powder	1.15E+09	3.52E+06	3.43E+05	1.05E+03	1.80E+06	5.50E+03	-208.0	1.5	-197.2	3.4
Depth Profiling	FR12@8_5	Area 4 in PDCM Powder	1.36E+09	4.15E+06	4.26E+05	1.30E+03	2.24E+06	6.85E+03	-162.3	1.3	-154.9	3.1

Table S3.4: Oxygen isotopic composition expressed as $\delta^{17}\text{O}(\text{‰})$ and $\delta^{18}\text{O}(\text{‰})$. The plasma discharge is generated in a mixture N₂O - Pentanol. The sample collected on the glass walls is noted *Powder*. The ionic counting rates are expressed in counts per second (cps) while the total number of counts (Tot) depend on the duration of the measurement. The ionic ratios $^{17}\text{O}/^{16}\text{O}$ and $^{18}\text{O}/^{16}\text{O}$ used to calculate the δ values are indicated as *PDCM used as the reference*. They correspond to the average values determined on the PDCM noted FR9 (cf. previous Table S3.3). Error bars are the statistical precision calculated on the total number of counts ($\pm 2\sigma$). No hotspot (i.e. no increase in the ionic intensities) is observed at the location (noted *Area*) where MIF effect was measured. This Table was used to construct the Figure 2 of the text.

S3.5:

EM on 16O, 17O and 18O
MgCl2 in Pentanol / CO2

17O/16O PDCM used as the reference = 3.717E-04
18O/16O PDCM used as the reference = 2.130E-03

FR35	Sample	Comments	16O Tot.	16O Cps	17O Tot.	17O Cps	18O Tot.	18O Cps	Delta 18 %	± 18 (2s) %	Delta 17 %	±17 (2s) %
Hot Spot 2	Robert-Juillet-2020_4.im	Hot Spot on Si Wafer	1.75E+06	6.40E+04	8.87E+02	3.25E+01	5.10E+03	1.86E+02	367.3	14.0	364.1	33.6
Hot Spot 4	Robert-Juillet-2020_4.im	Hot Spot on Si Wafer	6.34E+05	1.30E+05	2.47E+02	5.06E+01	1.37E+03	2.81E+02	14.7	27.0	48.4	63.6
Hot Spot 5	Robert-Juillet-2020_4.im	Hot Spot on Si Wafer	1.81E+06	5.85E+04	7.09E+02	2.29E+01	3.99E+03	1.29E+02	32.7	15.8	52.6	37.6
Hot Spot 7	Robert-Juillet-2020_5.im	Hot Spot on Si Wafer	1.60E+06	3.79E+04	6.51E+02	1.54E+01	3.49E+03	8.28E+01	25.7	16.9	95.2	39.2
Hot Spot 8	Robert-Juillet-2020_5.im	Hot Spot on Si Wafer	2.10E+06	6.49E+04	7.95E+02	2.46E+01	4.39E+03	1.36E+02	-18.8	15.1	18.5	35.5
Hot Spot 9	Robert-Juillet-2020_5.im	Hot Spot on Si Wafer	3.43E+06	8.49E+04	1.34E+03	3.33E+01	6.98E+03	1.73E+02	-44.2	12.0	55.4	27.3
Hot Spot 6	Robert-Juillet-2020_5.im	Hot Spot on Si Wafer	1.59E+06	1.09E+04	7.06E+02	4.86E+00	3.73E+03	2.57E+01	106.1	16.4	198.4	37.6
Hot Spot Outside 4	Robert-Juillet-2020_9.im	Hot Spot on Si Wafer	8.49E+06	8.94E+02	3.92E+03	4.13E-01	2.17E+04	2.29E+00	202.6	6.8	243.2	16.0
Hot Spot	Robert-Juillet-2020_9.im	Hot Spot on Si Wafer	1.39E+06		6.43E+02		3.46E+03		163.9	17.0	240.9	39.4
Hot Spot	Robert-Juillet-2020_9.im	Hot Spot on Si Wafer	3.34E+06	3.91E+02	1.61E+03	1.88E-01	9.12E+03	1.07E+00	280.3	10.5	294.2	24.9
Bulk Image	Robert-Juillet-2020_9.im	On Si Wafer	1.88E+07	2.07E+03	8.01E+03	8.82E-01	4.49E+04	4.94E+00	122.8	4.7	148.8	11.2

FR35	Sample	Comments	16O Tot.	16O Cps	17O Tot.	17O Cps	18O Tot.	18O Cps	Delta 18 %	± 18 (2s) %	Delta 17 %	±17 (2s) %
Random Area 1	Robert-Juillet-2020_10.im	On Si Wafer	1.38E+06	4.45E+03	9.23E+02	2.98E+00	7.02E+03	2.26E+01	1390.5	11.9	801.4	32.9
Random Area 2	Robert-Juillet-2020_10.im	On Si Wafer	5.00E+05	1.61E+03	3.26E+02	1.05E+00	2.21E+03	7.14E+00	1078.0	21.3	754.2	55.4
Random Area 3	Robert-Juillet-2020_10.im	On Si Wafer	5.70E+05	1.84E+03	3.64E+02	1.17E+00	2.82E+03	9.10E+00	1323.8	18.8	718.2	52.4
Random Area 4	Robert-Juillet-2020_10.im	On Si Wafer	2.55E+06	8.23E+03	1.55E+03	5.00E+00	1.11E+04	3.59E+01	1048.0	9.5	632.7	25.4
Random Area 5	Robert-Juillet-2020_10.im	On Si Wafer	1.71E+06	5.50E+03	1.01E+03	3.27E+00	7.32E+03	2.36E+01	1013.1	11.7	598.7	31.4
Random Area 6	Robert-Juillet-2020_10.im	On Si Wafer	1.25E+06	4.03E+03	7.93E+02	2.56E+00	5.79E+03	1.87E+01	1177.4	13.1	708.4	35.5
Random Area 7	Robert-Juillet-2020_10.im	On Si Wafer	2.42E+06	7.80E+03	1.50E+03	4.85E+00	9.58E+03	3.09E+01	860.2	10.2	671.8	25.8
Random Area 8	Robert-Juillet-2020_10.im	On Si Wafer	1.18E+06	3.82E+03	6.85E+02	2.21E+00	4.34E+03	1.40E+01	724.6	15.2	558.4	38.2
Random Area 9	Robert-Juillet-2020_10.im	On Si Wafer	9.73E+05	3.14E+03	5.47E+02	1.76E+00	3.72E+03	1.20E+01	796.6	16.4	512.6	42.8
Random Area 4	Robert-Juillet-2020_17.im	On Si Wafer	5.49E+06	1.58E+04	2.29E+03	6.59E+00	1.23E+04	3.56E+01	55.7	9.0	119.4	20.9

FR35	Sample	Comments	16O Tot.	16O Cps	17O Tot.	17O Cps	18O Tot.	18O Cps	Delta 18 %	± 18 (2s) %	Delta 17 %	±17 (2s) %
Hot Spot 1	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	1.93E+06	2.62E+05	9.80E+02	1.33E+02	4.73E+03	6.41E+02	149.0	14.5	364.2	31.9
Hot Spot 2	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	2.62E+05	4.50E+04	2.09E+02	3.59E+01	1.25E+03	2.15E+02	1240.9	28.3	1145.3	69.2
Hot Spot 3	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	1.14E+07	4.11E+05	5.27E+03	1.89E+02	2.60E+04	9.34E+02	65.8	6.2	237.5	12.8
Hot Spot 5	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	4.47E+06	3.45E+05	1.98E+03	1.53E+02	1.03E+04	7.93E+02	79.4	9.9	193.0	23.5
Hot Spot 7	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	7.20E+06	3.57E+05	3.12E+03	1.55E+02	1.63E+04	8.07E+02	61.4	7.8	167.0	17.9
Hot Spot 8	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	8.46E+06	3.45E+05	3.59E+03	1.46E+02	1.83E+04	7.48E+02	18.3	7.4	142.4	16.7
Hot Spot 9	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	6.01E+05	4.66E+04	3.60E+02	2.79E+01	2.14E+03	1.66E+02	669.3	21.6	613.0	52.7
Hot Spot 10	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	1.28E+06	1.08E+05	6.18E+02	5.20E+01	3.36E+03	2.83E+02	231.3	17.2	296.7	40.2
Hot Spot 11	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	1.57E+06	1.82E+05	7.54E+02	8.73E+01	3.79E+03	4.38E+02	132.3	16.2	291.6	36.4
Hot Spot 12	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	1.13E+06	2.58E+05	5.17E+02	1.18E+02	2.77E+03	6.33E+02	151.8	19.0	231.5	44.0
Hot Spot 13	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	2.29E+05	9.54E+04	1.61E+02	6.71E+01	8.88E+02	3.70E+02	820.4	33.6	891.3	78.8
Random Area 5	Robert-Juillet-2020_17.im	On Si Wafer	5.36E+06	1.55E+04	2.31E+03	6.67E+00	1.22E+04	3.52E+01	69.3	9.0	160.3	20.8
Random Area 6	Robert-Juillet-2020_17.im	On Si Wafer	2.34E+07	6.74E+04	1.01E+04	2.91E+01	5.32E+04	1.54E+02	69.9	4.3	159.5	10.0
Random Area 7	Robert-Juillet-2020_17.im	On Si Wafer	6.42E+06	1.85E+04	3.04E+03	8.78E+00	1.64E+04	4.74E+01	200.0	7.8	274.0	18.1
Random Area 8	Robert-Juillet-2020_17.im	On Si Wafer	3.32E+06	9.57E+03	1.50E+03	4.31E+00	8.38E+03	2.42E+01	186.2	10.9	212.8	25.9
Random Area 9	Robert-Juillet-2020_17.im	On Si Wafer	2.84E+07	8.20E+04	1.26E+04	3.64E+01	6.55E+04	1.89E+02	82.8	3.9	196.0	8.9
Hot Spot 6	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	8.52E+06	2.55E+05	3.50E+03	1.05E+02	1.83E+04	5.48E+02	9.5	7.4	106.4	16.9
Random Area 2	Robert-Juillet-2020_17.im	On Si Wafer	1.05E+07	3.03E+04	4.28E+03	1.23E+01	2.26E+04	6.52E+01	9.6	6.7	94.3	15.3
Hot Spot 4	Robert-Juillet-2020_17.im	Hot Spot on Si Wafer	7.32E+06	3.09E+05	3.09E+03	1.31E+02	1.59E+04	6.72E+02	19.3	7.9	135.8	18.0
Random Area 3	Robert-Juillet-2020_17.im	On Si Wafer	3.69E+07	1.06E+05	1.52E+04	4.39E+01	7.90E+04	2.28E+02	6.1	3.6	110.6	8.1

FR35	Sample	Comments	16O Tot.	16O Cps	17O Tot.	17O Cps	18O Tot.	18O Cps	Delta 18 %	± 18 (2s) %	Delta 17 %	±17 (2s) %
Hot Spot 3	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	8.71E+06	1.54E+05	3.74E+03	6.59E+01	1.91E+04	3.36E+02	27.4	7.2	155.6	16.3
Hot Spot 8	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	5.75E+06	1.25E+05	2.59E+03	5.64E+01	1.33E+04	2.89E+02	85.1	8.7	214.2	19.6
Hot Spot 10	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	1.42E+06	1.68E+05	6.33E+02	7.47E+01	3.14E+03	3.70E+02	34.1	17.9	195.8	39.7
Hot Spot 11	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	4.26E+06	1.24E+05	1.85E+03	5.38E+01	9.45E+03	2.75E+02	41.7	10.3	168.1	23.3
Hot Spot 13	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	3.35E+06	9.33E+04	1.54E+03	4.30E+01	7.75E+03	2.16E+02	85.4	11.4	239.0	25.4
Hot Spot 15	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	2.18E+06	8.59E+04	9.78E+02	3.85E+01	4.92E+03	1.94E+02	58.2	14.3	205.9	32.0
Hot Spot 16	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	4.29E+06	1.39E+05	1.91E+03	6.18E+01	9.54E+03	3.09E+02	43.9	10.2	195.3	22.9
Hot Spot 17	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	2.79E+06	5.33E+04	1.22E+03	2.32E+01	6.25E+03	1.19E+02	51.1	12.7	174.2	28.7
Hot Spot 21	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	9.50E+06	3.18E+04	4.20E+03	1.41E+01	2.20E+04	7.38E+01	89.4	6.7	190.1	15.4
Hot Spot 22	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	5.28E+06	1.10E+04	2.33E+03	1.37E+01	1.20E+04	7.03E+01	64.1	9.1	186.7	20.7
Random Area 10	Robert-Juillet-2020_18.im	On Si Wafer	1.94E+07	6.18E+04	8.59E+03	2.73E+01	4.47E+04	1.42E+02	79.2	4.7	188.9	10.8
Random Area 11	Robert-Juillet-2020_18.im	On Si Wafer	9.60E+06	3.05E+04	4.13E+03	1.31E+01	2.16E+04	6.87E+01	57.3	6.8	157.9	15.6
Random Area 12	Robert-Juillet-2020_18.im	On Si Wafer	1.65E+07	5.24E+04	7.43E+03	2.36E+01	3.83E+04	1.22E+02	91.2	5.1	213.5	11.6
Random Area 13	Robert-Juillet-2020_18.im	On Si Wafer	1.43E+07	4.54E+04	6.00E+03	1.91E+01	3.16E+04	1.00E+02	37.2	5.6	129.5	12.9
Random Area 14	Robert-Juillet-2020_18.im	On Si Wafer	1.61E+07	5.13E+04	6.79E+03	2.16E+01	3.61E+04	1.15E+02	51.2	5.3	131.6	12.1
Random Area 15	Robert-Juillet-2020_18.im	On Si Wafer	2.38E+07	7.58E+04	1.03E+04	3.26E+01	5.24E+04	1.66E+02	31.7	4.4	157.4	9.9
Random Area 16	Robert-Juillet-2020_18.im	On Si Wafer	6.13E+06	1.95E+04	2.73E+03	8.67E+00	1.47E+04	4.68E+01	127.9	8.2	197.6	19.1
Random Area 17	Robert-Juillet-2020_18.im	On Si Wafer	1.07E+07	3.40E+04	4.80E+03	1.53E+01	2.52E+04	8.01E+01	105.3	6.3	205.5	14.4
Random Area 18	Robert-Juillet-2020_18.im	On Si Wafer	1.21E+07	3.85E+04	5.34E+03	1.70E+01	2.80E+04	8.90E+01	84.6	6.0	184.2	13.7
Random Area 19	Robert-Juillet-2020_18.im	On Si Wafer	1.03E+07	3.28E+04	4.58E+03	1.45E+01	2.34E+04	7.45E+01	67.7	6.5	194.3	14.6
Random Area 20	Robert-Juillet-2020_18.im	On Si Wafer	7.59E+06	2.41E+04	3.63E+03	1.15E+01	1.79E+04	5.68E+01	105.8	7.5	286.6	16.8
Random Area 21	Robert-Juillet-2020_18.im	On Si Wafer	4.20E+06	1.34E+04	1.90E+03	6.03E+00	1.03E+04	3.27E+01	149.1	9.9	216.1	22.9
Random Area 22	Robert-Juillet-2020_18.im	On Si Wafer	7.27E+06	2.31E+04	3.41E+03	1.08E+01	1.79E+04	5.69E+01	156.4	7.5	263.5	17.1
Random Area 23	Robert-Juillet-2020_18.im	On Si Wafer	7.86E+06	2.50E+04	3.50E+03	1.11E+01	1.88E+04	5.98E+01	124.5	7.3	199.9	16.9
Random Area 24	Robert-Juillet-2020_18.im	On Si Wafer	9.66E+06	3.07E+04	4.43E+03	1.41E+01	2.31E+04	7.33E+01	121.1	6.6	232.5	15.0
Random Area 25	Robert-Juillet-2020_18.im	On Si Wafer	7.39E+06	2.35E+04	3.44E+03	1.09E+01	1.69E+04	5.36E+01	70.9	7.7	254.0	17.0
Hot Spot 4	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	5.52E+06	1.38E+05	2.22E+03	5.54E+01	1.15E+04	2.88E+02	-19.8	9.3	81.1	21.2
Hot Spot 9	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	3.51E+06	1.15E+05	1.45E+03	4.76E+01	7.47E+03	2.46E+02	0.5	11.6	110.4	26.3
Hot Spot 12	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	2.08E+06	9.00E+04	8.44E+02	3.65E+01	4.52E+03	1.95E+02	19.4	14.9	91.2	34.4
Hot Spot 14	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	3.37E+06	1.08E+05	1.33E+03	4.27E+01	7.14E+03	2.29E+02	-5.9	11.8	60.2	27.4
Hot Spot 19	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	3.09E+06	1.11E+05	1.23E+03	4.40E+01	6.55E+03	2.34E+02	-5.6	12.4	69.8	28.5
Hot Spot 5	Robert-Juillet-2020_18.im	Hot Spot on Si Wafer	3.35E+06	1.37E+05	1.33E+03	5.45E+01	7.34E+03	3.00E+02	27.7	11.7	67.	

Table S3.5: Oxygen isotopic composition expressed as $\delta^{17}\text{O}(\text{‰})$ and $\delta^{18}\text{O}(\text{‰})$. The plasma discharge is generated in a mixture MgCl_2 - Pentanol - CO_2 . The sample is collected on a silicon wafer. The ionic counting rates are expressed in counts per second (cps) while the total number of counts (Tot.) depends on the duration of the measurement. The ionic ratios $^{17}\text{O}/^{16}\text{O}$ and $^{18}\text{O}/^{16}\text{O}$ of the PDCM used to calculate the δ values are indicated as *Reference Std values*. The area of the PDCM is defined on the same wafer where isotopic hotspots were measured. In absence of identifiable hotspots in an ion image, the raster area ($20 \times 20 \mu\text{m}$) was divided in 9 *Random Areas*. *Bulk* is the measured (not calculated) average of these 9 areas. Note that isotopic effects are detectable even in hotspot free areas. This Table was used to construct the Figure 2 of the text. All sample names correspond to files registered in the NanoSims data files.

S4. Calculated MIF slopes in a 3 isotope diagram

Several theories have been developed to account for the MIF effect^(6,16, 25-30). They were all based on the results obtained on the isotopic composition of ozone for which numerous and precise laboratory parameters are available in the literature. They are centered on the interpretation of the empirical fit (the η factor) discovered experimentally by Janssen⁽²⁾, which accounts for the isotopic compositions of the different isotopomers of ozone:

$$\alpha = \alpha_{\text{MDF}} \cdot \eta$$

The η factor referred here to as the MIF factor (Mass Independent Fractionation factor) illustrates the fact that, for any isotopomer masses, there is a constant shift between the observed fractionation and the expected MDF factors α_{MDF} 's. The origin of η will be discussed elsewhere. Whatever its origin, η characterizes the reaction rate ratio of complexes having different symmetry such as $[^{16}\text{O}\dots^{17}\text{O}]^*/[^{16}\text{O}\dots^{16}\text{O}]^*$. Neglecting the MDF (i.e. $\alpha_{\text{MDF}}=1$), and extending this ratio to all possible isotopic reactions for a given chemical element gives the overall mass independent isotopic fractionation factor $m^{-16}\alpha$ ⁽¹⁸⁾:

$$m^{-16}\alpha = \frac{1+x(m)[\eta^{-1}-1]}{1+x(16)[\eta^{-1}-1]}$$

m designates 17 or 18 and $x(16)$, $x(17)$ and $x(18)$ stand for the relative abundances of ^{16}O , ^{17}O and ^{18}O (i.e. 99.76×10^{-2} , 380×10^{-6} and 2000×10^{-6} respectively). Note that, in this formalism, the isotopic fractionation factor is not anymore the rate constant ratio, but instead the reaction rate ratio. In the 3 isotope diagram, the slope $\Delta(\delta^{17}\text{O})/\Delta(\delta^{18}\text{O})$ depends only on the relative abundances of the 3 isotopes with no free parameters.

Taking the example of magnesium isotopes ($X(24)=0.7899$, $x(25)=0.100$, $x(26)=0.11$), the slope $\Delta(\delta^{25}\text{Mg})/\Delta(\delta^{26}\text{O})=0.985$, $\forall(\Delta\eta)$. A detailed theoretical treatment of the MIF effect is proposed in *Chem. Phys.* (in press).

S5. Details of the Pentane/N₂O analyses.

	delta 18		delta 17			delta 18		delta 17	
	‰	±2s	‰	±2s		‰	±2s	‰	±2s
FR12-8_1	-306	19.7	-294	27.1	FR12-8_3	-151	19.4	-136	25.9
FR12-8_1	-306	19.7	-287	27.0	FR12-8_3	-163	19.5	-127	26.0
FR12-8_1	-298	19.7	-269	26.9	FR12-8_3	-159	19.5	-136	26.0
FR12-8_1	-288	19.7	-274	26.9	FR12-8_3	-160	19.5	-154	26.1
FR12-8_1	-287	19.7	-266	26.8	FR12-8_3	-174	19.5	-150	26.1
FR12-8_1	-286	19.7	-259	26.7	FR12-8_3	-172	19.5	-144	26.1
FR12-8_1	-274	19.6	-252	26.7	FR12-8_3	-173	19.5	-143	26.1
FR12-8_1	-277	19.6	-245	26.6	FR12-8_3	-179	19.5	-156	26.2
FR12-8_1	-268	19.6	-244	26.6	FR12-8_3	-178	19.5	-145	26.1
FR12-8_1	-268	19.6	-243	26.6	FR12-8_3	-177	19.5	-142	26.1
FR12-8_1	-261	19.6	-242	26.6	FR12-8_3	-179	19.5	-162	26.2
FR12-8_1	-267	19.6	-251	26.6	FR12-8_3	-181	19.5	-156	26.2
FR12-8_1	-267	19.6	-249	26.6	FR12-8_3	-181	19.5	-161	26.2
FR12-8_1	-267	19.6	-255	26.7	FR12-8_3	-186	19.5	-157	26.2
FR12-8_1	-272	19.6	-239	26.6	FR12-8_3	-184	19.5	-168	26.3
FR12-8_1	-268	19.6	-243	26.6	FR12-8_3	-188	19.5	-159	26.3
FR12-8_1	-267	19.6	-247	26.6	FR12-8_3	-184	19.5	-158	26.3
FR12-8_1	-264	19.6	-246	26.7	FR12-8_3	-190	19.5	-168	26.3
FR12-8_1	-267	19.6	-247	26.7	FR12-8_3	-189	19.5	-159	26.3
FR12-8_1	-273	19.6	-247	26.7	FR12-8_3	-186	19.5	-172	26.4
FR12-8_2	-143	19.2	-126	25.2	FR12-8_4	-191	19.5	-170	26.1
FR12-8_2	-143	19.2	-120	25.2	FR12-8_4	-195	19.5	-186	26.3
FR12-8_2	-149	19.2	-117	25.2	FR12-8_4	-201	19.5	-192	26.3
FR12-8_2	-148	19.2	-124	25.2	FR12-8_4	-203	19.5	-175	26.3
FR12-8_2	-151	19.3	-126	25.2	FR12-8_4	-209	19.5	-194	26.4
FR12-8_2	-154	19.3	-121	25.2	FR12-8_4	-207	19.5	-191	26.4
FR12-8_2	-154	19.3	-134	25.3	FR12-8_4	-213	19.6	-204	26.5
FR12-8_2	-158	19.3	-132	25.3	FR12-8_4	-213	19.6	-207	26.5
FR12-8_2	-159	19.3	-139	25.3	FR12-8_4	-211	19.6	-194	26.4
FR12-8_2	-159	19.3	-132	25.3	FR12-8_4	-222	19.6	-201	26.5
FR12-8_2	-164	19.3	-148	25.4	FR12-8_4	-218	19.6	-211	26.6
FR12-8_2	-160	19.3	-146	25.4	FR12-8_4	-221	19.6	-218	26.6
FR12-8_2	-160	19.3	-132	25.3	FR12-8_4	-219	19.6	-222	26.7
FR12-8_2	-162	19.3	-134	25.3	FR12-8_4	-226	19.6	-227	26.7
FR12-8_2	-168	19.3	-159	25.4	FR12-8_4	-228	19.6	-224	26.7
FR12-8_2	-167	19.3	-134	25.4	FR12-8_4	-227	19.6	-229	26.7
FR12-8_2	-168	19.3	-143	25.4	FR12-8_4	-231	19.6	-222	26.7
FR12-8_2	-169	19.3	-137	25.4	FR12-8_4	-232	19.6	-220	26.7
FR12-8_2	-170	19.3	-148	25.4	FR12-8_4	-236	19.6	-212	26.7
FR12-8_2	-169	19.3	-141	25.4	FR12-8_4	-240	19.6	-230	26.8

Table S5: Pentanol/N₂O experiment. $\delta^{17,18}\text{O}$ measured by depth profiling at the location of the isotopic MIF hotspots. No hotspot in ion intensities was observed at the location of the isotopic MIF hotspots. Each step represents a sputtered thickness between 200 and 300 nm. Data are reported in the Figure S5 below.

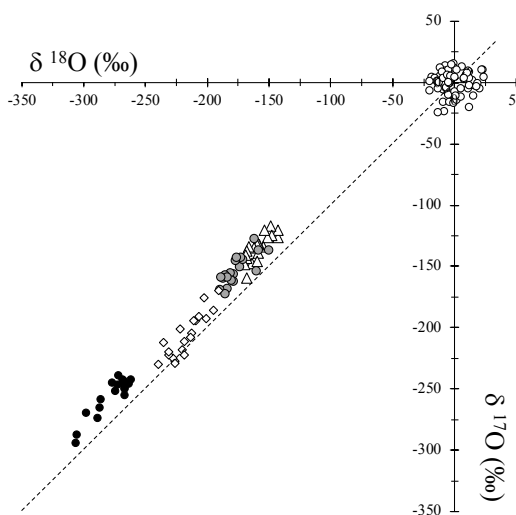


Figure S5: The $\delta^{17,18}\text{O}$ values reported for the samples FR12@8 from 1 to 4 in Table S3.4 are average values of ≈ 80 measurements for each sample. This figure shows the evolution of the $\delta^{17,18}\text{O}$ values averaged every 4 measurements. The PDCM used as a reference value for δ units is shown for comparison (cf. Table S3.3; open dots around 0-0‰). Black dots, open diamonds, grey dots and open triangle stand for FR12@ 8 from 1 to 4, respectively. Error bars (statistic and PDCM reproducibility) are not shown for clarity but never exceed ± 19 and ± 26 ‰ (2σ) for $\delta^{18}\text{O}$ and $\delta^{17}\text{O}$, respectively. The 1:1 line is drawn for reference. Note that the correlation defined by the data do not intercept the reference at 0-0‰. This may be caused by an analytical “matrix effect” but may also reflects a combined MDF and MIF effect.

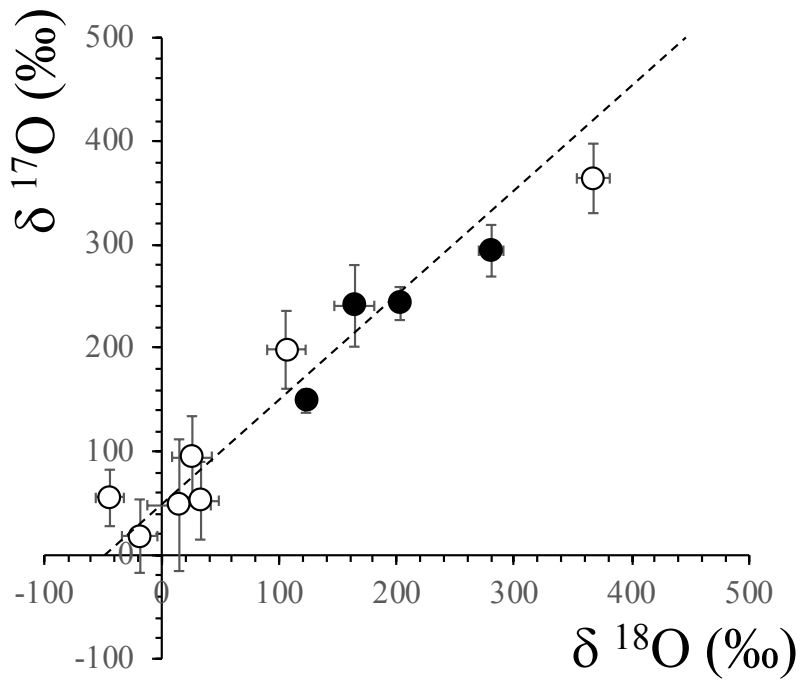


Figure S6: The $\delta^{17,18}\text{O}$ of individual emissivity hotspot measured in 3 different images: *Robert-Juillet-2020_4, _5 and _9* reported in Table S3.5. Black dots stand for the same *image 9*. The purpose of this plot is to show that the internal variations between hotspots being recorded in the same image, cannot result from bias in the instrumental setups. The dashed line is the reference 1:1 correlation line. The corresponding emissivity hotspots from *image 9* is shown in the Fig. 2b of the text.